

[www.newtonx.org](http://www.newtonx.org)

# NEWTON-X

a package for Newtonian  
dynamics close to the  
crossing seam

Documentation based on NEWTON-X version 1.4.0 (release August 15, 2013)



# 1 Table of contents

---

1	TABLE OF CONTENTS .....	III
2	ABOUT NEWTON-X .....	1
2.1	<i>General Information</i> .....	1
2.2	<i>Contact information</i> .....	1
2.3	<i>What is new in this version</i> .....	1
3	MIXED QUANTUM-CLASSICAL DYNAMICS SIMULATIONS .....	2
4	DEVELOPERS, COLLABORATORS AND CONTRIBUTORS .....	4
5	HOW TO REFERENCE NEWTON-X .....	5
6	QUICK START .....	6
6.1	<i>NEWTON-X Tutorial</i> .....	6
6.2	<i>The NXINP tool</i> .....	6
6.3	<i>Initial conditions generation</i> .....	7
6.4	<i>Dynamics Input</i> .....	7
6.4.1	Adiabatic dynamics (dynamics on one surface) .....	7
6.4.2	Non-adiabatic dynamics (Surface hopping) .....	8
6.4.3	Mixed adiabatic and non-adiabatic dynamics .....	8
6.5	<i>Creating the trajectory inputs</i> .....	8
6.6	<i>Running the dynamics</i> .....	9
6.7	<i>Where are the results?</i> .....	9
6.8	<i>Overview of the file structure</i> .....	9
6.8.1	Dynamics simulations .....	9
7	CAPABILITIES .....	11
7.1	<i>General features</i> .....	11
7.2	<i>Overview of the available interfaces</i> .....	12
8	HOW TO GET NEWTON-X .....	13
9	HOW TO INSTALL NEWTON-X .....	14
9.1	<i>Binary distribution</i> .....	14
9.2	<i>To install NEWTON-X you need</i> .....	14
9.3	<i>To run NEWTON-X you need</i> .....	14
9.4	<i>Installation procedure</i> .....	14
9.5	<i>Setup of third-party programs</i> .....	16
9.6	<i>Verification of installation</i> .....	17
9.7	<i>Compatibility between DALTON and NEWTON-X installations</i> .....	17
10	INITIAL CONDITIONS GENERATION .....	19
10.1	<i>How to execute INITICOND</i> .....	19
10.2	<i>Input parameters</i> .....	19
10.3	<i>What you need to execute</i> .....	23
10.3.1	Additional files .....	23
10.3.2	The JOB_AD directory .....	24
10.3.3	Save files option .....	25
10.4	<i>Output</i> .....	25
10.5	<i>Running in several computers</i> .....	26
11	TRAJECTORY-INPUT GENERATION, INITIAL CONDITIONS FOR MULTIPLE STATES, AND SPECTRA .....	27
11.1	<i>NXINP and options in the MAKEDIR.PL program</i> .....	27
11.2	<i>Transition spectra</i> .....	28
11.3	<i>Initial conditions for multiple states</i> .....	30
11.4	<i>Managing several trajectories</i> .....	30
11.5	<i>About energy restrictions</i> .....	31
12	DYNAMICS INPUTS .....	32
12.1	<i>What is necessary to run</i> .....	32
12.1.1	control.dyn .....	32
12.1.2	Geometry .....	34
12.1.3	Velocity .....	34

12.1.4	Freezing atoms.....	35
12.1.5	Specific input for quantum-chemistry electronic-structure calculations.....	35
12.1.5.1	Which electronic method to use.....	35
12.1.5.2	Using analytical models.....	35
12.1.5.3	COLUMBUS.....	36
12.1.5.4	TURBOMOLE.....	38
12.1.5.5	DFTB.....	39
12.1.5.6	GAUSSIAN.....	39
12.1.5.7	TINKER.....	44
12.1.5.8	DFTB+.....	44
12.1.5.9	GAMESS.....	44
12.1.5.10	Hybrid Gradients (QM/MM).....	45
12.1.6	Thermostat control.....	48
12.1.7	Non-adiabatic dynamics control.....	49
12.1.8	Time-derivative couplings.....	53
12.1.9	Wave function coefficients.....	53
12.1.10	Propagation of molecular orbitals.....	54
12.1.11	Boundaries.....	54
12.1.12	Stop conditions.....	54
12.2	<i>How to execute NEWTON-X</i> .....	55
12.3	<i>Output files</i> .....	55
12.4	<i>Restarting the job</i> .....	56
12.5	<i>Customized analysis</i> .....	57
13	STATISTICAL ANALYSIS.....	58
13.1	<i>What is needed to run</i> .....	58
13.2	<i>How to execute ANALYSIS</i> .....	59
13.3	<i>Output files</i> .....	59
14	NORMAL MODE AND ESSENTIAL DYNAMICS ANALYSIS.....	62
14.1	<i>Normal mode analysis</i> .....	62
14.1.1	Input parameters.....	62
14.1.2	Text Output.....	63
14.1.3	Graphical Output.....	63
14.2	<i>Trajectory alignment</i> .....	64
14.2.1	Input parameters.....	64
14.3	<i>Average Structure</i> .....	64
14.3.1	Input parameters.....	64
14.4	<i>Essential Dynamics</i> .....	65
14.4.1	Input parameters.....	65
14.5	<i>Python subroutine libraries</i> .....	65
14.6	<i>Required packages to run NMA analysis</i> .....	66
15	TOOLS.....	67
15.1	<i>Plotting energy x time</i> .....	67
15.2	<i>Plotting velocities and molecular orbitals</i> .....	67
15.3	<i>Smoothangle</i> .....	68
15.4	<i>Collectjumps</i> .....	68
15.5	<i>Diagnostic</i> .....	69
15.6	<i>Conversion tools</i> .....	69
15.7	<i>Split and merge initial conditions</i> .....	70
16	TECHNICAL DETAILS.....	71
16.1	<i>Templates and interfaces with new programs</i> .....	71
16.2	<i>Conversion factors</i> .....	72
16.3	<i>Format of internal files</i> .....	72
16.4	<i>Normal modes</i> .....	73
16.5	<i>Output files of the SH program</i> .....	74
16.6	<i>CIOVERLAP documentation</i> .....	74
16.6.1	CIOVERLAP program.....	75
16.6.2	CIS_CASIDA.....	76
16.6.3	CIS_SLATERGEN.....	77
16.6.4	CIVECCOMPARE.....	77
16.6.5	CIVECCONSOLIDATE.....	78
16.6.6	READSIFS.....	79
16.6.7	CIPC.X, MCPC.X.....	79
16.7	<i>Quick description of the programs</i> .....	79
16.7.1	Initial conditions.....	79
16.7.2	Dynamics.....	80

16.7.3	Tools.....	81
16.7.4	Statistical analysis.....	81
17	LINKS TO THIRD-PARTY PROGRAMS.....	83
18	REFERENCES.....	84

## 2 About NEWTON-X

---

### 2.1 General Information

NEWTON-X<sup>1,2</sup> is a general purpose program package for molecular dynamics in the electronic excited states, including mixed quantum-classical methods (surface hopping).

NEWTON-X modular development allows it to be easily linked to any quantum chemistry package that can provide energy gradients and (optionally) non-adiabatic coupling vectors.

In the current version, NEWTON-X can perform dynamics using COLUMBUS,<sup>3,4</sup> TURBOMOLE,<sup>5</sup> DFTB,<sup>6</sup> DFTB+, GAUSSIAN 03,<sup>7</sup> GAUSSIAN 09,<sup>8</sup> and GAMESS<sup>9</sup> program packages. Nonadiabatic dynamics using hybrid gradients (including QM/MM approach) is available as well for combinations between TINKER<sup>10</sup> and one of the following programs COLUMBUS, TURBOMOLE, DFTB+, and analytical models. Other third-party programs providing energies and nonadiabatic couplings can be easily integrated to work with NEWTON-X as well.

NEWTON-X code is distributed free of charge for non-commercial and non-profit uses. You may use the program freely and adapt the code to your needs. Please, if you have any enhancements, share them with us. You are, however, not allowed to re-distribute code or binaries in parts or total. Anyone intending to use NEWTON-X must contact us.

*NEWTON-X is shipped to the user "as is". There is no guarantee that it will work as described. Usage is at your own risk, there is no liability taken for any damage or loss of data and/or money. If you experience problems please tell the developers supplying the version-number – they are always grateful for any hint, but bear in mind that there is no kind of official support for NEWTON-X.*

### 2.2 Contact information

Mario Barbatti  
Max-Planck-Institut für Kohlenforschung  
Kaiser-Wilhelm-Platz 1  
D-45470 Mülheim an der Ruhr, Germany  
[barbatti@kofo.mpg.de](mailto:barbatti@kofo.mpg.de)  
[www.mpi-muelheim.mpg.de/private/barbatti](http://www.mpi-muelheim.mpg.de/private/barbatti)

The program can be downloaded from:  
[www.newtonx.org](http://www.newtonx.org)

Support can be obtained through the discussion forum at:  
<https://groups.google.com/forum/?fromgroups#!forum/newtonx>

### 2.3 What is new in this version

NEWTON-X version 1.4.0 is a major update over version 1.3.1. The main modification is the inclusion of nonadiabatic dynamics at TDDFT level with GAUSSIAN 09 and at CC2 and ADC(2) levels with TURBOMOLE.

## 3 Mixed quantum-classical dynamics simulations

---

Mixed quantum-classical approaches<sup>11</sup> are probably the most employed class of methods to perform excited-state molecular dynamics simulations including non-adiabatic effects. In these approaches, which include the surface hopping and the Ehrenfest methods, the nuclear time evolution is treated classically by means of the Newton's equations, while the time evolution of the population of each electronic state is treated separately. In the surface hopping method,<sup>12</sup> the time evolution of the population is obtained in two steps: first, non-adiabatic transition probability between each pair of states is computed and a stochastic algorithm is applied to decide in which state the classical trajectory is propagated in the next time step. Second, statistics over a large set of independently computed trajectories allows getting the fraction of trajectories (occupation) in each state as a function of time. The main hypothesis behind the surface hopping approach is that the *occupation* and the quantum *population* of each electronic state are the same if an infinite number of trajectories are computed.<sup>13</sup> A recent review of the method is done in Ref.<sup>14</sup>

There are several proposed ways to evaluate the non-adiabatic transition probabilities, since the most simple methods which just assume that the probability is the unity if the energy gap between the states is smaller than some threshold, to more sophisticated approaches which take into account the variation of wavefunction coefficients<sup>15</sup> or compute the Landau-Zener transition probability.<sup>16</sup> The most reliable procedure to compute the non-adiabatic transition probability for surface hopping simulations is the Tully's fewest switches algorithm.<sup>13</sup> In this approach, the time-dependent Schroedinger equation (TDSE) is integrated simultaneously to the classical trajectory.<sup>17</sup> To cope with the lack of non-local information introduced by the on-the-fly approach, non-local terms in the TDSE are neglected and the nuclear wavefunction is supposed to be entirely localized at the classical position determined by the Newton's equation. The integration of this semi-classical version of the TDSE gives the adiabatic populations of the electronic states, which are then used to compute the probability using the fewest-switches formula.

The integration of the TDSE depends on nonadiabatic coupling terms connecting different states. If adiabatic representation is used to expand the molecular wavefunction, non-adiabatic coupling vectors should be computed. Alternatively, if diabatic representation is used, non-diagonal Hamiltonian matrix elements should be computed. Either way, the computation of the non-adiabatic coupling terms are the bottleneck for non-adiabatic dynamics approaches. These terms are not usually available for most of quantum chemical methods and when they are, their computational cost increases with the square of the number of electronic states.<sup>18</sup> These difficulties have motivated, on one hand, the search for approximated hopping algorithms as those mentioned above, and on the other hand the computation of the coupling terms based on wavefunction overlaps.<sup>15,18-22</sup> Besides that, it has been an important achievement the development of procedures for analytical computation of them at the multireference configuration interaction (MRCI) and at the multiconfigurational self-consistent field (MCSCF) methods.<sup>23,24</sup>

One consequence of the hyperlocalization of the nuclear wavefunction in mixed quantum-classical approaches is that non-diagonal terms in the density matrix do not vanish with time as they should do.<sup>25,26</sup> In surface-hopping, this lack of decoherence results in an excessive number of hopping events from lower to upper states, which disturbs the accomplishment of the occupation/population hypothesis mentioned above. Decoherence can be imposed by applying an *ad hoc* correction to the

adiabatic population every time step, which forces the non-diagonal terms in the density matrix to damp to zero within a certain time constant.<sup>26</sup>

When a hopping between two states takes place, it usually does through finite energy gap. In order to keep the total energy constant in the subsequent trajectory, it is necessary to correct the kinetic energy, for example, by rescaling the momentum or by adding more momentum at the direction of the non-adiabatic coupling vector.<sup>17,27</sup> It may also happen that the stochastic algorithm attempts to make a hopping from a lower to an upper state in a region where there is not enough energy to do so. Such cases have been usually treated by forbidding the hopping occurrence.<sup>17</sup> The momentum can be kept or reversed afterwards. Another possibility is to take the time uncertainty principle to search for a geometry nearby where the hopping is allowed.<sup>28</sup>

Because of the stochastic nature of the fewest-switches surface-hopping approach, trajectories starting with the same initial conditions will give rise to different time development. Moreover, the initial conditions should reflect the initial phase space distribution. Therefore the averages that define the state occupation should in principle be performed over this double ensemble of trajectories starting in different points of the phase space, several times in each one. Because of computational limitations, this procedure is usually reduced to a single ensemble of trajectories starting in different points of the phase space only once in each one.

The initial condition ensemble can be generated in a diversity of ways. For instance, the simulation of an instantaneously excited wave packet into the Franck-Condon region may be done by selecting geometries and velocities from a dynamics in the ground state, regarding this dynamics run for long enough period as to allow an adequate sampling of the phase space. Alternatively, each nuclear degree of freedom can be treated within the harmonic approximation and a Wigner distribution can be build.

Most of the methods and algorithms mentioned in this short introduction are implemented in the NEWTON-X program.



## 4 Developers, Collaborators and Contributors

---

The NEWTON-X program has been developed in a multi-institutional collaboration, involving researchers from several countries.

People working in the general NEWTON-X development are:

Mario Barbatti	Max-Planck-Institut für Kohlenforschung	Germany
Hans Lischka	University of Vienna/Texas Tech	Austria/USA
Matthias Ruckebauer	Frankfurt University	Germany
Felix Plasser	University of Vienna	Austria

Many other people have contributed in the past or are still contributing to development of specific algorithms in NEWTON-X. They are:

*Surface hopping routines and initial condition distributions*

Giovanni Granucci	University of Pisa	Italy
Maurizio Persico	University of Pisa	Italy

*Time-dependent non-adiabatic couplings*

Jiri Pittner	J. Heyrovsky Institute	Czech Republic
--------------	------------------------	----------------

*NEWTON-X/GAMESS interface*

Aaron West	Iowa State University	USA
Theresa Windus	Iowa State University	USA

*NEWTON-X/GAUSSIAN 09 interface*

Rachel Crespo-Otero	Max-Planck-Institut für Kohlenforschung	Germany
---------------------	---	---------

Contributions from the following colleagues are also acknowledged:

Mirjana Eckert-Maksić, Sergio A. Losilla Fernández, Vladimir Lukeš, Peter Szalay, Bernhard Sellner, Atila Tajti, Mario Vazdar, Markus Weihs, Gunther Zechmann, Tomas Zeleny.

## 5 How to reference NEWTON-X

---

Please, cite NEWTON-X as:

M. Barbatti, G. Granucci, M. Persico, M. Ruckebauer, M. Vazdar, M. Eckert-Maksić and H. Lischka, *J. Photochem. Photobiol. A* **190**, 228 (2007).

M. Barbatti, G. Granucci, M. Ruckebauer, F. Plasser, J. Pittner, M. Persico, H. Lischka, *NEWTON-X: a package for Newtonian dynamics close to the crossing seam*, version 1.4, [www.newtonx.org](http://www.newtonx.org) (2013).

References to specific methods, algorithms and third-party programs used in NEWTON-X are given along this documentation. For a summary, see Chapter 7.

## 6 Quick start

---

### 6.1 NEWTON-X Tutorial

A NEWTON-X tutorial with step-by-step procedures for several examples is available at [www.univie.ac.at/newtonx/docs/tutorial.pdf](http://www.univie.ac.at/newtonx/docs/tutorial.pdf)

### 6.2 The NXINP tool

Most of NEWTON-X inputs are prepared with nxinp program. To execute nxinp, type:

```
$NX/nxinp
```

The first screen should look like:

```
=====
                          NEWTON-X
Newtonian dynamics close to the crossing seam
                          www.newtonx.org
=====

                          MAIN MENU

1. GENERATE INITIAL CONDITIONS

2. SET BASIC INPUT

3. SET GENERAL OPTIONS

4. SET NON-ADIABATIC DYNAMICS

5. GENERATE TRAJECTORIES AND SPECTRUM

6. SET STATISTICAL ANALYSIS

7. EXIT
```

Select one option (1-7):\_

The next sections will guide you through each one of these options. By now, it is enough to note that nxinp is self-explaining. When you select one of the options, say, option 2 (Set basic input), you are asked about a series of parameters. A sequence of input may be, for example:

```
=====
                          NEWTON-X
Newtonian dynamics close to the crossing seam
                          www.newtonx.org
=====

                          SET BASIC OPTIONS
```

```
nat: Number of atoms.
```

NEWTON-X: Newtonian dynamics close to the crossing seam

```
There is no value attributed to nat
Enter the value of nat : 12
Setting nat = 12

nstat: Number of states.
The current value of nstat is: 2
Enter the new value of nstat : 3
Setting nstat = 3

nstatdyn: Initial state (1 - ground state).
The current value of nstatdyn is: 2
Enter the new value of nstatdyn :
Setting nstatdyn = 2

dt: Time step for the classical equations.
The current value of dt (fs) is: 0.5
Enter the new value of dt (fs) : 0.1
Setting dt = _
```

Each parameter contains a short description and, most of time, an attributed default value. To use the default values, just press <ENTER>. More information about each parameter can be found in this documentation.

## 6.3 Initial conditions generation

### *Input*

- 1- Prepare geom file with equilibrium geometry. Use xyz2nx to create geom from xyz files.
- 2- Prepare force.out file with the harmonic frequencies using, for example, TURBOLOME.
- 3- Run nxinp program. Select option 1. GENERATE INITIAL CONDITIONS. nxinp will help you to select the input parameters to control de initial condition generation. Exit nxinp.

### *Run*

```
$NX/initcond.pl > initcond.log
```

### *Output*

final\_output: Initial conditions.  
initcond.log: Log file.

### *Further options*

Optionally, you can control the number of excitation quanta in each vibrational modes by providing a file qvector (see Chapter 10) together with the other input files.

It is also possible to pick points from previous dynamics calculations or to generate random velocities to be used as initial conditions.

## 6.4 Dynamics Input

### 6.4.1 Adiabatic dynamics (dynamics on one surface)

#### *Input*

- 1- Create a directory called JOB\_AD containing a set of input files for geometry optimization (1 step) with the chosen electronic structure program.
- 2- Run nxinp program. First, select option 2. SET BASIC INPUT. nxinp will help you to select the input parameters to control the dynamics. After that, select option 3. SET GENERAL OPTIONS if you want to change some more technical parameter. Exit nxinp.

#### *Hint*

Check control.dyn file. If the keyword “thres” is defined there, be sure that its value is 0 (thres = 0).

## 6.4.2 Non-adiabatic dynamics (Surface hopping)

### Input

- 1- Create a directory called JOB\_NAD containing an input for non-adiabatic coupling (single point) with the chosen electronic structure program..
- 2- Run nxinp program (`$NX/nxinp`). First, select option 2. `SET BASIC INPUT`. nxinp will help you to select the input parameters to control the dynamics. After that, select option 3. `SET GENERAL OPTIONS` if you want to change some more technical parameter. Optionally, select option 4. `SET NON-ADIABATIC DYNAMICS` to change the non-adiabatic-dynamics options. Exit nxinp.

### Hint

Check control.dyn file. If the keyword “thres” is defined there, be sure that its value is 100 (`thres = 100`).

## 6.4.3 Mixed adiabatic and non-adiabatic dynamics

### Input

- 1- Create a directory called JOB\_AD containing a set of input files for geometry optimization (1 step) with the chosen electronic structure program.
- 2- Create a directory called JOB\_NAD containing an input for non-adiabatic coupling (single point).
- 3- Run nxinp program (`$NX/nxinp`). First, select option 2. `SET BASIC INPUT`. nxinp will help you to select the input parameters to control the dynamics. Optionally, select option 3. `SET GENERAL OPTIONS` if you want to change some more technical parameter. Optionally, select option 4. `SET NON-ADIABATIC DYNAMICS` to change the non-adiabatic-dynamics options. Exit nxinp.

### Hint

Check control.dyn file. The keyword “thres” must be defined there. Its value is the energy difference threshold (eV) bellow to which the non-adiabatic dynamics starts. *This option is not fully tested.*

## 6.5 Creating the trajectory inputs

### Input

- 1- Copy final\_output file created after the initial conditions generation, section 6.3, into the same directory containing all files and directories created in the dynamics input, section 6.4.
- 2- If the jobs will run in a batch system, also copy the submission script file to that directory. In `$NX/./batch` you may find several examples of submission scripts.
- 3- Run nxinp program (`$NX/nxinp`). Select option 5. `GENERATE TRAJECTORIES AND SPECTRUM`. nxinp will help you to select the input parameters to control the trajectory inputs generation. At the end of the input selection, NEWTON-X will automatically generate the trajectory directories. This process can take some few minutes.

### Output

NEWTON-X: Newtonian dynamics close to the crossing seam

The trajectories inputs were written to `TRAJECTORIES/TRAJn`, where  $n$  is the trajectory number.

## 6.6 Running the dynamics

1- Go to each `TRAJECTORIES/TRAJn` (see section 6.5) and run

```
$NX/moldyn.pl > moldyn.log &
```

or, if is this the case, submit the job to the batch system.

2- Alternatively, go to `TRAJECTORIES` and run

```
$NX/submit.pl
```

It will allow you to automatically submit several sequential jobs to the batch system.

## 6.7 Where are the results?

1- During or after the dynamics, go to directory `TRAJECTORIES/TRAJn/RESULTS` and run:

```
$NX/plot to generate "energy x time" graph with GNUPLOT.
```

```
molden dyn.mld to see motion with MOLDEN or any visualization package (xyz format).
```

```
$NX/arrow to generate a MOLDEN file for a specific time step, containing the velocity and, in the case of dynamics with COLUMBUS, also molecular orbitals and non-adiabatic coupling vectors.
```

2- `dyn.out` file contains details about the geometry, velocity, energy and wave function (adiabatic coefficients) along the trajectory. `tprob` contains information about the hopping probability at each time step. `en.dat` contains information about the energy of each state. `sh.out` contains further information about the TDSE integration.

3- In `TRAJECTORIES/TRAJn`, the standard output (`moldyn.log`) contains the log information of the job and information about states, gradients and non-adiabatic couplings. The standard output is also written to `RESULTS/nx.log`.

4- In `TRAJECTORIES/TRAJn/DEBUG`, `runnx.error` contains occasional error messages. `log.conv` contains the information about convergence of ab initio calculations.

5- `TRAJECTORIES/TRAJn/INFO_RESTART` contains a complete set of files to restart the dynamics from the last time step that run.

## 6.8 Overview of the file structure

### 6.8.1 Dynamics simulations

The basic structure of directories and files during the dynamics simulations is shown in Figure 1.

In the case of hybrid calculations (QM/MM) the structure is very similar. The main difference is that in this case the `JOB_AD` and `JOB_NAD` directories have a substructure specifying the several jobs. The inset in Figure 1 shows an example of this substructure for a QM/MM job with COLUMBUS and TINKER.

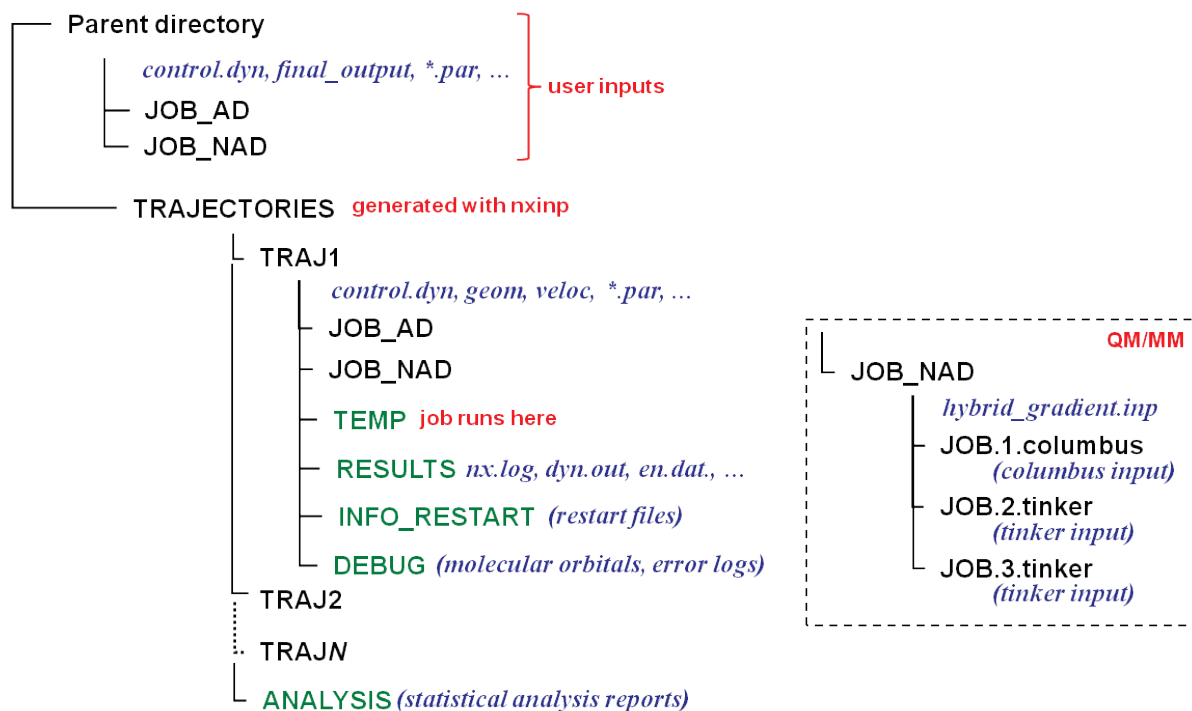


Figure 1 Basic structure of files and directories for a dynamics simulation. The inset shows an example for QM/MM simulations.

# 7 Capabilities

---

## 7.1 General features

Features	Options
Dynamics	On-the-fly dynamics with the velocity-Verlet algorithm <sup>29,30</sup> Mixed quantum classical (non-adiabatic) dynamics <sup>11</sup> Surface hopping fewest switches algorithms <sup>29,30</sup> Dynamics with non-adiabatic coupling vectors and time-derivative couplings <sup>18,19</sup> Dynamics with Local Diabatization method <sup>20</sup> Decoherence corrections <sup>26</sup> Lagrangean extrapolation of orbitals <sup>31</sup> Damped dynamics (kinetic energy always null) Andersen thermostat <sup>32,33</sup> Hybrid dynamics (QM/MM) <sup>34</sup>
Interfaces	COLUMBUS (MCSCF,MRCI) <sup>3,4</sup> TURBOMOLE (TD-DFT, RI-CC2, ADC(2)) <sup>5</sup> DFTB (TD-DFTB) <sup>6</sup> DFTB+ (DFTB) TINKER <sup>10</sup> GAUSSIAN (CASSCF, TD-DFT) <sup>7,8</sup> DFT-MRCI <sup>35</sup> GAMESS (MCSCF, CCSD(T), MP2) <sup>9</sup>
Initial conditions	Wigner distribution Quantum and classical harmonic oscillator distributions Pick points from previous dynamics Random velocity generation <sup>36</sup>
Spectrum simulation	UV photoabsorption cross section <sup>37-39</sup> UV absorption and emission spectra.
File management	Friendly input via nxinp facility Automatic management of files and directories in multiple trajectories
Output and analysis	Statistical analysis of results (internal coordinates and forces, energies, wave function properties) Normal Mode Analysis, Essential Dynamics Analysis <sup>40,41</sup> Recomputation of internal coordinates On-the-fly graphical outputs (MOLDEN, GNUPLOT)



## 7.2 Overview of the available interfaces

Program	Version	Method	Initial conditions / Spectrum		Dynamics			
			Frequency reading	Spectrum simulation	Adiabatic	Nonadiabatic (surface hopping)		
						NAC	CIO	LD
COLUMBUS	5.9	MRCI / MCSCF						
	7	MRCI / MCSCF						
TURBOMOLE		TD-DFT						
		CC2 / ADC(2)						
DFTB		TD-DFTB						
DFTB+		DFTB						
GAUSSIAN	03	CASSCF						
	09	TDDFT						
TINKER		MM						
DFT-MRCI		DFT-MRCI						
GAMESS	*	MCSCF						
		CCSD(T)/MP2						
MOLDEN		-						
ANALYTICAL		User defined						

### Available feature

NAC – Based on nonadiabatic coupling vectors.

CIO – Based on wavefunction overlaps.

LD – Based on local diabatization.

\* Initial conditions and adiabatic dynamics have been tested with GAMESS version 11 AUG 2011 (R1). Nonadiabatic jobs have been tested with a development version based on that version.

## 8 How to get NEWTON-X

---

NEWTON-X code is distributed free of charge for non-commercial and non-profit uses. To request a copy, visit [www.newtonx.org](http://www.newtonx.org).

The third-party quantum chemistry and visualization programs are not distributed with NEWTON-X. They should be directly obtained from the respective distributors and owners.

## 9 How to install NEWTON-X

---

### 9.1 Binary distribution

Binary files for a Linux-compiled NEWTON-X are distributed together with the source code. As first option, we strongly recommend to use this version. In this case, just **uncompress** the distribution file

```
tar -zxvf nx-<version>.tgz
```

or, alternatively,

```
gunzip nx-<version>.tgz  
tar -xf nx-<version>.tar
```

**Then, set the shell-variable \$NX to <install-path>/bin**

For c-shell users, this is done by:

```
setenv NX <install-path>/bin
```

For bash-shell users, this is done by:

```
export NX=<install-path>/bin
```

### 9.2 To install NEWTON-X you need

- The NEWTON-X source code. See Chapter 8 to see how to get the NEWTON-X source code.
- A FORTRAN 90 compiler installed, **preferentially the Intel Fortran** compiler.
- PERL 5.8.0 or higher installed ([www.perl.com](http://www.perl.com)).
- BLAS, LAPACK, and GSL libraries installed. More information on how to set the paths to these libraries can be found in the README.install file, which can be found in the NEWTON-X distribution.

### 9.3 To run NEWTON-X you need

- At least one of the interfaced third-party programs (COLUMBUS, TURBOMOLE, DFTB, TINKER, GAUSSIAN, GAMESS). See Chapter 7 of this documentation for an updated list of available interfaces.

For the full functionality, you also need:

- GNUPLOT ([www.gnuplot.info](http://www.gnuplot.info)).
- Some molecular visualization program able to read xyz files (e.g., MOLDEN or VMD).

### 9.4 Installation procedure

**Untar and uncompress** the source code file

```
tar -zxvf nx-<version>.tgz
```

or, alternatively,

```
gunzip nx-<version>.tgz  
tar -xf nx-<version>.tar
```

The files will be written to the NX-pack directory. Go to the **NX-pack/install subdirectory** and run the script

```
./installnx.pl
```

It will ask you for an installation address and for which packages you want to install. If the installation proceeds without problem, you will get in the screen something like:

```
=====
Welcome to the NEWTON-X installation routine
www.newtonx.org
=====

Please enter directory for the NEWTON-X installation
/home/users/NX          ! Give the path in your system

This directory seems not to exist - create it? (y/n) y
Created directory /home/users/NX

Select packages to be installed.
[0] Complete installation [Default option]
[1] Copy files and install tools (TOOLS)
[2] Initial conditions (TOOLS+INITCOND)
[3] Molecular dynamics (TOOLS+MOLDYN+MODEL)
[4] Time-dependent overlaps (TOOLS+CIOVERLAP)
Enter comma separated list or dash separated range: 0    ! Complete installation

Starting NEWTON-X installation at Mon May 17 09:27:16 CEST 2010

Compiler keyword:      ifort
Flag keyword:          -static -vec-report0
Program version:       1.2.0a (01/Jun/2010)
Fortran version:       /opt/intel/Compiler/11.1/069/bin/intel64/ifort
Perl version:          /usr/bin/perl
Original directory:    /home/users/NX-pack/install
Installation directory: /home/users/NX
BLAS:                  /opt/intel/mkl/10.1.3.027/lib/em64t/
GCC for LA libraries:  /usr/bin/g++
GCC for cioverlap:     /usr/bin/g++
BLAS for cioverlap:    /usr/lib64
LAPACK for cioverlap:  /usr/lib64
GSL for cioverlap:     /usr/lib64
GSLCBLAS for cioverlap: /usr/lib64
Packages selected:
  1 TOOLS
  2 INITCOND
  3 MOLDYN
  4 CIOVERLAP

Checking compilers, libraries and dependencies...
No problem detected for the selected packages.

Copying file .....
.....DONE
Compiling INITCOND .....DONE
Compiling MOLDYN .....DONE
Compiling TOOLS .....DONE
Compiling MODEL .....DONE
Compiling CIOVERLAP .....DONE
Setting permissions .....DONE

Installation completed!
Don't forget to set the environment variable $NX to /home/users/NX/bin
```

**Do not forget to set the shell-variable \$NX to <install-path>/bin after installation.**

For c-shell users, this is done by:

```
setenv NX <install-path>/bin
```

For bash-shell users, this is done by:

```
export NX=<install-path>/bin
```

A complete installation log is written to installnx.log file.

NEWTON-X: Newtonian dynamics close to the crossing seam

As default, `instalnx.pl` is set to use the Intel Fortran Compiler for Linux. To change the compiler and the compiler options follow the instructions contained in the `instalnx.pl` body. More information on compilers can be found in the `README.install` file, which can be found in the NEWTON-X distribution. NEWTON-X has been compiled and tested in 32-bit, EM64T, and Opteron machines.

## 9.5 Setup of third-party programs

NEWTON-X works using third-party programs, such as COLUMBUS, TURBOMOLE, GAUSSIAN, and others. These programs should be separately obtained from their distributors and installed according to their specific instructions.

Depending on the job, NEWTON-X assumes that a series of variables are defined in the system. They are:

Program	Variable	Example of file pointed by each variable
COLUMBUS	\$COLUMBUS	ls \$COLUMBUS > ... runc ...
DFTB	\$DFTB	ls \$DFTB > ... dftb ....
DFTB+	\$DFTPLUS	ls \$DFTPLUS > dftb+
GAUSSIAN 03	\$g03root	ls \$g03root/g03 > ... g03 ...
GAUSSIAN 09 <sup>1</sup>	\$g09root	ls \$g09root/g09 > ... g09 ...
DFT-MRCI	\$DFTCI	ls \$DFTCI > ... mrci ...
	\$KSCIHOME	ls \$KSCIHOME > ... dftparam.s.bhlyp ...
GAMESS	\$GAMESS	ls \$GAMESS > ... rungms ...
TURBOMOLE <sup>2</sup>	\$TURBODIR	ls \$TURBODIR > ... scripts bin ...
	\$PATH	echo \$PATH > ... :\$TURBODIR/bin/x86_64-unknown-linux-gnu:\$TURBODIR/scripts: ...
TINKER <sup>3</sup>	\$PATH	echo \$PATH > ...:/usr/bin/tinker-bin-linux:...

<sup>1</sup> In the case of GAUSSIAN 09, the GAUSSIAN profile must be sourced before executing NEWTON-X. Example (C-shell):  
source \$g09root/g09/bsd/g09.login

<sup>2</sup> For TURBOMOLE, directories containing the executables (e.g., \$TURBODIR/bin/x86\_64-unknown-linux-gnu) and scripts (e.g., \$TURBODIR/scripts) must be in the standard path.

<sup>3</sup> For TINKER, directories containing the executables (e.g., /usr/bin/tinker-bin-linux) must be in the standard path.

## 9.6 Verification of installation

To test the installation, create a directory to execute the tests. Move into it and select the tests that you want to perform by running

```
$NX/inp-testnx.pl
```

It will return a list of all available tests. You can select particular tests or all of them. It will create a file called test.inp containing a space-separated list with the number of the tests that should run.

Then, run the program

```
$NX/test-nx.pl > test-nx.log &
```

This program will run the selected tests. If test.inp is not present, all tests are selected by default. The tests are a series of quick pre-build examples. It will check whether they have normal termination or not, and it will compare the results to those of standard files. Check the results in test-nx.log.

Tests invoking GAMESS may need special adjustments for each system. When such tests are requested, inp-testnx.pl asks additionally a few questions about GAMESS environment (executable number, scratch directory, and execution script name). This information is written to games.par file, which is used by the jobs running under test-nx.pl command.

## 9.7 Compatibility between DALTON and NEWTON-X installations

Even after a successful installation, NEWTON-X jobs using COLUMBUS may return the following error message:

```
readsifs returned with error! Are the integer formats between Columbus and NX matching?
```

In such cases, the probable cause is an incompatibility issue between DALTON program (used by COLUMBUS) and NEWTON-X. It can be solved by recompiling CIOVERLAP program, by the following these steps:

1. Go to install directory.
2. Edit installnx.pl file and change the value of variable \$status\_record32. The default value is "on" (SIFS integer format = 64 bits), which means that DALTON is assumed to have been compiled for 64 bits. "off" (SIFS integer format = 32 bits) means that DALTON is assumed to have been compiled for 32 bits.

3. Run the installation program again, but this time select option  
[4] Time-dependent overlaps (TOOLS+CIOVERLAP).
4. **It is not necessary to recompile LA library.**  
A previous installation of the LA library already exists.  
Do you want to replace it? (y/n) **n**

## 10 Initial conditions generation

INITCOND is a set of programs developed to generate initial conditions to the molecular dynamics and spectrum simulation.

It is possible to generate initial conditions by sampling the data according to harmonic oscillator distributions (classical and quantum). It is also possible to pick at random points from a previous dynamics calculation performed with NEWTON-X or to generate random velocities for some specific geometry.

For NACT = 1, 2 or 3 the initial conditions are based on normal modes. The normal modes themselves should be generated in a separated run and given as input to NEWTON-X. NEWTON-X can read the normal modes directly from the output files of several third-party programs (see IPROG below).

Despite the initial format, the normal modes are internally transformed into mass-weighted normal modes **L**.

For NACT = 1 or 2, for each normal mode, an amplitude  $Q_n$  and a momentum  $\dot{Q}_n$  are randomly selected from a harmonic oscillator distribution. The Cartesian coordinates and velocities of all atoms are then determined from the normal coordinates by the inverse transformation<sup>41</sup>. If NACT = 3, only  $Q_n$  is randomly selected, while  $|\dot{Q}_n|$  is scaled as to give the harmonic oscillator energy. For NACT = 2, if the vibrational numbers are zero, the distributions matches the Wigner distribution for the quantum harmonic oscillator<sup>42,43</sup>. For higher vibrational quantum numbers, the distribution for each normal mode is

$$|\psi_{N_n}(Q_n)|^2 |\xi_{N_n}(P_n)|^2,$$

where  $\psi_{N_n}(Q_n)$  and  $\xi_{N_n}(P_n)$  are respectively the harmonic oscillator wavefunctions in the coordinate and momentum spaces.

If NACT = 4, random points are picked from a previously performed dynamics. If NACT = 5, random velocities are generated for a fixed velocity according to the algorithm described in Ref.<sup>36</sup>. If NACT = 6, single point electronic structure calculations are performed for a previously computed set of initial conditions.

### 10.1 How to execute INITCOND

Run

```
$NX/initcond.pl > initcond.log
```

### 10.2 Input parameters

File: *initqp\_input*

Namelist: *DAT*

Parameter	Default	Description
NUMAT	= [3]	number of atoms



NPOINTS	= [1]	number of initial conditions generated
NACT	= [2]	<p>1 - Each normal mode contains a given energy in the initial state, according to the vibrational quantum number, and the distribution of coordinates and momenta is that of a classical harmonic oscillator.</p> <p>2 - The distribution of coordinates and momenta is that of a quantum mechanical harmonic oscillator in the specified vibrational state. The sample of the coordinates and momenta is uncorrelated. In the ground vibrational state, this distribution matches the Wigner distribution for the harmonic oscillator.</p> <p>3 - The distribution of coordinates and momenta is that of a quantum mechanical harmonic oscillator in the specified vibrational state. Only coordinates are sampled. For each normal mode, the momenta are scaled to the coordinates up to the have the correct harmonic vibrational energy.</p> <p>4 - NPOINTS points are picked up from some previously run dynamics simulations.</p> <p>5 - Gaussian-distributed random velocities are generated for some specific geometry.</p> <p>6 - Single point electronic structure calculations for a previously computed set of initial conditions.</p>
ISEED	= [1]	<p>Random number seed</p> <p>0 Use standard value for the random number seed.</p> <p>-1 Use random seed.</p> <p>Any integer &gt; 0 is used as the seed itself.</p> <p>Use ISEED = -1 if the job will be split (see section 15.7).</p>
LVPRT	= [1]	<p>1 - Standard print level.</p> <p>2 - Debug print level.</p>
<b>If NACT ≤ 4</b>		
N_PICK	= [-1]	<p>Method to pick up points:</p> <p>-1 - pick random points from trajectories.</p> <p>n (integer) - pick time step n from trajectories.</p>
NIS	= [1]	Initial state (Ground state = 1).
NFS	= [2]	Final state. (Information for all states between NIS and NFS are stored.)
CHK_E	= [0]	<p>0 - Do not check the energies.</p> <p>1 - Check the energies between states NIS and NFS.</p> <p>For ground state dynamics, set CHK_E = 0.</p>
CMP_E	= [0]	<p>(Only if NACT = 4 and CHK_E = 1)</p> <p>0 - Read energies from dynamics output</p> <p>1 - Compute energies.</p>
PROG	= [1.0]	<p>(Only if CHK_E = 1 and NACT = 3 or CMP_E = 1 and NACT = 4)</p> <p>Program to compute vertical excitation energies:</p> <p>1.0 - COLUMBUS</p> <p>2.0 - TURBOMOLE RI-CC2 / ADC(2)</p> <p>2.1 - TURBOMOLE TD-DFT</p> <p>5.0 - DFTB</p> <p>9.0 - DFT-MRCI</p> <p>10.0 - GAMESS (MCSCF)</p> <p>10.1 - GAMESS (other methods)</p>
EVERT	= [5.0]	Required vertical energy (eV).
DE	= [0.5]	Allowed variation EVERT (eV).

**If NACT ≤ 3**

NEWTON-X: Newtonian dynamics close to the crossing seam

FILE_GEOM	= [geom]	file containing the equilibrium geometry of the molecule in COLUMBUS and NEWTON-X format. (See section 15.6 for conversion tools)
FILE_NMODES	= [nmodes]	file containing the normal coordinates and the vibrational frequencies (in $\text{cm}^{-1}$ ). Normally it is an output file from some quantum chemistry program (see option IPROG).
FILE_OUT	= [ini_qv]	output file from INIQP, containing NPOINTS set of generated Q and V.
FILE_VIB	= [qvector]	file with the vibrational quantum numbers. See description below.
ANH_F	= [1.0]	Multiply all frequencies by ANH_F. Useful to add anharmonic effects.
KVERT	= [1]	0 - Use the provided EVERT. 1 - Use EVERT of the equilibrium geometry that is calculated in the first step of the program.
IPROG	= [1]	Read vibrational modes from: 1 - GAMESS 2 - TURBOMOLE 3 - COLUMBUS 4 - GAUSSIAN 5 - MOLDEN 6 - DFTB 7 - ACES2
NM_FLAG	= [0]	(Only if IPROG = 5) If normal modes should be read from a MOLDEN file, then the type of normal mode should be given. 1 - Cartesian normal modes ( $\text{amu}^{-1/2}$ ) 2 - Normalized Cartesian normal modes 3 - Mass weighted normal modes The default 0 will terminate the program.
<b>IF NACT = 4</b>		
ADDRESS	= []	Path to the set of previous trajectories from which the initial conditions must be picked. The default address [/home/old_dyn/TRAJECTORIES] certainly should be changed.
TRAJI	= [1]	Initial trajectory to be read.
TRAJF	= [10]	Final trajectory to be read.
TI	= [0]	(Only if N_PICK = -1) Start to pick points only after time TI (fs) of each trajectory.
TF	= [-1]	(Only if N_PICK = -1) Do not pick points after time TF (fs) of each trajectory. Diagnostic.pl (see section 15.5) is always called in this kind of run. The time suggested in its output (diag.log) is used if it is smaller than TF. If TF = -1, always use the time suggested in diag.log file.
REORDER	= [1]	(Only if N_PICK = -1) Sort the points according to the trajectory number and time step. 0 - Do not sort. 1 -Sort.
ETOT_DEV	= [0.5]	Allowed variation in the total energy (eV).
POP_DEV	= [0.1]	Allowed variation in the norm of the adiabatic population.
<b>IF NACT = 5</b>		
FILE_GEOM	= [geom]	file containing the initial geometry of the molecule in the COLUMBUS and NEWTON-X format. (See section 15.6 for conversion tools)
EKIN	= [0]	Kinetic energy (eV).

NEWTON-X: Newtonian dynamics close to the crossing seam

TEMP = [0] Temperature (K).  
 If TEMP = 0 (default), this option generates random velocities corresponding to kinetic energy  $E_{kin}$  (microcanonical ensemble).  
 If TEMP > 0, canonical ensemble of random velocities is generated with mean kinetic energy  $E_{KIN}$  and standard deviation  $\sigma = k_B T (3N/2)^{1/2}$ .  
 In any case, translational and rotational velocities are zero.

**IF NACT = 6**

NIS = [1] Initial state (Ground state = 1).  
 NFS = [2] Final state. (Information for all states between NIS and NFS are stored.)  
 PROG = [1.0] Program to compute vertical excitation energies:  
 1.0 - COLUMBUS  
 2.0 - TURBOMOLE RI-CC2 / ADC(2)  
 2.1 - TURBOMOLE TD-DFT  
 5.0 - DFTB  
 6.5 - GAUSSIAN 09  
 9.0 - DFT-MRCI  
 10.0 - GAMESS (MCSCF)  
 10.1 - GAMESS (other methods)  
 20 - Hybrid energies (QM/MM)

## Example of initqp\_input file:

```
&DAT
nact=      2,
numat=     2,
npoints=   20,
file_geom= hf.geom,
file_nmodes=hf.nmodes,
file_out=   qvector,
evert=     5.0,
de=        0.25,
kvert=     1,
chk_e=     1,
prog =     1.0,
iprogram = 2.0,
nis =      1,
nfs =      2,
&END
```

## File : columbus.par

Optional file containing information for COLUMBUS jobs.

Parameter	Default	Description
MEM	[200]	COLUMBUS core memory in Mwords (1 GB = 134 Mwords). For COLUMBUS 7, MEM is internally converted to GB.

## File : turbomole.par

Optional file containing information for TURBOMOLE jobs.

Parameter	Default	Description
PARALLEL	[1]	Number of cores to use for parallel Turbomole (SMP only, no MPI!)

## File : dftb.par

Optional file containing information for DFTB jobs.

Parameter	Default	Description
DFTB_EXEC	= [dftb]	Name of the DFTB executable file. \$DFTB variable must be defined in the system. NEWTON-X will run \$DFTB/<dftb_exec>.

## File : dftb+.par

NEWTON-X: Newtonian dynamics close to the crossing seam

Optional file containing information for DFTB+ jobs.

Parameter	Default	Description
DFTBP_EXEC	= [dftb+]	Name of the DFTB+ executable file. \$DFTBPLUS variable must be defined in the system. NEWTON-X will run \$DFTBPLUS/<dftb_exec>.

File : dftci.par

Optional file containing information for DFT-MRCI jobs.

Parameter	Default	Description
maxiter	= [6]	Maximum number of MRCI cycles.

File: gamess.par

Parameter	Default	Description
VERNO	= [00]	Version number of GAMESS executable.
NCPUS	= 1	Number of computer processes to be run.
RUN_GAMESS	= [rungms]	Name of GAMESS execution script.
SCR	= [1]	GAMESS scratch directory: 0: Use default GAMESS options as defined in RUN_GAMESS script. 1: Create SCR directory inside TEMP directory of NEWTON-X.

### 10.3 What you need to execute

- initqp\_input with the set of parameters. (It can be generated with nxinp tool.)

#### 10.3.1 Additional files

If  $NACT \leq 3$ :

- file with the equilibrium geometry (see FILE\_GEOM keyword). Hint: the program tm2nx converts the coord file (TURBOMOLE format) to the NEWTON-X format.
- file with the vibrational modes (see FILE\_NMODES keyword).  
**Warning 1:** The atom order in FILENMODES and in FILE\_GEOM must be the same.  
**Warning 2:** The orientation of the normal mode coordinates in FILENMODES orientation must be consistent with the orientation of the optimized geometry in FILE\_GEOM.
- file with the vibrational quantum numbers (see FILE\_VIB keyword). The default is the ground state (0 quantum in each mode). If you want to change this default values, write to FILE\_VIB a list of quanta in each mode. Example:  

```

0 0 0 0 0
0 0 0 0 0
0 1

```

This example puts one quantum at the highest frequency mode of a system with 12 modes (6 atoms). Keep the format with 5 columns of integers. The order of the modes is the same as in FILE\_NMODE. If FILE\_VIB does not exist, the program assumes 0 for all modes. To set -0.5 for a mode makes this mode contribute with 0 to the initial energy.
- JOB\_AD directory containing input files for excited-state single point calculation with a third-party quantum chemistry program (only if CHK\_E=1).
- save\_file, optional file (see section 10.3.3).

NEWTON-X: Newtonian dynamics close to the crossing seam

If NACT = 4:

- The TRAJn directories of some dynamics calculations previously performed.

If NACT = 5:

- Geometry file with the initial geometry in NEWTON-X format.

If NACT = 6:

- A previously computed set of initial conditions in the NEWTON-X format renamed final\_output.old.
- JOB\_AD directory containing input files for excited-state single point calculation with a third-party quantum chemistry program.
- save\_file, optional file (see section 10.3.3).

### 10.3.2 The JOB\_AD directory

The JOB\_AD directory contains input files for excited-state single point calculation with a third-party quantum chemistry program.

For the cases that this directory is needed, set the oscillator strength to be calculated by the quantum chemistry program. This will give you the possibility of generating UV spectra as well as to use the transition probability to select the initial conditions.

**Specifically for COLUMBUS calculations (PROG = 1.0):** Input files must be prepared for a single point MCSCF or CI calculation in C<sub>1</sub> point group (no symmetry).

**For DFTB calculations (PROG = 5.0):** Input files must be prepared for excited-state spectrum calculation using code 10. The DFTB input parameter and geometry files must be named dftb.in and in.gen, respectively.

**For Gaussian 09 (PROG = 6.5):** Prepare input files for a single point vertical excitation at the TDDFT level. The input file must be called gaussian.com and the geometry must be given in Cartesian coordinates. The check point file named gaussian.chk containing the initial molecular orbitals may be provided as well. A suitable example of gaussian.com content is:

```
%chk=gaussian
%mem=2000MB
# B3LYP/6-31G(d) TD=NStates=2 NoSymm

methaniminium

1 1
N 0.000000 0.000000 0.637342
C 0.000000 0.000000 -0.703614
H -0.648747 0.572349 1.177883
H 0.648745 -0.572350 1.177883
H 0.618844 0.701080 -1.278050
H -0.618844 -0.701080 -1.278049
```

Note that the geometry is given in Angstroms in this version of NEWTON-X. Up to NEWTON-X 1.3, the keyword Unit=AU was mandatory. To avoid troubles with differences between input and standard

orientations in Gaussian, it is highly recommended to use NoSymm keyword in the Gaussian.com as well as in the computation of the normal modes (FILE\_NMODES).

NEWTON-X executes GAUSSIAN 09 by invoking the command:

```
g09 < gaussian.com
```

It is assumed that the user sources the g09 profile before running NEWTON-X.

**For DFTB+ calculations (PROG = 8.0):** Input files must be prepared for a DFTB calculations yielding energy and gradient. The DFTB input parameter and geometry files must be named dftb.in and in.gen, respectively. At the present, only ground state dynamics (no TD) is available with this program.

**For DFT-MRCI calculations (PROG = 9.0):** JOB\_AD should contain a complete set of input files for TURBOMOLE DFT calculation, including the auxiliary basis set, and the specific input file for the DFT-MRCI program. The DFT-MRCI input file must be named mrci.inp and should be set to the total number of excited states.

### 10.3.3 Save files option

When electronic structure calculations are performed during the initial condition generation ( $\text{NACT} \leq 3$  with  $\text{CHK\_E} = 1$  or  $\text{NACT} = 6$ ), it is possible to give a list of files or directories that should be saved for every initial condition. This is simply done by creating a file name save\_file and including a list of file names to be saved. save\_file must contain only one file name per line. For instance, you may want to save the molecular orbitals (mos file) and the scf log file (dscf.out) resulting from the TURBOMOLE calculations for each initial condition. In this case the following save\_file should be given together with the other input files:

Example of save\_file file:

```
dscf.out
mos
```

The required files are tar-compressed and written to the DEBUG directory renamed as ic<card>-filename.tgz, where <card> is the number of the initial condition as given in the final\_output file. For the example above, DEBUG will contain:

```
ic0-dscf.out.tgz
ic0-mos.tgz
ic1-dscf.out.tgz
ic1-mos.tgz
ic2-dscf.out.tgz
ic2-mos.tgz
...
ic0-<NPOINTS>.out.tgz
ic0-<NPOINTS>.tgz
```

The routine that saves the files runs right after the execution of the third-party program, before the cleanup of execution directory.

If the file that should be saved is inside a directory, the relative path must be provided. For instance, if we want to save the MCSCF log file of COLUMBUS jobs (mcsfsm.sp), save\_file should contain "LISTINGS/mcsfsm.sp". The full LISTINGS directory will be saved if save\_file contains "LISTINGS".

## 10.4 Output

The initial conditions are written to final\_output file. The data from spectrum simulation are written to the cross-section.dat, spectrum.dat and spectrum-hist.dat files.

NEWTON-X: Newtonian dynamics close to the crossing seam

Initial condition for multiple states are written to `final_output.[initial-state].[final-state]`. Thus, for example, `final_output.1.3` contains initial conditions for transitions from state 1 (ground state) to state 3. The `final_output` file is the same as `final_output.[NIS].[NFS]`.

DEBUG directory contains the error log information (`runnx.error`) and files that should be saved when `save_file` file is provided (see section 10.3.3).

## 10.5 Running in several computers

It is possible to split the spectrum and initial condition generation job to run in different computers and merge them again afterwards. See section 15.7 for general instructions.

Do not forget to set `ISEED = -1` when the jobs are split.

# 11 Trajectory-input generation, initial conditions for multiple states, and spectra

---

## 11.1 NXINP and options in the MAKEDIR.PL program

The section “5. GENERATE TRAJECTORIES AND SPECTRUM” of nxinp input program allows multiple different tasks, which can be selected in the sub menu:

```

=====
                          NEWTON-X
Newtonian dynamics close to the crossing seam
                          www.newtonx.org
=====

                          GENERATE TRAJECTORIES AND SPECTRUM

type:  What do you want to do?
       1 - Generate spectrum
       2 - Select initial conditions for multiple initial states
       3 - Generate trajectories
       4 - Return to main menu
The current value of type is: 3
Enter the new value of type  :
```

After selecting one of these options, you will be asked a series of question about the specificities of your job. At the end, the file mkd.inp will be generated and the program makedir.pl will be automatically running when you see the message:

```
Processing data: This may take some minutes. Please, wait...
```

If you want, after leaving nxinp, makedir.pl can be executed again by running:

```
$NX/makedir.pl > makedir.log
```

The options in makedir.pl may be changed by setting the following keywords in mkd.inp.

*Name: mkd.inp*

Parameter	Default	Description
TYPE	= [3]	1 - Generate spectrum 2 - Select initial conditions for multiple initial states 3 - Generate trajectories
NIS	= [1]	Initial state. 1 is the ground state. (Spectrum simulation.)
NFS	= [2]	Array of final states (space separated, e.g., 2 3 4). If not only spectrum, but trajectories input is required as well, only one final state is allowed.
SCREEN	= [0]	Energy restriction:

NEWTON-X: Newtonian dynamics close to the crossing seam



		0 - don't apply any restriction 1 - use the original energy restriction written in the final_output files 2 - apply new energy restriction
E_CENTER	= [0.0]	Center of the energy restriction (Only if SCREEN = 2) x - value of the center of restriction (eV) ref n - use the vertical excitation of final_output.nis.n file
E_VAR	= [0.5]	Width of the energy restriction. (eV) (Only if SCREEN = 2)
OS_CONDON	= [-1]	Oscillator strength: -1 - try to read from final_output file x - oscillator strength is always x (Condon approximation)
PROB_KIND	= [F]	Probabilities in the spectrum generation will be computed according to A - Einstein coefficients A (spontaneous emission) B - Einstein coefficients B (induced absorption or induced emission) F - Oscillator strength (photoabsorption cross section)
NORM	= [local]	Normalization of the Einstein's coefficients: local - Use energy-restricted data set global - Use complete data set
SEED	= [0]	Seed for the random number generation 0 - a default random number seed is used 1 - a randomized seed is used Any other positive integer is used itself as the random number seed.

IF TYPE = 1. The next three keywords define the parameters for the Gaussian broadening method.

L_SHAPE	= [lorentz]	Line shape: gauss - Normalized gaussian function lorentz - Normalized Lorentzian function
DELTA	= [0.05]	Phenomenological broadening of the spectrum using the Gaussian method <sup>1</sup> (in eV). The default $\Delta$ corresponds to a small broadening, equivalent to 0.5 nm.
TEMP	= [0]	Temperature correction for the spectrum (in K).
NREF	= [1]	Refraction index
EPS	= [0.005]	Distance between consecutive points in the spectrum using the Gaussian method. (in eV)
KAPPA	= [3]	$\kappa$ (integer) is used to define the range of the spectrum between $\Delta E_{\min} - \kappa \Delta$ and $\Delta E_{\max} - \kappa \Delta$ .

TYPE = 3. The next three keywords define the batch system behaviour.

TITLE	= [nx]	Title. Useful if the job runs in batch.
SUBFILE	= [pmold]	If the job run in batch, SUBFILE is the name of the submission script.
BATCHDEF	= [-N]	Definition of the batch system. Default is SGE but NEWTON-X attempts to read it from SUBFILE and change it when some suitable value is found.

## 11.2 Transition spectra

NEWTON-X computes the spectrum by means of the line broadening method<sup>1,39</sup>, by assigning to each initial condition a line shape function  $g$  with the height  $P$  and a width representing some phenomenological broadening ( $\Delta$ ) and plotting the sum ( $S$ ) of these line shape functions as a function of the transition energy  $E$ , i.e.,

NEWTON-X: Newtonian dynamics close to the crossing seam

$$S(E) = \sum_{n=1}^{N_{\text{points}}} P_n(\Delta E_n) g(E - \Delta E_n).$$

$E$  (eV) and  $S(E)$  (arbitrary unities) are written to spectrum.dat file. The line shape options are the Gaussian function

$$g_{\text{Gauss}}(E - \Delta E_{il}, \delta) = \left(\frac{2}{\pi}\right)^{1/2} \frac{\hbar}{\delta} \exp\left(-\frac{2(E - \Delta E_{il})^2}{\delta^2}\right)$$

or the Lorentzian function

$$g_{\text{Lorentz}}(E - \Delta E_{il}, \delta) = \frac{\hbar\delta}{(2\pi)} \left[ (E - \Delta E_{il})^2 + (\delta/2)^2 \right]^{-1}.$$

The intensity of each line  $P$  can be taken as the Einstein coefficient  $A$  and  $B$  or the oscillator strength  $f$ . Absolute values for the Einstein's coefficients<sup>44</sup> can be obtained from the log information written in makedir.log. First, note that the Einstein's coefficients are given by

$$A = C_A f \Delta E^2,$$

$$B = C_B A \Delta E^{-3},$$

with  $C_A = 2\pi e^2 / (\hbar \varepsilon_0 m c^3)$  and  $C_B = c^3 \hbar^2 / 8\pi$ . As usual,  $e$ ,  $\hbar$ ,  $m$ ,  $c$  and  $\varepsilon_0$  are, respectively, fundamental charge, Planck's constant, electron mass, speed of light, and free-space permittivity. Therefore, to compute the absolute values of the Einstein's coefficients, first check the units of vertical excitation energy in final\_output file. If they are in eV, then multiply coefficients  $A$  by  $0.43392 \times 10^8$  or  $B$  by  $0.11445 \times 10^{15}$ . If they are in hartree, then multiply coefficients  $A$  by  $0.32130 \times 10^{11}$  or  $B$  by  $0.56800 \times 10^{10}$ . The final units of  $A$  will be  $s^{-1}$  and of  $B$  will be  $m^3 s^{-2} J^{-1}$ .

Note that the algorithm assumes the same degeneracy factor for both states ( $g_i = g_j$ ) and refraction index  $n = 1$ . When spectrum involving multiple states is selected by attributing several values to the NFS array, final\_output.nis.isf files (ISF is each value in NFS array) for each transition must be present.

The file spectrum.dat is composed of two columns:  
 $E$  (eV)     $S$  (arbitrary units)

When the intensity is selected to be  $f$  (PROB\_KIND = F), in addition to the spectrum computed as above, the photoabsorption cross section is computed as well and written to the file cross-section.dat. The cross section is given by<sup>39</sup>

$$\sigma(E) = \frac{\pi e^2 \gamma}{2mc\varepsilon_0 n} \sum_{l \neq i}^{N_{fs}} \left[ \frac{1}{N_p^l} \sum_k^{N_p^l} f_{il}(\mathbf{R}_k) g(E - \Delta E_{il}(\mathbf{R}_k), \delta) \right],$$

where the internal sum runs over each of the NPOINTS initial coordinate  $\mathbf{R}_k$  and the external sum runs over the several states (NFS). The factor  $\gamma$  is given by<sup>37</sup>

$$\gamma = 1 - \exp(-E / k_B T).$$

In order to compare the simulations to the experiments it useful to remind that photoabsorption cross sections ( $\text{cm}^2$ ) and molar extinction coefficients ( $\varepsilon$  in  $\text{M}^{-1}\text{cm}^{-1}$ ) can be interconverted by the relation<sup>45</sup>

$$\sigma = 10^3 \ln(10) \frac{\varepsilon}{N_A},$$

where  $N_A$  is the Avogadro's number.

The error in the cross section due to the statistical sampling can be estimated by

$$\delta\sigma(E) \approx \frac{\pi e^2 \hbar \gamma}{2mc\varepsilon_0 n} \sum_{l \neq i}^{N_{fs}} \frac{1}{N_p^{1/2} (N_p - 1)^{1/2}} \left[ \sum_k^{N_p} (f_{il}(\mathbf{R}_k) g_L(E - \Delta E_{il}(\mathbf{R}_k), \delta) - \langle s_l \rangle)^2 \right]^{1/2},$$

where

$$\langle s_l \rangle = \frac{1}{N_p} \sum_{k'}^{N_p} \Delta E_{0l}(\mathbf{R}_{k'}) f_{il}(\mathbf{R}_{k'}) g_L(E - \Delta E_{il}(\mathbf{R}_{k'}), \delta).$$

The file cross-section.dat is composed of four columns:

$E$  (eV)  $\lambda$  (nm)  $\sigma$  ( $\text{\AA}^2 \cdot \text{molecule}^{-1}$ )  $\delta\sigma$  ( $\text{\AA}^2 \cdot \text{molecule}^{-1}$ )

During spectrum and initial condition generation, the  $\Delta E$  energies (eV) of all accepted initial conditions are written to spectrum-hist.dat file. A histogram produced with this data will also give the absorption spectrum. The graphical tool is available to produce simple histograms. From the directory containing spectrum-hist.dat, call `$NX/spectrum.pl arg`, where arg is an optional argument with the number of bins in the histogram. If ARG is not given, the bin width is assumed to be 0.05 eV. The program produces a simple GNUPLOT histogram that represents the spectrum. (Counts x eV) Average and standard deviation values are written to hist.out. Frequency table is written in hist.dat.

### 11.3 Initial conditions for multiple states

In Chapter 10 it was explained how to generate initial conditions for starting the dynamics in a single excited state. That procedure may be generalized to create initial conditions for multiple adiabatic states, weighting each one according to their oscillator strength.

Having the final\_output.[nis].[nfs] files in the same directory, run nxinp and select option

```
2 - Select initial conditions for multiple initial states
```

The energy restriction may be changed or not. After running makedir.pl, a new directory called `SELECTED_INITIAL_CONDITIONS` is created. Inside this directory, there are new final\_output.[nis].[nfs] files containing initial conditions for each nfs state. These files can be individually used to start the dynamics in each nfs state.

Suppose you want to generate initial conditions to start the dynamics simultaneously from  $S_2$  and  $S_1$  states. `SELECTED_INITIAL_CONDITIONS` directory should contain the files final\_output.1.2 and final\_output.1.3. The number of initial conditions in each one reflects the dipole transition probability from  $S_0$  into these two states. If, for example, if  $S_1$  state has  $\pi\pi^*$  character and  $S_2$  state has  $n\pi^*$  character, the number of points in final\_output.1.2 will be substantially larger than in final\_output.1.3. When the dynamics is performed the number of trajectories starting in  $S_1$  and  $S_2$  states should reflect this proportion.

### 11.4 Managing several trajectories

For managing several trajectories, it is useful to use the script makedir.pl. makedir.pl reads final\_output file generated by the initial condition generation procedure, and writes a new NEWTON-X input directory for each initial condition accepted.

After preparing the inputs using nxinp tool, be sure that you have in the same directory:

- final\_output (always)
- control.dyn (always)
- JOB\_AD (if necessary for the specific job)
- JOB\_NAD (if necessary for the specific job)
- jiri.inp (if necessary for the specific job)
- sh.inp (if necessary for the specific job)

- wf.inp (if necessary for the specific job)
- <third-party program>.par (if necessary for the specific job)
- therm.inp (if necessary for the specific job)
- freeze.inp (if necessary for the specific job)
- therm.freeze (if necessary for the specific job)
- boundaries.inp (if necessary for the specific job)
- mkd.inp (optional)
- pmold (if available, see below)

Run

```
$NX/makedir.pl > makedir.log
```

If TYPE = 3, makedir.pl will create the directories TRAJECTORIES/TRAJi, where i is the number of the accepted condition.

## 11.5 About energy restrictions

If energy restrictions are applied (SCREEN = 2) or spectrum generation is chosen (TYPE = 1), vertical excitation energy ( $\Delta E$ ) and the oscillator strength ( $f$ ) are collected from each initial condition in final\_output.nis.isf or final\_output files (where isf is each value in nfs array). For each initial condition within  $E\_CENTER \pm E\_VAR$ , the quantity  $f$ ,  $A = f \cdot \Delta E^2$  or  $B = A \cdot \Delta E^{-3}$  proportional to the oscillator strength or Einstein's coefficients is computed according to PROB\_KIND keyword. The normalization may be done by the maximum  $A$  or  $B$  values within the window restriction (NORM = local) or within the complete set of values (NORM = global). The relative "transition probability"  $P = f/f_{max}$ ,  $A/A_{max}$  or  $P = B/B_{max}$  is compared to a random number in order to select whether the initial condition will be accepted or not.

## 12 Dynamics inputs

---

### 12.1 What is necessary to run

For the input you need the following files:

Always:

control.dyn - with parameters to control de dynamics.

geom - with initial geometry.

veloc - with initial velocity.

Depending on the settings in control.dyn:

jiri.inp - with parameters to control non-adiabatic dynamics.

sh.inp - with parameters to control non-adiabatic dynamics.

<third-party>.par - with third-party program options.

therm.inp - with thermostat options.

freeze.inp - with frozen atoms information.

therm.freeze - with atoms not affected by thermostat.

boundaries.inp - with boundary condition options.

NX\_analysis - with instructions for customized analysis of the results.

stopsign.inp - with instructions to terminate the simulations.

You also need the input files for the program that will calculate the energies, gradients and non-adiabatic couplings.

JOB\_AD - directory containing input files for adiabatic calculation. Energy and gradient for one state.

JOB\_NAD - directory containing input files for non-adiabatic calculations. Energy, gradient and non-adiabatic coupling.

For adiabatic dynamics or non-adiabatic dynamics using time-derivative couplings<sup>18</sup>, only JOB\_AD is necessary.

For mixed, non- and adiabatic dynamics (using the THRES keyword), both JOB\_AD and JOB\_NAD are necessary.

For non-adiabatic dynamics using non-adiabatic coupling vectors, only JOB\_NAD is necessary.

#### 12.1.1 control.dyn

File control.dyn should contain the main parameters for the dynamics:

*File: control.dyn*

*Namelist: input*

Parameter	Default	Description
NAT	= [3]	Number of atoms.
NSTAT	= [2]	Number of states (dynamics will be performed on the highest one).

NEWTON-X: Newtonian dynamics close to the crossing seam

		NSTAT $\geq$ NSTATDYN.
NSTATDYN	= [nstat]	Initial state.
NDAMP	= [0]	0 : normal dynamics. 1 : velocity is dumped to zero at each time step.
KT	= [1]	Print output at each KT steps.
DT	= [0.5]	Time step (fs).
T	= [0.0]	Initial time (fs).
TMAX	= [10.0]	Maximum time (fs).
NINTC	= [3*Nat-6]	Number of internal coordinates to read from the ab initio program. Change default if the system is linear or has redundant coordinates.
MEM	= [200]	Core memory for the ab initio program (Mwords). (Relevant if PROG = 1)
KILLSTAT	= [1]	Finish dynamics if after a hopping the system remains more than timekill fs on state KILLSTAT.
TIMEKILL	= [0]	See KILLSTAT. 0 : deactivate KILLSTAT and TIMEKILL
PROG	= [1]	Program to compute energies, gradients and non-adiabatic coupling vectors (if available): 0 : Analytical model (see section "Using analytical models") 1 : COLUMBUS 2.0 : TURBOMOLE RI-CC2 / ADC(2) 2.1 : TURBOMOLE TD-DFT 5 : DFTB 6 : GAUSSIAN 7 : TINKER 8 : DFTB+ 10.0 : GAMESS 10.1 : GAMESS ADIABATIC 20 : Hybrid jobs
THRES	= [0 for PROG = 2.x and 5] = [100 for PROG = 0, 1 and 6]	Energy difference threshold to initiate non-adiabatic dynamics (eV).
LVPRT	= [1]	Amount of output to print and output files to keep: 0 : Minimal level 1 : Normal level 2 : Debug level (huge amount of data)
ETOT_JUMP	= [0.5]	Kill trajectory if total energy deviate more than ETOT_JUMP eV in one time step.
ETOT_DRIFT	= [0.5]	Kill trajectory if total energy deviate more than ETOT_DRIFT (eV) in comparison to the value in t = 0.
NXRESTART	= [0]	Restart options: 0 : new job. 1 : restart job using the content of INFO_RESTART directory. See details in Section 12.4.

---

[default values] written in inp.f90.

#### Example 1:

```
&input
Nat   = 6           ! Number of atoms
nstat = 3           ! Number of states
nstatdyn = 3       ! Initial state
```

NEWTON-X: Newtonian dynamics close to the crossing seam

```

dt      = 0.5      ! Time step (fs)
tmax    = 200.0    ! Total time (fs)
prog    = 2.1      ! Dynamics with TURBOMOLE (TD-DFT).
/&end      ! Do not forget &input and /&end.

```

Example 2: Ethylene  $S_1$ -state dynamics, calculating  $S_0$ ,  $S_1$  and  $S_2$  energies. Time step of 0.1 fs during 200 fs and output printed at each 2 fs. 100,000,000 words of memory. If potential energy difference between two consecutive states drops below 2.0 eV, start non-adiabatic dynamics. Finish dynamics if after a hopping to  $S_0$ , the system remains more than 10 fs on this state. Get energies, gradients and non-adiabatic couplings with COLUMBUS.

```

&input
Nat      = 6
nstat    = 3
nstatdyn = 2
kt       = 20
dt       = 0.10
tmax     = 200.0
mem      = 100
thres    = 2.0
killstat = 1
timekill = 10.0
/&end

```

Do not forget "&input" and "/&end".

## 12.1.2 Geometry

The geometry input is (free format, au):

*Name: geom*

```

Symbol_1  Z_1  x_1  y_1  z_1  M_1
Symbol_2  Z_2  x_2  y_2  z_2  M_2
:
Symbol_nat Z_nat x_nat y_nat z_nat M_nat

```

where  $Z$  is the nuclear charge,  $x,y,z$  are the Cartesian coordinates, and  $M$ , the atomic masses.

Example:

```

C      6.   -1.27572383   0.00000000   0.00000000   12.00000000
C      6.    1.27572383   0.00000000   0.00000000   12.00000000
H      1.   -2.34867651   0.00000000  -1.75798067    1.00782504
H      1.   -2.34867651   0.00000000   1.75798067    1.00782504
H      1.    2.34867651   0.00000000  -1.75798067    1.00782504
H      1.    2.34867651   0.00000000   1.75798067    1.00782504

```

## 12.1.3 Velocity

The velocity input is (free format, au):

*Name: veloc*

```

vx_1  vy_1  vz_1
vx_2  vy_2  vz_2
:
vx_nat vy_nat vz_nat

```

Example:

```

3.226196727134333E-004 -7.803939823701649E-004 3.501212063660452E-004
8.831202616257374E-005 1.103339279899924E-003 -7.468292758584672E-004
-1.994896289256706E-004 4.345679802152278E-004 -6.123248174920957E-004

```

NEWTON-X: Newtonian dynamics close to the crossing seam

```
-3.735660908785368E-003  2.225120145326573E-003  1.414904777249870E-003  
-2.272632671568894E-003 -2.177375422081543E-003  3.098061686476041E-003  
-6.469986389583130E-004  1.402416374342127E-004  2.362055042392439E-003
```

### 12.1.4 Freezing atoms

Cartesian coordinates of specific atoms can be kept frozen along the dynamics. For that, a list of atoms should be given in a file called `freeze.inp`. This file should be given together with the other input files. The atoms in the list are identified by its positions in the `geom` file and the list is blank separated.

Example of `freeze.inp`:

```
1 3
```

(Cartesian coordinate of atoms one and three in `geom` file will not change during the dynamics.)

If a file `freeze.inp` is present in the input, the atoms defined there will also not be affected by the thermostat.

The velocity of the atoms to be kept frozen should be set to zero in the `veloc` file (section 12.1.3). If the velocity of the frozen atoms is not initially set to zero, the program will stop with an error message.

Be aware that this algorithm may introduce spurious rotations and translation in the molecule depending on how the initial velocities of the remaining atoms were generated.

### 12.1.5 Specific input for quantum-chemistry electronic-structure calculations

Please, refer to the documentation of each particular program to see details on their inputs. Here, we present a brief summary of the main option that should be selected for some frequent jobs.

#### 12.1.5.1 Which electronic method to use

Nonadiabatic dynamics will always demand a compromise between quality and computational costs. In principle, the description of state crossing regions should be described by a multireference method like MRCI. However, often this is not affordable. In many cases (but not always), single references cases like TDDFT or ADC(2) have proven to give an adequate description of the nonadiabatic problem. You may keep the following thumb rules in mind when deciding which method to use in the simulation:

- Crossings between excited states at TDDFT and ADC(2) levels are in general well described if the ground state DFT is still single reference.
- Crossings between excited states will cause convergence problems at CC2 level.
- The description of crossings between the first excited state and the ground state is not reliable at TDDFT, CC2 or ADC(2) levels.

#### 12.1.5.2 Using analytical models

Analytical models for potential energies, gradients and non-adiabatic coupling vectors can be used in the molecular dynamics.

The analytical model is supposed to be a program in any language that reads molecular geometry from `geom` and velocities (if needed) from `veloc`, and writes the potential energies to `epot`, gradients to `grad`, and non-adiabatic coupling vectors to `nad_vectors`. The format of each one of these files is described in the section 16.3. If the analytical model requires additional files besides the executable file (for example, parameter inputs) they must be put in `JOB_NAD` directory. During the dynamics, the content of this directory is copied to the same location as the other NEWTON-X input files.

The specific location of the analytical model can be set via `analyt.par`. The options are given below.



Name: *analyt.par*

Parameter	Default	Description
ANMOD	[analytical.model]	File name of the executable containing the analytical model. The default is the two-dimension conical-intersection analytical model proposed by Ferretti et al. <sup>17</sup> .
PATH	[\$NX]	Absolute path to ANMOD file.

Example of *analyt.par* file:

```
anmod = my_model.x
path = /home/model/
```

The parameters of the two-dimension conical intersection model (default option) can be set via file *con\_int.dat* in *JOB\_NAD* directory. The options are given below.

Name: *JOB\_NAD/con\_int.dat*  
*namelist DAT*

Parameter	Default	Description
$\alpha$	= [3.0]	See Ref. <sup>17</sup> .
$\beta$	= [1.5]	
$K_x$	= [0.02]	
$K_y$	= [0.10]	
$\Delta$	= [0.01]	
$X_1$	= [4.0]	
$X_2$	= [3.0]	
$X_3$	= [3.0]	
$\gamma$	= [0.04]	

Example of *con\_int.dat*

```
&DAT
  alpha=3.0
  beta=1.5
  kx=0.02
  ky=0.10,
  delta=0.01
  x1=4.0
  x2=3.0
  x3=3.0
  gamma=0.04
&END
```

The atomic masses must be set directly in the geom file.

### 12.1.5.3 COLUMBUS

#### COLUMBUS 5.9

Adiabatic dynamics: JOB\_AD directory

Prepare a set of input files for a Columbus job (geometry optimization, one iteration, NROOT option). For MCSCF jobs, prepare input for CI gradient, but set "Maximum excitation level" to 0 in the CIDRT input. It is also possible to use MCSCF gradients, but in this case only single state dynamics are allowed (NSTAT = 1).

NEWTON-X: Newtonian dynamics close to the crossing seam

Non-adiabatic dynamics using non-adiabatic coupling vectors: JOB\_NAD directory

Prepare input files for a single-point non-adiabatic coupling calculation.

Non-Adiabatic dynamics using time-derivative couplings: JOB\_AD directory

Prepare a set of input files for a Columbus job (geometry optimization, one iteration, NROOT option). For MCSCF jobs, prepare input for CI gradient, but set "Maximum excitation level" to 0 in the CIDRT input.

**COLUMBUS 7.0**

Adiabatic and non-adiabatic (using coupling vectors or time-derivative couplings) dynamics are directly available at the SA-MCSCF and MR-CI levels. All inputs can be performed using the "non-adiabatic coupling" option in COLUMBUS and only the necessary terms are computed for each case. Optionally, adiabatic and time-derivative MR-CI dynamics may still be performed through "geometry optimization". Please set version=7.0 when using COLUMBUS 7.0.

Parameters to control COLUMBUS jobs

Name: columbus.par

Parameter	Default	Description
VERSION	= [5.9]	COLUMBUS version
MEM	= [200]	Core memory for COLUMBUS (Mwords). The same as MEM in control.dyn, but with priority over it. (1 GB = 134 Mwords). For COLUMBUS 7, MEM is automatically converted into GB.
CIRESTART	= [0]	0: do not use previous CI vector 1: use previous CI vector
MOCOEF	= [4]	0: use the same mocoef file in all time steps 1: use the mocoef file from the previous time step k: Lagrangean extrapolation of molecular orbitals at order k-1 ( $k \leq 11$ ) (see Section 12.1.10)
PRT_MO	= [20]	Save mocoef file to DEBUG file every PRT_MO timesteps. (Only if mocoef = 1.)
IVMODE	= [8]	Initial CI-vector generation mode. See COLUMBUS docs (ciudg program) for a list of options. CIRESTART keyword has priority over IVMODE keyword.
CITOL	= ["1E-4"]	Tolerance for MR-CI calculations
REDUCE_TOL	= [1]	0: keep rtolci and rtolbk in ciudgin as CITOL for all states. 1: use CITOL only to NSTATDYN. For all other states use 1E-3. 2: use CITOL only to NSTATDYN. Use 1E-3 for states coupled to NSTATDYN according to transmomin Columbus file. For all other states use 1E-1. Note that energies computed with 1E-1 are not reliable estimates. Although they are written to the output files, they do not have significance.
MC_CONV	= [0]	If mscf calculation do not converge: 0: warn and continue 1: kill trajectory
CI_CONV	= [0]	If CI calculation do not converge: 0: warn and continue 1: kill trajectory
QUAD_CONV	= [60]	Set value of NCOUPL in mscfn
MULL_POP	= [0]	1: print out the Mulliken populations from the COLUMBUS calculation to nx.log

NEWTON-X: Newtonian dynamics close to the crossing seam

#### 12.1.5.4 TURBOMOLE

##### Adiabatic dynamics: JOB\_AD directory

Prepare a set of input files for a TURBOMOLE job at TD-DFT or CC2 level without symmetry and copy them to JOB\_AD directory.

The initial geometry (TURBOMOLE coord file), the number of states (TURBOMOLE control file), and the state for which the gradient should be computed (TURBOMOLE control file) are automatically set by NEWTON-X during the program execution, overwriting the options in JOB\_AD.

For TD-DFT dynamics, TURBOMOLE control file is automatically changed to have:

```
$soes all NSTAT-1
$exopt NSTATDYN-1
```

For RI-CC2 or ADC(2) dynamics, TURBOMOLE control file is automatically changed to have:

```
$excitations
  irrep=a nexc=NSTAT-1
  exgrad states=(a NSTATDYN-1)
```

If TURBOMOLE auxbasis file is provided in JOB\_AD, NEWTON-X assumes that the resolution-of-identity (RI) should be used. For CC2 or ADC(2) dynamics, auxbasis must be always provided and the `ricc2` program is invoked every time step. For DFT dynamics, the presence of auxbasis file is optional. If it is provided, `ridft` program is invoked, otherwise `dscf` program is invoked.

The NEWTON-X input for CC2 and ADC(2) dynamics are exactly the same. The only difference is in the TURBOMOLE control file within JOB\_AD directory, which should be adequate to the desired method. During the execution, NEWTON-X search for ADC(2) keyword in the TURBOMOLE control file. If it is not found, the CC2 ground state energy is used. If it is found, MP2 ground state energy is used.

If you want to perform only adiabatic dynamics, set THRES = 0 (default in control.dyn file) to avoid that NEWTON-X starts the non-adiabatic dynamics calculations.

The TURBOMOLE mos file is not updated during the dynamics.

##### Non-Adiabatic dynamics: JOB\_AD directory

Prepare the TURBOMOLE input files as explained in the previous item. The nonadiabatic dynamics options are described in section 12.1.7.

Set THRES = 100 in control.dyn file.

##### Parallel Turbomole (SMP)

If you have the SMP parallelized version of Turbomole executables available (`dscf_smp`, `grad_smp`, `egrad_smp` and `ricc2_smp`) you can use them by specifying `parallel = <ncores>` in the file `turbomole.par` with `<ncores>` being the number of cores to use. The environment variable `$OMP_NUM_THREADS` is set automatically according to this number and needs not to be set by the user.

##### Parameters to control TURBOMOLE jobs

NEWTON-X: Newtonian dynamics close to the crossing seam

Name: *turbomole.par*

Parameter	Default	Description
PARALLEL	= [1]	Number of cores to use for parallel Turbomole (smp, no mpi!)

#### Inputs for old versions of NEWTON-X

For NEWTON-X versions prior the 1.0.8, additional input files should be provided for TD-DFT simulations if NSTATDYN = 1 and NSTAT > NSTATDYN. In this case, the following files are required in JOB\_AD directory:

control.opt - TURBOMOLE control file for ground state gradient (GRAD) calculation.  
 control.sp - TURBOMOLE control file for excited state single point (ESCF) calculation.  
 For excited state gradient calculations (NSTATDYN > 1), no additional files are needed.

#### **12.1.5.5 DFTB**

Adiabatic dynamics in the ground and excited states can be performed using the time-dependent density functional theory tight binding (TD-DFTB) method. Information about the program can be obtained at [www.dftb.org](http://www.dftb.org).

#### Input for ground state dynamics: JOB\_AD directory

Prepare DFTB input for conjugated-gradient ground-state calculation (code 4 in the DFTB program). The DFTB input and geometry information must be called dftb.in and in.gen, respectively.

#### Input for excited-state dynamics: JOB\_AD directory

Prepare TD-DFTB input for conjugated-gradient excited-state calculation (code 4 in the DFTB program). The DFTB input and geometry information must be called dftb.in and in.gen, respectively.

#### Parameters to control the DFTB jobs

Name: *dftb.par*

Parameter	Default	Description
DFTB_EXEC	[dftb]	Name of the DFTB executable file. \$DFTB variable must be defined in the system. NEWTON-X will run \$DFTB/<dftb_exec>.
OTHER_STATE	[1]	It is possible to perform dynamics on one surface, and at the same time to keep track of the properties of other states (energies and oscillator strengths). In this case, however, DFTB must run twice per time step, which makes the calculation more expensive. 0 - Do not monitor other states 1 - Monitor other states
MULT	[S]	Multiplicity of the states. Only states with the same multiplicity are allowed. S - singlet T - triplet

#### **12.1.5.6 GAUSSIAN**

##### **CASSCF level**

Surface-hopping non-adiabatic dynamics between the ground and the first excited state can be performed at CASSCF level with GAUSSIAN 03.

#### Content of JOB\_NAD directory

NEWTON-X: Newtonian dynamics close to the crossing seam

Prepare input files for a single point conical intersection calculation at CASSCF level. The input file must be called gaussian.com and the geometry must be given in Cartesian coordinates. The check point file named gaussian.chk containing the initial molecular orbitals must be provided as well. A suitable example of gaussian.com content is:

```
%chk=gaussian
%mem=20000000
#P OPT=(Conical,MaxCycle=1) CAS(4,3) IOp(5/7=200) Guess=read 3-21G Nosymm

methaniminium

1 1
N 0.000000 0.000000 0.637342
C 0.000000 0.000000 -0.703614
H -0.648747 0.572349 1.177883
H 0.648745 -0.572350 1.177883
H 0.618844 0.701080 -1.278050
H -0.618844 -0.701080 -1.278049
```

Note that OPT=(Conical,MaxCycle=1), Nosymm and Guess=read keywords and options should be given exactly as in this example. IOp(5/7=MaxIt) controls the maximum number of iterations in the CASSCF calculation.

### Parameters to control GAUSSIAN jobs

Name: g03.par

Parameter	Default	Description
MOCOEF	= [1]	0: use the initial molecular orbitals in all time steps 1: use the check point file from the previous time step as the source of molecular orbitals
PRT_MO	= [20]	Save gaussian.chk file to DEBUG file every PRT_MO timesteps. (Only if mocoef = 1.)

NEWTON-X executes GAUSSIAN 03 by invoking the command:

```
. $g03root/g03/bsd/g03.profile;$g03root/g03/g03 gaussian.com
```

If this path is not adequate for your system, you can change it in \$NX/run-g03.pl program.

### **TDDFT level**

Surface-hopping non-adiabatic dynamics with an arbitrary number of states can be performed at TDDFT level with GAUSSIAN 09.

### Content of the JOB\_AD directory

Prepare input files for a single point calculation at TDDFT level. The input file must be called gaussian.com and an optional check point file named gaussian.chk.

For nonadiabatic dynamics, NEWTON-X calls rwdump program of Gaussian. **\$g09root/g09/bsd/g09.login should be sourced either in the user profile or in the submission script before running the job.**

To run a calculation in parallel, the number of processors has to be specified with the keyword %nproc in the GAUSSIAN input. It has also to be specified in the submission script.

In file *gaussian.com*, the DFT functional, basis set and TD keyword with the proper options should be specified. **The user should not use the keywords `Guess=Read` and `TD(Read)`** in the input file. These options must be defined in *g09.par* file (see below). NEWTON-X adds `NoSymm` keyword automatically to avoid any problems due to differences between input and standard orientations.

The basis set may be directly given in the route or defined through `GEN` keyword. **For nonadiabatic dynamics, the basis set must be also provided in an additional file called *basis*.** Different ways of defining the basis sets are illustrated in the next examples.

- Example 1: The simplest case. The same basis set is used for all atoms and it is defined in the GAUSSIAN library.

Example of content of *gaussian.com* file:

```
%chk=gaussian
%rwf=gaussian
%mem=200mw
#TD(Nstates=3,Root=2) 3-21G BHandHLYP nosymm

test bs1

0 1
Si      -0.010544      0.010835      0.808752
C        0.015819     -0.050721     -0.931069
H         0.009875      1.149573      1.543340
H         0.172225     -1.176299      1.509994
H         0.071356      1.042598     -1.452426
H        -0.149187     -0.712162     -1.664254
```

Example of *basis* file:

```
3-21G
```

In this example, there is a redundancy in the basis set definition, which appears in *gaussian.com* and in *basis*. You should carefully check whether both files contain the same definitions.

- Example 2: Basis set is not defined in the GAUSSIAN library. Then you have to use the `GEN` keyword in *gaussian.com* and provide the basis set. In addition you have to copy the basis set to the *basis* file.

Example of GAUSSIAN input:

```
%chk=gaussian
%rwf=gaussian
%mem=200mw
#TD(Nstates=3,Root=2) GEN BHandHLYP nosymm

test bs1

0 1
Si      -0.010544      0.010835      0.808752
C        0.015819     -0.050721     -0.931069
H         0.009875      1.149573      1.543340
H         0.172225     -1.176299      1.509994
H         0.071356      1.042598     -1.452426
H        -0.149187     -0.712162     -1.664254

H 0
S 2 1.00
    5.4471780      0.1562850
    0.8245470      0.9046910
S 1 1.00
    0.1831920      1.0000000
```

NEWTON-X: Newtonian dynamics close to the crossing seam

```

****
C      0
S      3      1.00
      172.2560000          0.0617669
      25.9109000          0.3587940
      5.5333500          0.7007130
SP     2      1.00
      3.6649800          -0.3958970          0.2364600
      0.7705450          1.2158400          0.8606190
SP     1      1.00
      0.1958570          1.0000000          1.0000000
****
Si     0
S      3      1.00
      910.6550000          0.0660823
      137.3360000          0.3862290
      29.7601000          0.6723800
SP     3      1.00
      36.6716000          -0.1045110          0.1133550
      8.3172900          0.1074100          0.4575780
      2.2164500          0.9514460          0.6074270
SP     2      1.00
      1.0791300          -0.3761080          0.0671030
      0.3024220          1.2516500          0.9568830
SP     1      1.00
      0.0933392          1.0000000          1.0000000
****

```

Content of *basis* file:

```

H      0
S      2      1.00
      5.4471780          0.1562850
      0.8245470          0.9046910
S      1      1.00
      0.1831920          1.0000000
****
C      0
S      3      1.00
      172.2560000          0.0617669
      25.9109000          0.3587940
      5.5333500          0.7007130
SP     2      1.00
      3.6649800          -0.3958970          0.2364600
      0.7705450          1.2158400          0.8606190
SP     1      1.00
      0.1958570          1.0000000          1.0000000
****
Si     0
S      3      1.00
      910.6550000          0.0660823
      137.3360000          0.3862290
      29.7601000          0.6723800
SP     3      1.00
      36.6716000          -0.1045110          0.1133550
      8.3172900          0.1074100          0.4575780
      2.2164500          0.9514460          0.6074270
SP     2      1.00
      1.0791300          -0.3761080          0.0671030
      0.3024220          1.2516500          0.9568830
SP     1      1.00
      0.0933392          1.0000000          1.0000000

```

In this case, the basis set defined in *basis* will be used in the dynamics even if *gaussian.com* has a different definition.

- Example 3: Different basis sets for different atoms. GEN keyword is also needed in this case.

### Example of GAUSSIAN input:

```
%chk=gaussian
%rwf=gaussian
%mem=200mw
#TD(Nstates=3,Root=2) GEN BHandHLYP nosymm

test bs1

0 1
  Si      -0.010544      0.010835      0.808752
  C        0.015819     -0.050721     -0.931069
  H         0.009875      1.149573      1.543340
  H         0.172225     -1.176299      1.509994
  H         0.071356      1.042598     -1.452426
  H        -0.149187     -0.712162     -1.664254
```

### Content of *basis* file:

```
1 0
6-31G(d)
****
2 0
6-31G(d)
****
3 0
3-21G
****
4 0
3-21G
****
5 0
6-31G
****
6 0
6-31G
```

In this case, the basis set defined in *basis* will be used in the dynamics even if *gaussian.com* has a different definition.

### Parameters to control GAUSSIAN jobs

Name: *g09.par*

Parameter	Default	Description
MOCOEUF	= [0]	0: compute the initial guess at every time step. 1: use the check point file from the previous time step as the source of molecular orbitals. 2: use the initial molecular orbitals in all time steps.
TD_ST	= [0]	0: compute the excited states without previous reference. 1: read the states from the checkpoint file, the checkpoint file is controlled by MOCOEUF (1 or 2). We don't recommend using this option in non-adiabatic dynamics calculations.
PRT_MO	= [20]	Save gaussian.chk file to DEBUG file every PRT_MO timesteps.
LD_THR	= [14]	N: Controls the linear dependency threshold ( $10^{-N}$ ) in G09 calculations.

NEWTON-X executes GAUSSIAN 09 by invoking the command:

```
. $g09root/g09/bsd/g09.profile;$g09root/g09/g09 gaussian.com
```

If this path is not adequate for your system, you can change it in \$NX/run-g09.pl program.

NEWTON-X: Newtonian dynamics close to the crossing seam



### 12.1.5.7 TINKER

Adiabatic ground-state dynamics with forcefield-gradients and –energies can be performed with TINKER.

#### Content of the JOB\_AD-directory

An TINKER-XYZ file called tinkin.xyz and the corresponding keyword-file tinkin.key have to be present. The file tinkin.key has to state at least the name of the parameter file ('parameters <filename>') and 'digits 12'.

The command executed by NEWTON-X is

```
testgrad tinkin.xyz -k tinkin.key y n n
```

If the TINKER-executables are not in the \$PATH, testgrad.x can also be provided as binary in the JOB\_AD directory.

### 12.1.5.8 DFTB+

Adiabatic dynamics in the ground can be performed using the density functional theory tight binding (DFTB) method.

#### Input for ground state dynamics: JOB\_AD directory

Prepare DFTB+ input for conjugated-gradient ground-state calculation (code 4 in the DFTB program). The DFTB+ input and geometry information must be called dftb.in and in.gen, respectively.

#### Parameters to control the DFTB+ jobs

Name: dftb+.par

Parameter	Default	Description
DFTBP_EXEC	[dftb+]	Name of the DFTB+ executable file. \$DFTBPLUS variable must be defined in the system. NEWTON-X will run \$DFTBPLUS/<dftbp_exec>.

### 12.1.5.9 GAMESS

#### Adiabatic dynamics: JOB\_AD directory

Prepare a gradient input file for a GAMESS job at some adiabatic level (e.g. CCSD(T) or MP2) without symmetry and copy the file to the JOB\_AD directory. The input file should be labeled either gamessinput\_original.inp (without the \$DATA group) or nx\_gamess.inp (with the \$DATA group). Either way, starting coordinates for the run always get taken from the NEWTON-X geom file.

The initial geometry, the number of states, and the state for which the gradient should be computed are automatically set by NEWTON-X during the program execution. Please ensure the iroot variable (in \$det or similar group) in nx\_gamess.inp is the same as nstatdyn in control.dyn file.

If you want to perform only adiabatic dynamics, set THRES = 0 in control.dyn file to avoid non-adiabatic dynamics calculations.

Analytic gradients are available for state-averaged runs and are activated with wtsok=.t. in the \$det group ( and no NUMGRD=.t. in \$contrl).

GAMESS input examples exist in \$NX/test-nx.

NEWTON-X: Newtonian dynamics close to the crossing seam

Non-adiabatic dynamics using non-adiabatic coupling vectors: JOB\_NAD directory

NA couplings are available at the SA-MCSCF level.

Specific NA coupling selection schemes are also available through the jiri.inp file with napick=.t. in \$cpconv. All NA couplings can be calculated with napick=.f. in \$cpconv along with appropriate input to jiri.inp.

Prepare input files for a single-point non-adiabatic coupling calculation.

Prepare either:

1)

gamessinput\_original.inp: runtyp=nacme without \$DATA and \$VEC

gamessinput\_original.vec: \$VEC group

2)

nx\_gamess.inp: runtyp=nacme with \$DATA and \$VEC

Parameters to control GAMESS jobs

Name: gamess.par

Parameter	Default	Description
VERNO	= [00]	Version number of GAMESS executable.
NCPUS	= 1	Number of computer processes to be run.
RUN_GAMESS	= [rungms]	Name of GAMESS execution script.
SCR	= [0]	GAMESS scratch directory: 0: Use default GAMESS options as defined in RUN_GAMESS script. 1: Create SCR directory inside TEMP directory of NEWTON-X.
MOCOEF	= [1]	0: use the initial molecular orbitals in all time steps 1: use the mocoef file from the previous time step k: Lagrangean extrapolation of molecular orbitals at order k-1 (k ≤ 11) (see Section 12.1.10)

**12.1.5.10 Hybrid Gradients (QM/MM)**

The hybrid gradients module allows adiabatic and non-adiabatic dynamics with combinations of programs (COLUMBUS, TINKER, TURBOMOLE and Analytical Model). The current implementation is described in Ref.<sup>34</sup>.

General explanations of hybrid calculations

Energies and gradients for subsets of atoms are treated with different programs and the partial gradients are then joint into a resulting total (hence 'hybrid') energy and gradient. For that purpose the set of atoms of the whole system is split in disjoint regions. These regions need not follow physical reasoning (they often will, but they can also pick e.g. single atoms out of molecules), but are logical entities for the definition of the single partial calculations.

The whole calculation is split into jobs. Each of these jobs can treat one or more regions of atoms and the partial result are multiplied by a user-defined factor before being added to the total. One region can be treated by multiple jobs and to care about 'double counting' of atoms is left to the user completely.

For COLUMBUS and TURBOMOLE there is the possibility to include regions only as point charges and not as atoms with basis-set. A hybrid setup (which is done in the JOB\_AD or JOB\_NAD

NEWTON-X: Newtonian dynamics close to the crossing seam

directory) consists of some general information about the hybrid setup ('control' parameters), the definition of the atoms, some of their properties and membership to the regions and the definition of the partial jobs, which regions they are concerned with and how they shall be put together to the overall result. The main input file is named `hybrid_gradients.inp`.

The nonadiabatic couplings, oscillator strengths and other nonadditive properties are given by only one job. It is possible to restrict the hybrid NAD-vectors to some regions. **The NAD-vector components on all atoms not belonging to these regions are replaced by Zeros in this case.** Also for back hoppings only the kinetic energy of these nad-regions is regarded as available energy.

For the treatment of bonded interaction between two regions link atoms can be inserted in bond connecting them. The gradient- and nonadiabatic coupling vector elements of these link atoms will be distributed to the two atoms between which it has been inserted. To avoid over-polarization effects, the point charges near the link atom can be set to zero and, so as to retain the overall charge, re-distributed to a set of other atoms. The job, where the link atom is inserted is (suggestively) called 'QM\_JOB' regardless of the method really used in this job. Similarly is the naming of 'QM\_ATOM' and 'MM\_ATOM'.

#### Format of the hybrid\_gradients.inp file

Each section begins with `$<name>` and ends with `$end`, where `<name>` can be 'control', 'job' or 'atoms'. Parameters for sections '\$control' and '\$job' are set in key=value pairs. All key=value pairs have to be separated with blanks.

Name: `hybrid_gradient.inp`

Section	Keyword	Description
<b>\$control</b>	NATOMS	number of atoms (optional)
	PROPERTIES	ID of the job, that gives the properties (oscillator strength, nonadiabatic couplings, optional)
	NADREGIONS	array of regions to which the hybrid nad-vector shall be restricted
<b>\$job</b>	ID	unique numeric (integer) identification for the job
	PROGRAM	name of the program to be used for this job ( <i>columbus, tinker, turbomole</i> )
	REGIONS	array of regions, that shall be treated by this job ( <i>comma seperated</i> )
	FACTOR	pre-factor with that the result of this job enters the total result (energy, gradients)
	POINTCHARGES	array of regions, that shall be treated as point-charges ( <i>for columbus and turbomole, comma seperated, optional</i> )
<b>\$atoms</b>	(For each atom give in one single line and in this order: <name>	Label for an atom (compulsory). Can be anything without blanks, comma and similar. The type of atom is recognized from the nuclear charge.
	<nuclear charge>	Nuclear charge.
	<x,y,z coords>	Cartesian coordinate (Bohr).
	<pointcharge>	Effective point charge for this atom. If this atom is not treated as point charge this can be left to 0.0000.
	<mass>	Atomic mass for this atom (amu).

	<region>	The region in which this atom is.
link	QM_JOB	ID of the job where the link atom shall be inserted.
	QM_ATOM	number (=position in input) of the atom connected to the link atom at the QM-side.
	MM_ATOM	number (=position in input) of the atom connected to the link atom at the MM side.
	RATIO	ratio of distances QMat→LINKat/QMat→MMat.
	ZERO	comma seperated list of atoms for which the point charge shall be set to 0.000 in the QM_JOB.
	SCATTER	comma seperated list of atoms to which the accumulated charge of the ZERO-atoms shall be distributed.

The section \$atoms is an extension of the NEWTON-X geom. with the two additional columns, <pointcharge> and <region>.

Each atom has to be in **exactly one** region.

All values in section \$atoms have to be separated with blanks.

Every section has to be ended with \$end. No section (except '\$jobs') is allowed to appear twice.

There are no further formatting rules to the file.

Example of hybrid\_gradient.inp file:

```
$job ID = 1 regions = 1,2 program = columbus pointcharges = 2 factor = 1 $end
$job ID = 2 regions = 1,2 program = tinker factor = 1 $end
$job ID = 3 regions = 1 program = tinker factor = -1 $end
$control properties = 1 nadregions = 1 natoms = 9 $end
$atoms
  C 6.0 -0.58698100 -0.10086826 0.12744020 12.00000000 0.0000 1
  O 8.0 1.68492145 0.00195156 0.55748655 15.99491464 0.0000 1
  N 7.0 -1.73272850 -2.27661770 -0.35180374 14.00307401 0.0000 1
  H 1.0 -1.76941972 1.58570542 0.11857328 1.00782504 0.0000 1
  H 1.0 -3.64014444 -2.31932440 -0.70744017 1.00782504 0.0000 1
  H 1.0 -0.69532029 -3.86742067 -0.35771869 1.00782504 0.0000 1
  OW 8.0 -6.90020375 -2.40598424 -1.31691242 15.99491464 -0.8340 2
  HW 1.0 -7.75621060 -2.20340156 -2.90625351 1.00782504 0.4170 2
  HW 1.0 -8.26652850 -2.54571452 -0.12804111 1.00782504 0.4170 2
$end
```

This example is explained in details in the NEWTON-X tutorial.

Example including link atoms (some lines not displayed):

```
# 2-butene with cyclohexane rings attached to the sides
# Z2-scheme is used for point charges, i.e. first and second
# neighbour to the link atom have zero charges and the
# their charge is distributed to their next-neighbours on the MM side

$job ID=1 regions=1,2,3 program=turbomole turbo_method=td-dft pointcharges=2,3 factor=1 $end
$job ID = 2 regions = 1,2,3 program = tinker factor = 1 $end
$job ID = 3 regions = 1 program = tinker factor = -1 $end
$link qm_job=1 qm_atom=1 mm_atom=11 ratio=0.713 zero=11,12,13,14 scatter=15,16,17 $end
$link qm_job=1 qm_atom=4 mm_atom=34 ratio=0.713 zero=34,35,36,37 scatter=38,39,40 $end
$control properties = 1 nadregions=1 natoms = 56 $end

$atoms
  C 6.0 1.44087838 0.85091533 -3.47469613 12.00000000 0.0000 1 #atom 1
                                     # linked to atom 11
  C 6.0 1.23567680 -0.48733580 -6.00452731 12.00000000 0.0000 1
  C 6.0 2.80448775 -2.25355699 -6.92713507 12.00000000 0.0000 1
  C 6.0 5.12903987 -3.34336584 -5.64481794 12.00000000 0.0000 1 #atom 4
                                     # linked to atom 34
  H 1.0 1.17513186 2.91664300 -3.78035178 1.00782504 0.0000 1
```

```

H 1.0 3.36724819 0.62483228 -2.66873550 1.00782504 0.0000 1
H 1.0 -0.40125500 0.07262218 -7.17351567 1.00782504 0.0000 1
H 1.0 2.39776578 -3.02448967 -8.82452971 1.00782504 0.0000 1
H 1.0 4.98819669 -5.44423296 -5.69313068 1.00782504 0.0000 1
H 1.0 5.18046121 -2.82088113 -3.61093594 1.00782504 0.0000 1
CT 6.0 -0.52739233 -0.13060275 -1.57905516 12.00000000 -0.1200 2 #atom 11
HC 1.0 -0.25927609 -2.20323737 -1.33315643 1.00782504 0.0600 2
HC 1.0 -2.45999068 0.14865342 -2.36823880 1.00782504 0.0600 2
CT 6.0 -0.33457412 1.20138205 0.99721226 12.00000000 -0.1200 2
HC 1.0 1.56440222 0.83732631 1.83045298 1.00782504 0.0600 2
HC 1.0 -0.45614398 3.28079700 0.68304340 1.00782504 0.0600 2
: : : : : : : : : : : : : : : : : : : : : : : : : :
CT 6.0 -2.23690095 2.08870485 5.28849685 12.00000000 -0.1200 2
HC 1.0 -0.32070164 1.91506358 6.14394752 1.00782504 0.0600 2
HC 1.0 -2.48154867 4.11306830 4.76133507 1.00782504 0.0600 2
CT 6.0 7.60087913 -2.51304285 -6.92702736 12.00000000 -0.1200 3 #atom34
HC 1.0 9.21362621 -3.51540650 -6.01449183 1.00782504 0.0600 3
HC 1.0 7.56925645 -3.13134234 -8.93790950 1.00782504 0.0600 3
CT 6.0 8.05428868 0.35505686 -6.77227957 12.00000000 -0.1200 3
HC 1.0 7.92399962 0.94155983 -4.75349648 1.00782504 0.0600 3
HC 1.0 6.51305828 1.36826754 -7.78612331 1.00782504 0.0600 3
CT 6.0 10.63311478 1.21580066 -7.83079266 12.00000000 -0.1200 3
HC 1.0 12.15373961 0.13167423 -6.84744154 1.00782504 0.0600 3
: : : : : : : : : : : : : : : : : : : : : : : : : :
$end

```

### Generating and updating the remaining input files

After having written `hybrid_gradient.inp` file, the complete set of input files can be created by run

```
$NX/hybrid_read_onefile.pl
```

Subdirectories for the different jobs will be created as well. Appropriate geom files for each job are provided in these subdirectories. Set up the third-party single jobs inputs in the subdirectories in the same way as for normal NEWTON-X dynamics. If you chose to include some atoms as point-charges to a COLUMBUS- or TURBOMOLE-job you have to set the jobs up appropriately. Refer to the documentations of the programs for information on how to do that (`hybrid_read_onefile.pl` will provide some useful files containing most of things, that have to be done extra to a normal setup).

In the subdirectories for partial hybrid jobs to be computed with COLUMBUS including point-charges two files 'potential.xyz' and 'elpotin' are provided. Do not delete these files!

In the subdirectories for partial hybrid jobs to be computed with TURBOMOLE including point-charges two files 'pointcharges' and 'control-additions' are provided. Do not delete the pointcharges-file. The contents of the file 'control-additions' have to be inserted in the control-file at an appropriate position after setting up the TURBOMOLE calculation.

If any change is done to `hybrid_gradient.inp`, `hybrid_read_onefile.pl` should be run again to update the remaining files.

#### 12.1.6 Thermostat control

Thermal-equilibration may be obtained during the dynamics by using a thermostat<sup>33</sup>. In the current NEWTON-X version, the Andersen thermostat is available<sup>32</sup>. The Andersen-Lowe thermostat<sup>46</sup> is being implemented and it will available soon. The parameters of the thermostat may be adjusted by using `nxinp` tool, in the input section "Set General Options". The options are given below.

Name: `therm.inp`  
namelist `therm`

Parameter	Default	Description
KTHERM	= [1]	Thermostat type: 0 - No thermostat. 1 - Andersen <sup>32</sup> . 2 - Andersen-Lowe <sup>46</sup> (in development).

NEWTON-X: Newtonian dynamics close to the crossing seam

---

TEMP	= [300]	Temperature (K).
KTS	= [1]	Turn the thermostat on at time step KTS. (KTS ≥ 1, integer)
LTS	= [-1]	Turn the thermostat off at time step LTS. (LTS > KTS, integer) -1 - Do not turn it off.
NSTHERM	= [1]	If the system is in the excited state: 0 - turn the thermostat off. 1 - apply the thermostat.
GAMMA	= [0.2]	Collision frequency (fs <sup>-1</sup> ).
RADIUS	= [10.0]	Collision radius (bohr). (Only in Andersen-Lowe thermostat.)
ISEED	= [1]	Random number seed for the thermostat. 0 - default seed value. 1 - generate random seed. > 1 - use this value (integer) as the seed.
LVP	= [1]	Print level. (In any case, relevant information is written to NEWTON-X log.) 1 - Do not print thermostat log file. 2 - Print minimum thermostat log file. 3 - Print debug-level thermostat log file.

---

Example of therm.inp file:

```
&therm
  ktherm = 1
  kts = 1
  lts = -1
  nstherm = 1
  temp = 300
  gamma = 0.6
  iseed = 0
  lvp = 3
&end
```

You can choose a subset of atoms that will not be affected by the thermostat. For that, a list of atoms should be given in a file called therm.freeze. This file should be given together with the other input files. The atoms in the list are identified by its positions in the geom file and the list is blank separated.

Example of therm.freeze:

```
1 3
```

(Atoms one and three in geom file will not be affected by the thermostat.)

If a file freeze.inp is present in the input, the atoms defined there will also not be affected by the thermostat (see section 12.1.4).

### 12.1.7 Non-adiabatic dynamics control

The non-adiabatic dynamics is controlled by two input files, sh.inp and jiri.inp.

SH is a stand-alone module for surface-hopping. It reads nuclear velocities and non-adiabatic couplings and gives as output the electronic wavefunction expansion coefficients on the adiabatic set of states. Switch from one adiabatic PES to another is ruled by the Tully's fewest switches algorithm<sup>13</sup>.

The parameters for non-adiabatic dynamics can be set via nxinp tool, at the input option “Set non-adiabatic dynamics”. The options are given below.

Name: *sh.inp*

*namelist shinp*

Parameter	Default	Description
VDOETH	= [0]	Dynamics with non-adiabatic coupling vectors or with time-derivative couplings (HST model <sup>18,19</sup> ). -1 - Local-diabatization method. <sup>20</sup> 0 - Compute non-adiabatic coupling vectors. 1 - Compute time-derivative couplings (COLUMBUS and TURBOMOLE TD-DFT). <sup>18</sup> The default is reset to 1 if PROG = 2.1.
SEED	= [1]	0 a default random number seed is used. 1 a randomized seed is used >1 random number seed
INTEGRATOR	= [5]	selects the integrator of the TDSE: 0 - a "home-made" integrator, see Ref. <sup>17</sup> . 1 - standard 4th order Runge-Kutta. 2 - Adams Moulton predictor-corrector, 5th order. 3 - Adams Moulton predictor-corrector, 6th order. 4 - Unitary propagator. 5 - Butcher, 5th order. <sup>47</sup> 6 - Unitary propagator for Local-Diabatization method. Good choices are usually 3, 4, and 5. For vdoeth = -1 (local diabatisation) integrator must be 6.
PHASE	= [1]	Integrate the phase along the trajectory (only for debug purposes): 0 - No. Phase is always zero. 1 - Yes.
NOHOP	= [0]	Force the hopping at the certain time step. 0 - The trans. prob. are computed and hopping is allowed (normal surface hopping). -1 - Hopping is not allowed at any time. n - Hopping is forced at (and only at) timestep n (n = positive integer).
FORCESURF	= [1]	Force the hopping to the surface FORCESURF (ground state = 1).
NRELAX	= [0]	number of cycles after a surface hopping, in which other hopping is forbidden.
MS	= [20]	number of subtime-steps for integration of the time dependent Schroedinger equations is dt/ms, where dt is the time-step of the classical trajectory defined in control.dyn file. In the ms-1 substeps between t and t+dt, the non-adiabatic coupling vector and the velocity vector is obtained by linear interpolation. The potential energy is interpolated with a cubic polynomial. 0 - Do not interpolate. 1 - Time-step will be divided by one. The results is the same as using ms = 0, but now the interpolation programs are called. For vdoeth = -1 (local diabatisation), ms must be 0.
GETPHASE	= [1]	0 use phase provided by the overlap of CI vectors. 1 use phase provided by the scalar product between h(t) and h(t-dt). Usually this is the best choice for multistate dynamics.
TULLY	= [1]	Fewest-switches algorithm: 0 - Tully <sup>13</sup> . 1 - Hammes-Schiffer and Tully <sup>19</sup> .
DECAY	= [0.1]	Apply the decoherence correction to the time-dependent coefficients $C_K$ of state $K$ <sup>26</sup> . The value of DECAY is assumed by the variable $\alpha$ in the equations: $C'_K = C_K \exp(-\Delta t / \tau_{KM}) \quad \forall K \neq M,$ $C'_M = C_M \left[ \frac{1 - \sum_{K \neq M}  C'_K ^2}{ C_M ^2} \right]^{1/2},$ $\tau_{KM} = \frac{\hbar}{ E_K - E_M } \left( 1 + \frac{\alpha}{E_{kin}} \right),$ where $M$ is the current state and $E_{kin}$ is the nuclear kinetic energy. $\alpha$ (DECAY) must be given

in atomic units (hartree). The recommended value is  $\alpha = 0.1$  hartree. -1 means normal calculation (without correction). After integrating the TDSE to obtain the coefficients  $\mathbf{C}$  and using the fewest-switches to determine the current state  $M$ , the equations above are used to obtain the coefficients  $\mathbf{C}'$ . The  $\mathbf{C}'$  are then used to continue the time evolution of the electronic wave function.

PROBMIN	= [0]	Do not hop if probability is smaller than PROBMIN.
MOM	= [1]	After a frustrated hopping: -1 - Invert momentum direction. 1 - Keep momentum direction.
ADJMOM	= [0]	After a hopping adjust momentum: -1 - along the momentum direction. 0 - along the non-adiabatic coupling vector $\mathbf{h}$ . 90 - along the gradient difference vector $\mathbf{g}$ . For any value larger or equal to 0, ADMOM is assumed to be an angle $\alpha$ in degrees, which defines the unitary vector $\hat{\mathbf{e}} = \sin(\alpha) \frac{\mathbf{g}}{\ \mathbf{g}\ } + \cos(\alpha) \frac{\mathbf{h}}{\ \mathbf{h}\ }$ The momentum is adjusted in the direction of $\hat{\mathbf{e}}$ . Note: Until NEWTON-X version 0.10a, the definition of ADMOM and the default value were different. Check old input files before running. The default is reset to -1 if VDOTH = 1.
POPDEV	= [0.05]	Kill trajectory if total adiabatic population deviate more than POPDEV from the unity.

Example of sh.inp file:

```
&shinp
  seed=0,
  integrator=4,
  nrelax=0,
  ms=10,
  getphase=1,
&end
```

The following keywords in jiri.inp file allow additional control over the non-adiabatic coupling calculations by restricting which states should be included in the calculations. In combination with THRES defined in control.dyn, it is possible to substantially reduce the computational costs by excluding couplings that should not contribute to the hopping probabilities<sup>18</sup>. The default (THRES = 0, KROSS = 1, CASCADE = 0, CURRENT = 1, NEVER\_STATE = 0, INCLUDE\_STATE = 0) implies that all possible coupling vectors between the current and the other states will be computed. When time derivative couplings are used (VDOTH = 1), the default of NEVER\_STATE = 1, meaning that non-adiabatic coupling with the ground state are not computed.

Internally, the restrictions implied by these keywords are independently applied to the complete list of  $N_C = N_{\text{stat}}(N_{\text{stat}}-1)/2$  possible coupling vectors and the result is a restricted list of couplings. This information is written every timestep to the file transmomin. Thus, the list of coupling vectors that should be computed is dynamically updated along the simulation. If a transmomin file is given with the input in JOB\_NAD it will be replaced by the new one already in step 0.

Name: jiri.inp

Namelist: jirinp

Parameter	Default	Description
KROSS	= [1]	0 - do not calculate non-adiabatic couplings between non-consecutive states. For example, between $S_0$ and $S_2$ . 1 - calculate non-adiabatic couplings between non-consecutive states.

NEWTON-X: Newtonian dynamics close to the crossing seam



---

CASCADE	= [0]	Compute non-adiabatic couplings only for states below the current state and state: 0 - above and below (e.g., in S <sub>2</sub> get S <sub>2</sub> -S <sub>1</sub> and S <sub>2</sub> -S <sub>3</sub> ) 1 - only below (e.g., in S <sub>2</sub> get S <sub>2</sub> -S <sub>1</sub> but not S <sub>2</sub> -S <sub>3</sub> )
CURRENT	= [1]	Compute non-adiabatic couplings only for pairs of states including the current state.
NEVER_STATE	= [0]	Array of states for which the non-adiabatic coupling vectors should never be computed. E.g., if never_state=1,2: don't compute couplings with S <sub>0</sub> or S <sub>1</sub> (1 – ground state). The array must be coma separated. The maximum number of states in the array is 10. The default (0) means do not exclude any state. The default is 1 when VDOTH = 1.
INCLUDE_PAIR	= [0]	Define the pairs of states for which the non-adiabatic coupling vector for the following pairs should be always computed, when the current state is one of the states in the pair. E.g., include_pair=1,2,1,3: when the molecule is in S <sub>0</sub> , always compute couplings S <sub>0</sub> -S <sub>1</sub> and S <sub>0</sub> -S <sub>2</sub> (1 – ground state). The array must be coma separated. The maximum number of pairs is 5.
E_CI	= [0.2]	Check the energy difference between every pair of states and report it when this difference drops below the e_ci threshold (in eV). The information is written to RESULTS/report.CI and it is useful to locate conical intersections.
If VDOTH = 1 in sh.inp:		
CI_CONS	= [1]	0 - Compute all determinant overlap terms. 1 - Neglect determinant overlap terms when the hank is too high or when the determinants are orthogonal.
CIO_OPTIONS	COLUMBUS: =["-t 1e-5 -e 2 -i"] TDDFT, TURBOMOLE: =["-s transmomin -a -t 5e-5 -e -1"]	Options to the cioverlap program (given in quotation marks). See detailed description in section 16.6.
CISC_OPTIONS	= [,-o -c"]	Options to the cis_casida program (given in quotation marks). See detailed description in section 16.6
NCORE	= [0]	Number of core orbitals that should be ignored in the CIS Casida procedure (using the -i option of the cis_casida executable).
NDISC	= [0]	Number of virtual orbitals that should be ignored in the CIS Casida procedure (using the -I option of the cis_casida executable).

---

Example of jiri.inp file:

```
&jirinp
  kross      = 0
  cascade    = 1
/&end
```

Do not forget "&jirinp" and "/&end".

There are eight possible combinations of the keywords KROSS, CASCADE and CURRENT. Each one corresponds to a different model. The models are called: two-state (TS), three-state (3S), horizontal coupling (HC), lower diagonal (LD), partial coupling (PC), neighbour coupling (NC), lower triangular, and complete coupling (CC)<sup>18</sup>. In the Table below, the keyword combination for each of these models is given, as well as the number N<sub>c</sub> of coupling vectors computed in each time step. An

NEWTON-X: Newtonian dynamics close to the crossing seam

example of which coupling vectors are actually computed in the case of dynamics with  $N_{\text{stat}} = 5$  states with the molecule instantaneously in  $N_{\text{statdyn}} = 3$  state is also shown in the Table. The bold line shows the default option.

kross	cascade	current	Model	$N_c$	Example: NSTAT=5, NSTATDYN=3										
					21	31	32	41	42	43	51	52	53	54	
0	1	1	TS	1			x								
0	0	1	3S	2			x			x					
1	1	1	HC	$N_{\text{statdyn}}-1$		x	x								
0	1	0	LD	$N_{\text{statdyn}}-1$	x		x								
<b>1</b>	<b>0</b>	<b>1</b>	<b>PC</b>	<b><math>N_{\text{stat}}-1</math></b>		<b>x</b>	<b>x</b>			<b>x</b>				<b>x</b>	
0	0	0	NC	$N_{\text{stat}}-1$	x		x			x					x
1	1	0	LT	$N_{\text{statdyn}}(N_{\text{statdyn}}-1)/2$	x	x	x								
1	0	0	CC	$N_{\text{stat}}(N_{\text{stat}}-1)/2$	x	x	x	x	x	x	x	x	x	x	x

In addition to these eight models, other combination can be built by using the keywords NEVER\_STATE, INCLUDE\_PAIR, and THRES (in control.dyn). For example, the following jiri.inp:

```
&jirinp
  kross      = 0
  cascade    = 1
  current    = 1
  include_pair = 1,2
/&end
```

will result in a modified TS model for which the coupling vector from state 1 (ground) to state 2 is computer whenever the molecule is in state 1.

### 12.1.8 Time-derivative couplings

Dynamics using time-derivative couplings to compute the non-adiabatic surface hopping probabilities can be performed with COLUMBUS (MCSCF, MRCI) and TURBOMOLE (TD-DFT). The method was proposed by Hammes-Schiffer and Tully<sup>19</sup> and the current implementation is described in Ref.<sup>18</sup>.

For TD-DFT this method should be used exclusively for surface hopping between excited states, because the linear-response TD-DFT cannot adequately describe the mutireference character of the electronic wavefunction near conical intersections between the ground and the first excited state.

### 12.1.9 Wave function coefficients

The real and imaginary parts of the coefficients of the time dependent Schroedinger equation are automatically updated to the file wfrun during the dynamics. The default is to start the dynamics with 1.0 at the initial state and 0.0 for all other states. It is possible to change this default by just including wf.inp as an additional input file, with the format:

```
Re(C_1) Im(C_1)
Re(C_2) Im(C_2)
...
Re(C_NSTATDYN) Im(C_NSTATDYN)
...
Re(C_NSTAT) Im(C_NSTAT)
```

For example, the default for a two-state dynamics ( $S_0$  and  $S_1$ ), starting in  $S_1$  is:

```
0.0 0.0
1.0 0.0
```

One possibility to start with 20% at  $S_2$  and 80% in  $S_1$  is:

```
0.894427191 0.0
0.447213596 0.0
```

NEWTON-X: Newtonian dynamics close to the crossing seam

### 12.1.10 Propagation of molecular orbitals

The keyword MOCOEF appearing in some of the <third-party>.par files controls how the molecular orbital coefficients are transferred from one step to the following. The original (time step  $n = 0$ ) coefficients may be used, or the coefficients from step  $n$  may be given as guess for step  $n+1$  (COLUMBUS, GAUSSIAN, AND GAMESS).

A set of  $k$  molecular orbital coefficients computed between steps  $n-k$  and  $n$  can also be extrapolated to create the guess for step  $n+1$  (COLUMBUS AND GAMESS). This is done by means of Lagrangean Extrapolation of Molecular Orbitals (LEMO) algorithm described in Ref. <sup>31</sup>. Briefly the molecular orbital coefficient guess ( $\mathbf{c}_{guess}$ ) for step  $n+1$  is given in terms of the previously converged coefficients ( $\mathbf{c}_{conv}$ ) by

$$c_{guess,i,j}^{n+1} = \sum_{l=0}^{k-1} L_{k-1,l} c_{conv,i,j}^{n-l} \quad (1 \leq i \leq n_{bas}, 1 \leq j \leq n_{bas}),$$

where  $\mathbf{L}$  is the set of Lagrangean coefficients for polynomial interpolation and  $n_{bas}$  is the number of basis functions.

The treatment of orbital rotations in the version of LEMO algorithm implemented in NEWTON-X differs from the one proposed in Ref. <sup>31</sup>. In NEWTON-X, the overlap factor

$$S_i^{n,n-k} = \frac{\sum_{j=1}^{n_{bas}} c_{conv,i,j}^n c_{conv,i,j}^{n-k}}{\sum_{j=1}^{n_{bas}} (c_{conv,i,j}^n)^2} \quad (1 \leq i \leq n_{mo})$$

is computed for a sub-set of  $n_{mo}$  orbitals, usually the occupied ones. If  $S_i^{n,n-k} < S_{min}$ , then the orbital  $i$  is not extrapolated and it is given by

$$c_{guess,i,j}^{n+1} = c_{conv,i,j}^n.$$

In the current version,  $n_{bas}$  and  $n_{mo}$  are automatically set by NEWTON-X reading the COLUMBUS input files.  $S_{min}$  is set to 0.7.

For a given extrapolation order  $k-1$ ,  $k$  sets of molecular orbital coefficients are necessary. In the beginning of the dynamics, while the step number is still smaller than  $k$ , the LEMO algorithm is executed at lower extrapolation order, using the molecular orbital sets available.

### 12.1.11 Boundaries

The system can be included in rigid boundaries (at the moment only spherical). If the file 'boundaries.inp' is present in the treatment of boundaries is activated.

In the spherical boundary, if any molecule attempts to move outwards, it undergoes an elastic collision and their radial velocity is inverted.

Name: boundaries.inp

Parameter	Description
INSPHERE	Include system in a rigid sphere. At the second line give the radius and Cartesian coordinates for the center of the sphere in Bohr. rr.r xx.x yy.y zz.z

### 12.1.12 Stop conditions

It is possible to define quite arbitrary set of conditions that when satisfied will cause the end of the simulations. This is done through a series of conditions connected by logic operators written to stopsign.inp file.

Each condition X (integer) with operator Y (and, or) is defined by four keywords.

Name: *stopsign.inp*

Parameter	Description
%comm_Y[X]	Single line Perl command to collect information to be checked.
%text_Y[X]	Description of the command.
%oprt_Y[X]	Any of: ==, !=, eq, ne, <, <=, >, >= (Perl operators)
%thrs_Y[X]	Threshold or value to be tested.

No defaults are available. An arbitrary number of conditions can be set.

Using the information from *stopsign.inp*, NEWTON-X automatically writes a Perl program called *ss-script.pl*. *ss-script.pl* is run every timestep, checks the conditions and writes the values for each one to *ss-Y-X* file. Then, it reads these values and if the conditions to stop are satisfied (considering all AND and OR operations), it writes the instruction “%STOP” to *ss-result* file. *moldyn.pl* checks the contents of *ss-result* every time step. If it finds %STOP, the trajectory simulation is ended.

Example: Stop trajectory if the energy gap between the ground state and the first excited state is smaller than 0.2 eV.

In this case, only one condition should be checked. The *stopsign.inp* may look like:

```
%comm_and[1]  %= "open(IN,"epot"); $e0=<IN>; $e1=<IN>; $de=($e1-$e0)*27.21138386;
open(OUT,">ss-and-1"); print OUT $de;" # %comm_and[1] is a single line!
%text_and[1]  %= "E1-E0 (eV) "
%oprt_and[1]  %= <=
%thrs_and[1]  %= 0.2
```

%comm\_and[1] is a single line instruction in Perl to read the values of ground state (\$e0) and excited state (\$e1) energies from *epot* file, to compute the difference and convert it to eV (\$de) and to print this value to *ss-and-1* file.

The condition to be satisfied is:

```
%text_and[1]  %oprt_and[1] %thrs_and[1]
"E1-E0 (eV) "      <=          0.2
```

## 12.2 How to execute NEWTON-X

Having all input files at the same directory, type:

```
$NX/moldyn.pl > moldyn.log &
```

where \$NX is the variable containing the path to the NEWTON-X files.

## 12.3 Output files

The output files are written to RESULTS directory.

- 1) The main output with a survey of the dynamics is written to *dyn.out*.
- 2) *nx.log* contains information about gradients, non-adiabatic coupling vectors and state configurations along the trajectory.
- 3) File *dyn.mld* contains a sequence of cartesian coordinates in conventional XYZ format with the dynamics history. It can be read by most of graphic programs such as MOLDEN, MOLEKEL or VMD.
- 4) File *intec* contains the internal coordinates for each time step.

- 5) File en.dat contains the potential energy (au) of each state for each time step (fs) in the following order: Time, Epot(S0), Epot(S1), ..., Epot(nstat), Epot(nstatdyn), suitable to produce Energy x time graphs.
- 6) At each time step, the random number, the cycle and the transition probabilities (in this order) are written to tprob. Hopping events are marked with 20 in the place of the actual random number. The file tprobbyfs gives similar information but in fs<sup>-1</sup> units.
- 7) At each time step, the populations, the wave function normalization and the product  $\mathbf{v} \cdot \mathbf{h}$  are written to sh.out.
- 8) File properties contains oscillator strengths and transition dipole moments for each time step (when available).
- 9) File report\_CI contains informs the trajectory point with energy gaps smaller than E\_CI. Useful to locate conical intersections.
- 10) At every hopping the geometry and velocity are written out as hopp\_geom.<surf1>.<surf2>.<time> and hopp\_veloc.<surf1>.<surf2>.<time>

In DEBUG directory:

- 11) File log.conv (DEBUG directory) gives information about the convergence of the ab initio calculations during the dynamics.
- 12) Dynamics using COLUMBUS program keeps the molecular orbital coefficients for each time step t in the respective subdirectory DEBUG/COL.t.
- 13) Error messages are written to runnx.error file.
- 14) With lvprt  $\geq 2$ , lots of debug data are written to this directory.

In INFO\_RESTART directory:

- 15) At each time step the INFO\_RESTART directory is updated with all necessary information to restart the trajectory if it is necessary. Specific information about the restart status can be found in INFO\_RESTART/restart.inf.

## 12.4 Restarting the job

To restart trajectories:

- 1) Replace the original control.dyn by the modified control.dyn written to the INFO\_RESTART directory. (It may be necessary to set a new value for TMAX in this file.)
- 2) Run the job again in the usual way.

When control.dyn contains NXRESTART = 1, NEWTON-X uses the files contained in INFO\_RESTART directory as new inputs input files. INFO\_RESTART/control.dyn, however, is not used and should be copied to the input directory as indicated in point 1) above. If any keyword should be changed in the restarted job, this must be done in the INFO\_RESTART files (except in the case of control.dyn).

With NXRESTART = 1, the contents of DEBUG, RESULTS and INFO\_RESTART directories are not deleted and NEWTON-X produces a continuous output with the previous and the restarted job

NEWTON-X: Newtonian dynamics close to the crossing seam

together. In the case of jobs running in a batch system, be sure of copying these directories to the node machine together with the other input files.

The content of INFO\_RESTART directory can also be used to restart an independent job, without reference to the previous one. In this case, use NXRESTART = 0 or simply delete this keyword in control.dyn.

It is advisable to back up the trajectory before restating it.

## 12.5 Customized analysis

If an executable file NX\_analysis is present, it will be executed after every time step. This option requires some familiarity with the NEWTON-X file structure but it gives the possibility of performing specific customized tasks without having to modify the source code. An example NX\_analysis file could look like this:

```
#!/bin/bash
echo "Performing custom analysis"
/mypath/analysis1.x >> ../RESULTS/ana1.log
/mypath/analysis2.x JOBEX_1.columbus/WORK/ciudgls* >> ../RESULTS/ana2.log
```

This set of instructions in the example will run the programs analysis1.x and analysis2.x provided by the user and then write the results to ana1.log and ana2.log, respectively. Note that NX\_analysis is executed inside TEMP (see Figure 1). Therefore, the RESULTS directory is reached by ../RESULTS.

## 13 Statistical analysis

ANALYSIS is a set of programs developed to analyse the results of the molecular dynamics simulation performed with NEWTON-X.

The program evaluates the mean value and the standard deviation over several trajectories for several properties. The result is given in function of time.

The input for statistical analysis can be done using `nxinp`, option

```
6. SET STATISTICAL ANALYSIS
```

Then a sequence of submenus gives a series of options which are discussed below. The input and the analysis execution should be done in the directory containing the TRAJi results.

### 13.1 What is needed to run

1. The ANALYSIS program reads the results written into the TRAJi/RESULTS directories, where *i* is the number of each trajectory. This structure of directories is automatically generated by the program `makedir.pl` (see section 6.5).

2. File `prop.inp` should contain the main parameters for the analysis:

Parameter	Default	Description
ITRJ	= [1]	initial trajectory to be analysed.
JTRJ	= [10]	final trajectory to be analysed.
TMIN	= [0]	initial time for the analysis (fs).
TMAX	= [100]	final time for the analysis (fs).
DT	= [0.5]	time step (fs) with which outputs are written. If, for example, dynamics run with DT = 0.5 fs and KT = 3 in <code>control.dyn</code> , the outputs were written every DT*KT = 1.5 fs. Therefore, the time step for analysis should be DT = 1.5. If TRAJ(ITRJ)/ <code>control.dyn</code> is found, DT and KT are read from this file and default is DT*KT, otherwise default is 0.5 fs.
PROPTYPE	= [1]	Kind of properties to be analysed. 1. Energy. 2. Wave function. 3. Internal coordinates. 4. Internal forces (only for PROG = 1 in NEWTON-X (COLUMBUS dynamics)). 5. Velocity autocorrelations function.

*The next keywords are needed only if PROPTYPE = 1-4:*

NSTAT	= [2]	Number of states to be analysed.
-------	-------	----------------------------------

*The next keywords are needed only if PROPTYPE = 2:*

COMPLETE_DATA	= [0]	Complete data for broken trajectories. 0. Do not complete data. Other positive value. If the last time in the trajectory is larger than COMPLETE_DATA, repeat last set of data until TMAX. Neglect trajectories whose last time is smaller than COMPLETE_DATA.
---------------	-------	--

NEWTON-X: Newtonian dynamics close to the crossing seam

The next keywords are needed only if *PROPTYPE* = 3 or 4:

NIC	= [1]	Number of internal coordinates to be analysed.(Maximum = 100)
ICLIST	= [1]	Array with the number of internal coordinates to be analysed. Example: if the stretch corresponds coordinate 1 and the torsion to 12, and one wants to analyse both, ICILIST =1,12 and nic = 2.
BMAT	= [0]	0. Get the internal coordinates from the output files. (Only for PROG = 1 in NEWTON-X (COLUMBUS dynamics)) 1. Run cart2int.x program from COLUMBUS to get the internal coordinates. In this case, an intcfl file with definition of the internal coordinates is required. A standard cart2intin input file will be automatically generated, unless the user provides one.
NAT	= [2]	Number of atoms (relevant only if BMAT = 1).

Example of prop.inp file:

```
&collect
  itrj      = 1,
  jtrj      = 20,
  tmin      = 0.0,
  tmax      = 50.0,
  dt        = 0.5,
  proptype  = 1,
  nstat     = 2,
  nic       = 3,
  iclist    = 8,9,12,
  bmat      = 0,
&end
```

Do not forget "&collect" and "&end".

## 13.2 How to execute ANALYSIS

Got to directory TRAJECTORIES and execute

```
$NX/diagnostic.pl
```

The execution of diagnostic.pl before running the analysis program is optional but recommendable.

Run nxinp to create the input file (see section 13.1).

Having prop.inp and diag.log (optional output of diagnostic.pl) in the TRAJECTORIES directory, execute

```
$NX/analysis.pl >analysis.log &
```

If PROPTYPE = 3 and BMAT = 1, this directory must also contain an intcfl file, with the definitions of the internal coordinates (see section 13.1). See COLUMBUS documentation to get specific information about intcfl.

If you change PROPTYPE and run analysis.pl again, the previous results will not be deleted.

## 13.3 Output files

All output files are written to ANALYSIS directory.

1) prop.proptype (proptype = 1..5) contains the history of the dynamics in the format:

Trajectory	Time	Prop (1)	Prop (2)	...	Prop (nprop)
-----	---	-----	-----	...	-----
itrj	tmin	...	...	...	...

NEWTON-X: Newtonian dynamics close to the crossing seam



itrj	...	...	...	...	...
itrj	tmax	...	...	...	...
itrj+1	tmin	...	...	...	...
...	...	...	...	...	...
jtrj	tmax	...	...	...	...

where Prop(i) is each one of the properties analysed.

## 2) Order of the properties in prop.protype files:

Prop type	Traj	Time	Prop (1)	Prop (2)	Prop (3)	Prop (4)	Prop (5)	Prop (nprop)
1	n	t	Epot	Etot	Ekin	E0	E1	E_nstat
2	n	t	PS	CS	PS.CS	A0	A1	A_nstat
3	n	t	C_1	C_2	...			C_nic
4	n	t	F_1	F_2	...			F_nic
5	n	t	VAF					

where:

n - number of trajectory

t - time (fs)

Epot - current potential energy (atomic unit)

Etot - total energy (atomic unit)

Ekin - kinetic energy (atomic unit)

E0 - ground state potential energy (atomic unit)

E1 - first excited state potential energy (atomic unit)

PS - previous surface

CS - current surface

PS.CS - real number formed by PS and CS. Useful to monitoring the hoppings. Examples:

2.2 = previous surface 2, current surface 2 (no hopping)

1.3 = previous surface 1, current surface 3 (1 -> 3 hopping)

1 = ground state.

A0 - Adiabatic population of the ground state.

A1 - Adiabatic population of the first excited state.

The adiabatic population is computed from the real and imaginary part of the electronic wave function

( $\psi$ ) given in dyn.out. For state  $j$ , it is:

$$A_j = \text{Re}(\psi_j)^2 + \text{Im}(\psi_j)^2.$$

**C\_i** - internal coordinate i (Angstrom, rad).

F\_i - internal force for coordinate i (units?).

VAF - Velocity autocorrelation function for trajectory  $n$  defined as

$$VAF(t, t_0) = \frac{S_n(t, t_0)}{S_n(0, t_0)},$$

$$S_n(t, t_0) = \frac{1}{N_{at}} \sum_{i=1}^{N_{at}} \mathbf{v}_{i,n}(t_0) \cdot \mathbf{v}_{i,n}(t_0 + t),$$

where  $\mathbf{v}_{i,n}$  is the velocity vector for atom  $i$ .

## 3) Order of the properties in mean\_value.protype files:

Prop type	Time	N	Prop (1)	Prop (2)	Prop (3)	Prop (4)	Prop (5)	Prop (6)	Prop (7)	Prop (nprop)
1	t	k	<Epot>	D(Epot)	<Etot>	D(Etot)	<Ekin>	D(Ekin)	<E0>	D (Enstat)
2	t	k	f_0	<A0>	D(A0)	f_1	<A1>	D(A1)	<A2>	D (Anstat)
3	t	k	<C_1>	D(C_1)	<C_2>	D(C_2)	...			D (C_nic)
4	t	k	<F_1>	D(F_1)	F_2	D(F_2)	...			D (F_nic)
5	t	k	c*t	<VAF>	D(VAF)					

where:

$k$  - number of used in the computation of the mean value and of the standard deviation.

$f_i$  - relative amount of trajectories in the state  $i$ .

$\langle P \rangle$  - mean value of  $P$  over  $k$  trajectories.

$D(P)$  - Standard deviation of  $P$ . The standard deviation is computed as the square root of the bias-corrected variance:

$$S[P(t)] = \left[ \frac{1}{k-1} \sum_{i=1}^k (P(t,i) - \langle P(t) \rangle)^2 \right]^{1/2}.$$

$c$  - speed of light in cm/fs.

4) When `proptype = 3` and `bmat = 1`, file `intec_new` is added to each TRAJi/RESULTS directory. With `proptype = 3` you may experience discontinuities of  $\pi$  and/or  $2\pi$  of the torsion as a function of time. You can use `smoothangle.pl` (in your NEWTON-X-directory) to correct this, but read the documentation first. There are some caveats!

5) The `prop.proptype` files contain the histories of all trajectories sequentially. You can use `splithist.pl` (in your NEWTON-X-directory) to split this in single files for all trajectories. You get a set of files named `prop.proptype.trajnr`.

6) With the tool `collectjumps.pl` (in you NEWTON-X-directory) you can collect some information about hopping events.

# 14 Normal Mode and Essential Dynamics Analysis

---

The idea of the Normal Mode Analysis (NMA) is to describe the molecular motion in terms of its normal mode displacements. Another approach for analysing dynamics motions is called Essential Dynamics <sup>40</sup>. It is a principal component analysis of the geometric displacements intended to find important motions in the dynamics. This is performed by diagonalizing the covariance matrix. The eigenvectors give the modes of interest; the corresponding eigenvalues represent the variance of these modes.

The NMA package in NEWTON-X <sup>41</sup> was developed to perform normal mode analysis, essential dynamics and other related tasks using the output structure of the program. It is driven by the program `nma.pl` which allows for several options. The main option `nma` performs the normal mode analysis. Before that, superposition can be carried out with option `align`. As reference it is possible to take an equilibrium structure or the average structure over all trajectories, created with option `av_struc`. Essential dynamics can be carried out with option `ess_dyn` (after optional superposition). Output from `ess_dyn` has the form of a normal mode collection and can be used as alternative input for `nma`.

## 14.1 Normal mode analysis

First a reference structure is subtracted from the coordinate vector. This difference vector is multiplied with the inverse of the normal mode matrix for the coordinate transformation. Several averages are printed out and optionally plots are created.

Go to TRAJECTORIES directory and run:

```
$NX/nma.pl nma
```

### 14.1.1 Input parameters

The input is contained in `nma.inp` which has to be in the folder from which the script is executed. The normal mode matrix is read in from a Molden input file.

*File: nma.inp*

Parameter	Description
REF_STRUC_FILE	File with a reference structure (either an equilibrium structure or the average structure created with <code>av_struc.py</code> ).
REF_STRUC_TYPE	File type of this file (xyz, tmol, ...).
VIBRATION_FILE	MOLDEN input file that contains the normal modes.
FIRST_TRAJ	Index of the first trajectory
LAST_TRAJ	Index of the last trajectory
DT	Length of time step (fs)
NUM_STEPS	Maximum number of analysed time steps.

ABS_LIST	List of normal modes for which the absolute value is taken because of symmetry. The numbering is according to the Molden input file. Without this setting all non-totally symmetric modes should average out to 0.
NEG_LIST	List of normal modes where the negative value is taken in <b>total_std.txt</b> and <b>cross_av_std.txt</b> . This is only for convenience when viewing the results.
ANA_INTS	For which time intervals the averaging is carried out.
PLOT	Plots are automatically created. For this the matplotlib/pylab package has to be installed.
DESCR	Name of the subdirectory into which results are written.

---

#### Example of nma.inp file:

```
# input for nma.py
ref_struc_file = '../coord'
ref_struc_type = 'tmol'
vibration_file = '../molden.input'
first_traj = 1
last_traj = 50
dt = .5
num_steps = 201
abs_list = [7,8,9,11,12,15,16,17,19,20,22,26,28,29,30,32,33,36,37]
neg_list = abs_list
ana_ints = [[0,101],[101,201],[0,201]]
plot = True
descr = ''
```

### 14.1.2 Text Output

Output text files are space separated tables and can be read into a plotting program. All values are in Angstrom.

- Trajectory specific information in in each **RESULTS** directory
  - **nma\_<descr>.txt** contains the direct transformation of the coordinates for one trajectory. After plotting, the time evolution of each coordinate can be observed.
  - **nma\_<descr>\_av.txt** and **nma\_<descr>\_std.txt** contain average and standard deviation of the trajectory in the analysed time intervals
- Several averages are put into the **NMA/<descr>** folder:
  - **mean\_against\_time.txt**, **std\_against\_time.txt** - For every timestep the average and standard deviation over all trajectories. These files show the coherent motions.
  - **total\_std.txt** - Total standard deviation over all time steps and trajectories, one number per time interval analysed. It is representative of the total (random and coherent) activity of a normal mode. For a harmonic motion, the standard deviation is directly related to the amplitude.
  - **cross\_av\_std.txt** - Standard deviation of the average trajectory. Through the first averaging random motions are cancelled out and only coherent activity is seen.

### 14.1.3 Graphical Output

Graphics output for averages and for single trajectories can be created if the matplotlib/pylab package is installed.

- Average (created if plot=True in the input file or with 'python nma.py plot')
  - bar graphs in **NMA/<descr>/bar\_graphs**

representing the standard deviation shown in **total\_std.txt** and **cross\_av\_std.txt**.

- in **NMA/<descr>/time\_plots** the time evolution of the normal modes with cross-trajectory-standard-deviation is seen.
- Plots for single trajectories can be drawn by specifying the trajectory index, e.g. '**nma.py plot 1 2 3**' or '**nma.py plot all**', if specific modes are supposed to be plotted into one figure: '**nma.py plot 1 2 3 modes 22 25 27**'. The figures are put into each trajectory's **RESULTS** directory.

## 14.2 Trajectory alignment

Use this option for aligning a set of trajectories using a least squares fit. **Aligning will cancel out the rotational and translational normal mode but leave the other results basically unchanged.** Fitting can be a problem when structures are strongly changing, for more information see Ref. <sup>48</sup>.

Go to TRAJECTORIES directory and run:

```
$NX/nma.pl align <first_traj> <last_traj>
```

<first\_traj> and <last\_traj> are, respectively, the numbers of the initial and final trajectories in the set of interest.

### 14.2.1 Input parameters

*File: align.inp*

Parameter	Description
REF_STRUC_FILE	File with a reference structure (either an equilibrium structure or the average structure created with av_struc.py).
REF_STRUC_TYPE	File type of this file.
OUT_DIR	Output directory.

Example of align.inp

```
# input for align.sh
ref_struc_file = '../..../opt_vib/coord'
ref_struc_type = 'tmol'
out_dir='Aligned_Trajs'
```

## 14.3 Average Structure

Creates the average structure of a set of trajectories.

Go to TRAJECTORIES directory and run:

```
$NX/nma.pl av_struc
```

### 14.3.1 Input parameters

*File: av\_struc.inp*

Parameter	Description
FIRST_TRAJ	Index of the first trajectory
LAST_TRAJ	Index of the last trajectory
NUM_STEPS	Maximum number of time steps in a trajectory

Example of av\_struc.inp:

```
# input for av_struc.py
first_traj = 1
last_traj = 50
```

```
num_steps = 601
```

## 14.4 Essential Dynamics

Finds the most active linear motions through Principal Component Analysis (diagonalization of the covariance matrix)<sup>40</sup>.

Go to TRAJECTORIES directory and run:

```
$NX/nma.pl ess_dyn
```

### 14.4.1 Input parameters

File: *ess\_dyn.inp*

Parameter	Description
REF_STRUC_FILE	File with a reference structure (it is only used for reading in the atom types)
REF_STRUC_TYPE	File type of this file.
FIRST_TRAJ	Index of the first trajectory
LAST_TRAJ	Index of the last trajectory
NUM_STEPS	Number of analysed time steps. This has to be at least as large as the maximum number of time steps in any trajectory.
ANA_INTS	For which time intervals the averaging is carried out.
DESCR	Name of the subdirectory into which results are written.

Example of *ess\_dyn.inp*:

```
# input for ess_dyn.py
ref_struc_file = '../..../opt_vib/coord'
ref_struc_type = 'tmol'
first_traj = 1
last_traj = 10
num_steps = 2001
ana_ints = [[0,501], [501,2001], [0,2001]]
descr = ''
```

The output consists of MOLDEN vibration files for the analysed time intervals. **total\_cov** contains the results with the total covariance over all trajectories and time steps. **cross\_av** shows the essential dynamics of the average trajectory.

For time dependent results, output from this script can be further analysed with *nma* option.

## 14.5 Python subroutine libraries

These libraries are the basis for the provided scripts and could be used for other programming tasks. For more information, look at the documentation included in the python files (either in the source code or with the python 'help(...)' command).

- **chk\_dep.py** : check whether necessary packages are installed
- **traj\_manip.py**: operations for manipulation of trajectories
- **struc\_linalg.py**: package for performing linear algebra operations on structures
- **plotting.py**: plotting routines
- **superposition.py**: superposition of molecules with a quaternion fit
- **vib\_molden.py**: code for parsing a MOLDEN vibration file
- **file\_handler.py**: basic file operations

## 14.6 Required packages to run NMA analysis

In order to run the NMA programs, the packages below should be installed in the system. All of them may be found as part of the UNIX distribution (and should be quickly installed with e.g. '**yum install <package>**') or they can be downloaded from the URLs specified.

- NUMPY Python package ([numpy.scipy.org](http://numpy.scipy.org))
- OPENBABEL-PYTHON package ([openbabel.org/wiki/Python](http://openbabel.org/wiki/Python))
- PYTHON-MATPLOTLIB package (optional for plotting, [matplotlib.sourceforge.net](http://matplotlib.sourceforge.net))

# 15 Tools

---

## 15.1 Plotting energy x time

From RESULTS directory, call

```
$NX/plot
```

The program plot produces a simple gnuplot graph for dynamics by reading en.dat.

The potential energies (atomic units) for all states are plotted in function of time (fs) with lines. For non-adiabatic dynamics, the current potential energy is also plotted, but with points.

To plot the energy difference (eV) between two states along time (fs), call

```
$NX/plotdiff <option1=[arg1] option2=[arg2] ...>
```

The options are

Option	Argument	Description
FILE	filename	Writes the graph to 'filename' on the disk instead of displaying it on the screen. png format is used
TITLE	'Plot title'	Specifies the title of the plot. Default is a combination of the trajectory name and the states, i.e. TRAJn: E(state 1)-E(state 2) where n is the number of the trajectory.
NSTAT1	statenumber	Takes an integer value statenumber which specifies state number 1. If not given as a parameter, the user will be prompted to enter a value interactively.
NSTAT2	statenumber	Takes an integer value statenumber which specifies state number 2. If not given as a parameter, the user will be prompted to enter a value interactively.

To plot the potential energy (atomic units) of several trajectories at once versus time (fs), go to TRAJECTORIES directory and call

```
$NX/plotall
```

## 15.2 Plotting velocities and molecular orbitals

From RESULTS directory, call

```
$NX/arrow
```

The program arrow produces a MOLGEN output file for a certain time step defined by the user. The velocity and non-adiabatic coupling vectors (if available) at this time step are written as a normal mode and can be visualized with MOLGEN compatible programs. For COLUMBUS dynamics is also possible to visualize the molecular orbitals.

The program reads the information from dyn.out, moldyn.log and JOB\_(N)AD and multiplies the velocities components by 100.

The output is the file arrow.mld.



### 15.3 Smoothangle

Analysing the torsional mode may give discontinuities of  $\pi$  and/or  $2\pi$  of the torsion as a function of time. This should be fixed with this script. It is simply looking for change of more than  $\pi/2$  and then correcting this assuming that only jumps of  $\pi$ ,  $3\pi/2$  or  $2\pi$  occur.

Beware! Of course the script cannot know which columns are angles and which are not. It will try to correct *\*all\** values that vary more than  $\pi/2$  from line to line (what can be a problem if you have timesteps bigger than 1.5!). If you expect one value to do so, then don't use this script. This is going to be changed in future versions.

Usage:

1) Prepare the prop.3 file with \$NX/analysis.pl

2) run

`$NX/smoothangle.pl`

The original prop.3 file will be saved as prop.3.old and the new prop.3 file contains the smoothed data.

### 15.4 Collectjumps

Collects information about surface hoppings (max 2 states for the moment).

Usage:

1) prepare file collectjumps.inp with parameters in lines. Possible parameters are:

Parameter	Default	Description
NTRAJ	= [100]	Number of trajectories to analyse
NFIRST	= [1]	First trajectory to analyse
MAXTIME	= [500]	Maximum time, that can occur
ENERGY		Print out energies to jumps
GEOMETRY		Print out selected columns from prop.3
GCOLS		Columns to read from prop.3.j in everyday counting (starting with 1); comma separated list
GLABELS		Labels for geometry columns (comma separated and in the same order as GCOLS)

There is not much checking whether your input is sensible, so you are responsible for this alone. You can write whatever else you want as long as lines beginning with one of the keywords and then a '=' sign is followed by a fit argument lines beginning with anything else will be ignored - write a novel if you feel like it.

2) prepare prop.i.j - files with splithist.pl

prop.2.j files are mandatory (they contain info about hoppings)

prop.1.j and prop.3.j are needed if you want info from there

3) run

`$NX/collectjumps.pl`

Output files are:

jump\_collection.csv -> comma separated values with info

NEWTON-X: Newtonian dynamics close to the crossing seam

jump\_collection.txt -> rather statistical information

## 15.5 Diagnostic

The script diagnostic.pl goes through the TRAJi directories and collects information about error termination and problems with conservation of energy and adiabatic population. A diagnostic of the trajectories is written to diag.log, including up to which time step the information in each trajectory is reliable.

The file diag.log can read by the analysis program (see Chapter 13) and by the initial condition generation program in some cases (see Section 6.3).

To run diagnostic.pl, go to TRAJECTORIES directory and execute:

```
$NX/diagnostic.pl
```

In the beginning of the execution, diagnostic.pl will ask some few questions about path, number of trajectories and time. These same information may be directly provided by means of diag.inp file. The parameters in diag.inp are define below.

*Name: diag.inp*

Parameter	Default	Description
AD	= [current path]	Path to the trajectories.
ITRAJ	= [1]	Initial trajectory to be checked.
FTRAJ	= [no default]	Final trajectory to be checked.
TMIN	= [0.0]	Initial time (fs). It must be common to all trajectories.
ETOT_DEV	= [0.5]	Allowed variation in the total energy (eV).
POP_DEV	= [0.1]	Allowed variation in the norm of the adiabatic population.

Example of diag.inp file:

```
ad = /home/my_system/TRAJECTORIES
itraj = 1
ftraj = 10
tmin = 100
```

## 15.6 Conversion tools

Convert geometry files:

Program	From	To	Usage	Input file	Output file
nx2tm	NX/COLUMBUS	TURBOMOLE	<code>\$NX/nx2tm</code>	geom	coord
nx2xyz	NX/COLUMBUS	XYZ	<code>\$NX/nx2xyz</code>	geom	geom.xyz
nx2dftb	NX/COLUMBUS	DFTB	<code>\$NX/nx2dftb</code>	geom	in.gen
tm2nx	TURBOMOLE	NX/COLUMBUS	<code>\$NX/tm2nx</code>	coord	geom
xyz2nx	XYZ	NX/COLUMBUS	<code>\$NX/xyz2nx &lt; &lt;file&gt;</code>	<file>	geom.
dftb2nx	DFTB	NX/COLUMBUS	<code>\$NX/dftb2nx</code>	in.gen	geom.
finaloutput2dynmld.pl	NX	XYZ	<code>\$NX/ finaloutput2dynmld.pl</code>	final_output	dyn.mld
xyz2zmat	XYZ	ZMAT	<code>\$NX/xyz2zmat</code>	geom.xyz	geom.zmat
nx2gamess.f90	NX/COLUMBUS	GAMESS \$DATA	<code>\$NX/nx2gamess</code>	geom	gamessgeom

## 15.7 Split and merge initial conditions

To split the spectrum and initial condition generation jobs in several jobs to run in several computers, just prepare the input explained in Chapter 10. Use `ISEED = -1` in `initqp_input` file otherwise all jobs will generate the same set of initial conditions. Then run

```
$NX/split_initcond.pl
```

This program will ask a few simple questions (like in how many jobs the parent job should be split), then it will create a directory called `INITIAL_CONDITIONS`, and inside it will create a sequence of subdirectories called `I1`, `I2`, ..., each one containing a complete set of input files.

Copy each subdirectory to a different computer and run the jobs normally.

Script `submit_ic.pl` can be adapted to make a batch submission to your system.

Per default a calculation for the equilibrium geometry is performed in every `I*` directory to compute the reference energy, which is written to a file `epot0`. If you provide this file before the calculation, the initial calculation is skipped. To do this, you can run the calculation in `I1` first and copy `I1/TEMP/epot0` to the other directories.

To merge the jobs, after the calculations, copy the directories `I1`, `I2`, ... back to `INITIAL_CONDITIONS` directory and from inside this directory run

```
$NX/merge_initcond.pl
```

This program will ask the number of jobs to be merged and it will create a new directory called `I_merged` with merged results.

## 16 Technical details

---

### 16.1 Templates and interfaces with new programs

NEWTON-X is written in such way that is easy to interface it with any quantum chemistry package that can provide gradients and other properties necessary for the dynamics. Directory source/template contains some templates to guide the interfacing.

First, it is necessary to have a conversion toll to transform NEWTON-X geometries to the third-party program format. nx2prog.f90 template will help with this.

Second, it is necessary to tell NEWTON-X how to call the new program and how to read and write the properties generated by it. For this purpose, runprog.pl template helps to program this part.

Third, the usage of the new program in the initial condition and absorption spectrum generation, demands a set of instructions of how to call the program and extract energies and oscillator strengths. The template runprog-initcond.pl helps to implement this task.

Besides these specific programs, some intervention in the common code is also required. Normally, only the files moldyn.pl (main NEWTON-X driver), CPAN/colib\_perl.pm (NEWTON-X libraries) and nxinp (input tool) need some changes, which may be identified by comparing the references to previously interfaced programs.

Currently the following interfaces are defined or reserved for future implementation:

Key	Range	Program	Subkey	Method
0:	$0.00 \leq \text{PROG} < 0.95$	Analytic surface		
1:	$0.95 \leq \text{PROG} < 1.95$	COLUMBUS		
2:	$1.95 \leq \text{PROG} < 2.95$	TURBOMOLE		
			2.0:	RI-CC2 /ADC(2)
			2.1:	TD-DFT
3:	$2.95 \leq \text{PROG} < 3.95$	ACES2/CFOUR		
4:	$3.95 \leq \text{PROG} < 4.95$	MOPAC		
5:	$4.95 \leq \text{PROG} < 5.95$	DFTB		
6:	$5.95 \leq \text{PROG} < 6.45$	GAUSSIAN 03		
	$6.45 \leq \text{PROG} < 6.95$	GAUSSIAN 09		
7:	$6.95 \leq \text{PROG} < 7.95$	TINKER		
8:	$7.95 \leq \text{PROG} < 8.95$	DFTB+		
9:	$8.95 \leq \text{PROG} < 9.95$	DFT/MRCI		
10:	$9.95 \leq \text{PROG} < 10.95$	GAMESS	10.0	MCSCF
			10.1	Other methods
20:	$19.95 \leq \text{PROG} < 20.95$	Hybrid jobs		

If you want to contribute to the official version of NEWTON-X, by building a new interface, please, contact us before to reserve a key.

## 16.2 Conversion factors

Physical and mathematical constants, parameters and conversion factors are defined in source/modulus/units\_mod.f90 and source/CPAN/colib\_perl.pm and used throughout NEWTON-X. Some of them are:

Quantity	Atomic units	Other used units
Energy	1 (hartree)	27.21138386 eV
Mass	1 (electron mass)	1/1822.888515 amu
Time	1	24.188843265×10 <sup>-3</sup> fs
Length	1 (bohr)	1/0.52917720859 Å

## 16.3 Format of internal files

NEWTON-X is build to be mostly application independent. This means that as soon the third-party program calculates some property, this property is written in a standard format that NEWTON-X can read without knowing which program really produced it. Note that, although NEWTON-X reads free format, it is essential to keep the predefined ordering. The following standards are currently defined:

### *Geometry*

During the calculations, the geometry will be updated and written with full double precision Format: free, au.

Name: geom

```

Symbol_1 Atomic_number_1 x_1 y_1 z_1 mass_1
Symbol_2 Atomic_number_2 x_2 y_2 z_2 mass_2
.
.
Symbol_Nat Atomic_number_Nat x_Nat y_Nat z_Nat mass_Nat

```

### *Gradient*

Format: free, au

Name: grad

```

gx_1 gy_1 gz_1
gx_2 gy_2 gz_2
:
gx_nat gy_nat gz_nat

```

### *Potential energy*

Format: free, au

Name: epot

```

Epot_1
Epot_2
:
Epot_nstat

```

### *Non-adiabatic coupling vectors*

Name: nad\_vectors

Format: free, au

```

V(1,2)1,x V(1,2)1,y V(1,2)1,z
V(1,2)2,x V(1,2)2,y V(1,2)2,z
:
V(1,2)Nat,x ...
:
V(1,3)1,x ...
:
V(NS-1,NS)Nat,x ... V(NS-1,NS)Nat,z

```

where NS = NSTAT.

NEWTON-X: Newtonian dynamics close to the crossing seam

The order follow the lines of the triangular matrix:

```

      1      2      3      4      ..      NS-1      NS
1
2      1,2
3      1,3      2,3
4      1,4      2,4      3,4
:
NS-1 ..
NS      1,NS      2,NS      3,NS      4,NS      ..      NS-1,NS

```

Final order:

```

1,2
1,3
2,3
1,4
2,4
:
NS-1,NS

```

### Wave function

Name: wfrun (during the calculations) and wf.inp (input)

Format: free

```

      A1_real          A1_imag
      :
      A_nstat_real     A_nstat_imag

```

### Oscillator strength

Oscillator strengths and transition dipole moments are not written in a standard format to be read *a posteriori* by NEWTON-X. These properties are read directly from the quantum chemistry program outputs by the routine `osc_strength` in library `CPAN/colib_perl.pm`.

### State configuration

NEWTON-X still do not have a standard defined to write the state configuration (e.g., coefficients of the main configurations of the CI vector). For each program, this information is only grepped and written to the standard output at each time step.

## 16.4 Normal modes

In the initial condition generation, NEWTON-X might need to read the normal modes generated by a third-party program. This is done according to the following scheme:

IProg	NM_FLAG	Program	Input	Unit
1	1	GAMESS	$l_c$	(amu) <sup>-1/2</sup>
2	2	TURBOMOLE	$l_{nc}$	1
3	3	COLUMBUS	$l_{mw}$	1
4	2	GAUSSIAN	$l_{nc}$	1
5	<i>x</i>	MOLDEN	<i>nm_flag</i>	-
6	1	DFTB	$l_c$	(amu) <sup>-1/2</sup>
7	3	ACES2	$l_{mw}$	1

In this Table:

$l_c$  - Cartesian normal modes  
 $l_{nc}$  -  $l_c \mu^{1/2}$  normalized Cartesian normal mode  
 $l_{mw}$  - mass-weighted normal mode  
 $\mu$  - reduced mass  
amu - g/mol, atomic mass unity

Thus, when IPROG is set to 1 (read GAMESS output), NEWTON-X assumes that the normal modes are given as Cartesian normal modes in  $(\text{amu})^{-1/2}$ , and the NM\_FLAG is internally set to 1. If IPROG = 5 (read MOLDEN file), NM\_FLAG must be given as an input as well because any kind of normal modes may be written to this type of file.

## 16.5 Output files of the SH program

In the RESULTS directory:

1) sh.out  
log file

2) tprob  
transition probabilities:  
random number | current step | trans\_prob\_1 | ... | trans\_prob\_nstat  
Values for a specific time step are printed only when at least one of the trans\_prob values is larger than  $10^{-7}$ .

In the TEMP directory:

3) wfrun (overwritten)  
current value of the electronic wave function coefficients

4) popev\_info  
direct access working file

5) veloc and control.d  
updated in the case of surface hopping

6) irk  
hopping state in the last step  
0 - normal, 1 - hopping, 2 - frustrated hopping

7) sh.log  
Log information

## 16.6 CIOVERLAP documentation

(Documentation based on the original manual written by Jiri Pittner, September 01, 2010.)

Executables involved:

cioverlap  
cis\_casida  
cis\_slatergen  
civecompare  
civeconsolidate  
readsifs  
mcp.x, cip.x

The functionality, input options and input files of these programs are explained in the following subsections. The CIOVERLAP set of programs is a stand-alone package developed by Jiri Pittner to compute the overlap of two CI wavefunctions.

The options and input files necessary to run the CIOVERLAP program during the dynamics run are internally created and updated by NEWTON-X. The command line options of the core program (CIOVERLAP) can be controlled by the user through the keyword CIO\_OPTIONS and CISC\_OPTIONS in jiri.inp file (see section 12.1.7). These options are described below, in section 16.6.1.

### 16.6.1 CIOVERLAP program

Core program to compute overlap of two CI wave functions

<i>Command line options:</i>	<i>Description:</i>
-s screening_mask_file	name of file with transmomin format, tells cioverlap which overlap matrix elements are needed and the Slater det. prescreening is based on this info
-a	align phases of "new" w.f. in rows (bras); use file "phases.old" to patch phases of wavefunctions from previous step
-b	align phases of "new" w.f. in columns (kets); use file "phases.old" to patch phases of wavefunctions from previous step
-o	use file "phases.old" to patch phases of wavefunctions from previous step
-t screeningthr	double; threshold for the Slater det. prescreening procedure
-e excitrank	integer; excitation rank for the Slater det. prescreening procedure determinant pairs which are mutually more than excitrank-excited will be omitted. For TDDFT the value -1 suppresses this screening and unnecessary generation of excitlistfile
-i inactive	integer; number of orbitals which are never excited from generally, it is $\geq$ ncore. This parameter influences only efficiency of the computation if omitted or lower than correct value is given; incorrectly high value will cause erroneous result
-A activerange_from activerange_to	[type] integer; default value 0 calculates overlap in basis of Slater determinants (standard input described below) value 1 calculates overlap in basis of FULL CI GUGA wave functions (standard input is different)
<i>Standard input:</i>	<i>Description</i>
nbas	integer; total number of basis functions, includes core and discarded ones
ncore	integer; number of frozen core orbitals into ncore count only such core orbitals which are not included in slaterfile
ndisc	integer; number of discarded virtual orbitals
nelec	integer; total number of electrons
Sraw	square matrix of dimension 2*number of AOs; obtained as AO overlap matrix of "doubled molecule" combining the bra's and ket's geometry
braLCAO	LCAO matrix of bra, in AO,MO order



ketLCAO LCAO matrix of ket, in AO,MO order

<b>External files:</b>	<b>Description:</b>
slaterfile	input, binary file of signed 16-bit integers contains a list of the Slater determinants forming the N-electron basis alpha spinorbitals with +, beta with - sign ncore orbitals are not included in these determinants generated by cipc.x, mcpc.x, cis_slatergen, or civeconsolidate also output if sorting to lexical order has been requested (presently inactive)
slaterpermfile	output, permutation of slater dets. if sorting to lexical order has been requested presently inactive
eivectors1	input, binary file with two 32-bit integers followed by doubles dimensions and content of the matrix of bra CI wave functions
eivectors2	input, binary file with two 32-bit integers followed by doubles dimensions and content of the matrix of ket CI wave functions
phases.old	input, binary file of doubles phases by which "old" eivectors were scaled what is old and new depends on option -a/-b
Phases	output, binary file of doubles phases by which "new" eivectors should be scaled to align phases what is old and new depends on option -a/-b
Excitlistfile	input/output, binary file of integers auxiliary file for the screening by excitation rank (-e) if not existent/empty it will be created otherwise it will be read

*Execution example:*

```
(echo $nbas $ncore $ndisc $nelec; cat SMAT $BASEDIR/tmp.old/WORK/lcao lcao) |
cioverlap -s transmomin -b -t 1e-5 -e 2
```

## 16.6.2 CIS\_CASIDA

Auxiliary program to generate CIS-like wavefunction coefficients from TDDFT response functions.

<b>Command line options:</b>	<b>Description:</b>
-o	orthonormalize the wavefunctions after applying Casida's formula
-c	use Casida's formula with excitation and orbital energies to scale response function -> wave function coefficients. Note that the resulting w.f. will not be orthonormal unless -o is used simultaneously. By default this option is off and response matrix coefficients are used directly, leading to orthonormal CIS-like wave functions.
-i inactive_occ	integer; neglect all excitations from the number of lowest orbitals in the TDDFT response function
-I inactive_virt	integer; neglect all excitations to the number of highest orbitals in the TDDFT response function

<b>Standard input:</b>	<b>Description:</b>
ncore	integer; number of frozen core orbitals which were not active in TDDFT

NEWTON-X: Newtonian dynamics close to the crossing seam

nocc	integer; number of occupied orbitals which were active in TDDFT, including those to be later excluded by -i option
nvirt	integer; number of virtual orbitals which were active in TDDFT, including those to be later excluded by -I option
orbener	integer size + vector of doubles, size ncore+nocc+nvirt KS orbital energies
energies	integer size + vector of doubles ground and excited state energies
tddft_coefs	vector of matrices response function coefficients for all excited states

---

<b>External files:</b>	<b>Description:</b>
casidawf	output, binary file with the wave function suitable as eivectors input for cioverlap

---

*Execution example:*

```
(echo $ncore $nocc $nvirt; cat orbener energies tddft_coefs)|cis_casida
```

### 16.6.3 CIS\_SLATERGEN

Auxiliary program to generate CIS-like Slater determinant basis for TDDFT runs.

---

<b>Command line options:</b>	<b>Description:</b>
-i inactive_occ	integer; neglect all excitations from the number of lowest orbitals in the TDDFT response function must be same as in cis_casida
-I inactive_virt	integer; neglect all excitations to the number of highest orbitals in the TDDFT response function must be same as in cis_casida

---

<b>Standard input:</b>	<b>Description:</b>
<i>All must be same as for cis_casida</i>	
ncore	integer; number of frozen core orbitals which were not active in TDDFT
nocc	integer; number of occupied orbitals which were active in TDDFT, including those to be later excluded by -i option
nvirt	integer; number of virtual orbitals which were active in TDDFT, including those to be later excluded by -I option

---

<b>External files:</b>	<b>Description:</b>
slaterfile	output, binary file of signed 16-bit integers file with Slater determinant basis suitable as input to cioverlap

---

*Execution example:*

```
echo $ncore $nocc $nvirt |cis_casida
```

### 16.6.4 CIVECCOMPARE

Debugging tool to compare CI wave functions from different sources.

This program is normally not used in the overlap calculation; it can be used to compare e.g. CASSCF and FCI wave functions or CI wave functions computed by different programs.

<b><i>Command line options:</i></b>	<b><i>Description:</i></b>
-f	flip spins in w.f. 1
-g	flip spins in w.f. 2
eivectors1	input file name
slaterfile1	input file name
eivectors2	input file name
slaterfile2	input file name
<b><i>Standard input:</i></b>	<b><i>Description:</i></b>
freeze1	integer; number of electrons to be frozen from w.f.1 before comparison
freeze2	integer; number of electrons to be frozen from w.f.2 before comparison
nelec	integer; number of active electrons, must be same for w.f. 1 and 2 after the freezing threshold double; threshold to consider difference in CI coefficients significant
<b><i>External files:</i></b>	<b><i>Description:</i></b>
only files named on the command line (see above)	

***Execution example:***

```
echo 14 0 3 .00001 | \
/home/pittner/cioverlap-1.0/civeccompare -f \
/home/pittner/cipc/COL.0.50/WORK/eivectors1 \
/home/pittner/cipc/COL.0.50/WORK/slaterfile \
/home/pittner/mcpc/COL.0.50/WORK/eivectors1 \
/home/pittner/mcpc/COL.0.50/WORK/slaterfile
```

**16.6.5 CIVECCONSOLIDATE**

Auxiliary program to simplify "raw" CI wave functions obtained from cipc.x or mcpc.x.

Since cipc.x and mcpc.x in general generate (up to a permutation of spinorbitals) identical Slater determinants repeatedly from different GUGA spin adapted functions, this program shortens the CI wave function expansion by sorting and merging them to a unique order. Omitting its execution should only decrease efficiency, but might also slightly affect the result if the prescreening is involved.

<b><i>Command line options:</i></b>	<b><i>Description:</i></b>
eivector_original	input file name
slaterfile_original	input file name
eivector	output file name
slaterfile	output file name
consolidatefile	input/output - data for the consolidation transformation, created if

not existing yet

<b>Standard input:</b>	<b>Description:</b>
nelec	integer; number of electrons
orbnumshift	integer; shift orbital numbers by subtracting this during processing

*Execution example:*

```
echo $nelec 0 |civeconsolidate eivectors2.org slaterfile.org eivectors2 slaterfile
consolidatefile
```

### 16.6.6 READSIFS

Auxiliary program to read Columbus integral files.

Converts Columbus integral files in SIFS format to a form suitable for cioverlap.

<b>Command line options:</b>	<b>Description:</b>
-1	process only one-electron integrals
aoints	aoints file name

<b>External files:</b>	<b>Description:</b>
aoints, aoints2	input, SIFS integral files
aoints1S	output, binary files with integer dimensions followed by double precision contents of individual matrices (symmetric in packed storage)

*Execution example:*

```
dalton.x
readsifs -1 aoints >readsifs1s
bin2smatrix aoints1S >SRAW
```

(In NX cio\_end is employed instead of bin2smatrix to convert the overlap matrix for formatted input to cioverlap.)

### 16.6.7 CIPC.X, MCPC.X

Programs borrowed from COLUMBUS to generate slaterfile and eivectors file for cioverlap. See appropriate sections of Columbus manual or run them and follow the interactive menu. Columbus has to be compiled with the “-assume byterecl” option, as shown in the machine configuration file “linux64.ifc.byterecl”.

## 16.7 Quick description of the programs

### 16.7.1 Initial conditions

- `initcond.pl`  
Perl script. Control the process.
- `initqp.f90`  
Fortran 90. Generate an harmonic oscillator distribution.  
Written by Giovanni Granucci, Pisa.

- `readall.f90`  
Fortran 90. Read result of `initqp` to prepare the input to quantum chemistry calculations.
- `run_X_initcond.pl`  
Perl script. Execute and read X program.
- `writeall.f90`  
Fortran 90. Write output after checking the energies.
- `faux.f90`  
Fortran 90. Read and write interface between user input and the Perl control.
- `weight.f90`  
Fortran 90. Read and write the normal modes.  
Transform to mass-weighted modes if necessary.

## 16.7.2 Dynamics

- `moldyn01.f90`, `moldyn02.f90`, `moldyn03.f90` and `moldy04.f90`  
Fortran 90 codes to integrate the Newton's eq. in one time step, by means of the Velocity Verlet Algorithm [Swope et al. 1982].  
The first code computes  $r(t+dt)$  and  $v(t+dt/2)$ .  
The second code computes  $v(t+dt)$ ,  $E_{kin}$ , and  $E_{tot}$ .  
The third code writes outputs.  
The fourth code calculates the velocity after hopping in the interpolation process.
- `moldyn.pl`  
Perl script to control the several time steps of the dynamics.
- `inp.f90`  
Read `control.dyn` and re-write the parameters to be read by the other programs (Fortran 90).
- `run-<third-party program>.pl`  
Perl script to execute a third-party program.
- `readai_col.pl`  
Perl script to read and write the energies and gradients output from COLUMBUS program.
- `faux.f90`  
Initialize the counter. (Fortran 90)
- `typeofdyn.f90`  
From the potential energy values, it check the type of dynamics.  
(Fortran 90)
- `write_pair.f90`  
Check restrictions and writes the list of pairs of states for which the non-adiabatic coupling should be computed.  
(Fortran 90)
- `sh.f90`  
Transition probability and surface hopping. Written by G. Granucci.

(Fortran 90)

### 16.7.3 Tools

- `makedir.pl`  
Read output of initial conditions generation and create input directories for NEWTON-X. (Perl script)
- `md2tm.f90`  
Transform geometry from NEWTON-X format to TURBOLMOLE format. (Fortran 90)
- `md2col.f90`  
Transform geometry from NEWTON-X format to COLUMBUS format. (Fortran 90)
- `plot`, `plotall`, `plotdiff` (Perl script)  
Read `en.dat` and produce a gnuplot graphs of Energies x Time.
- `arrow` (Perl script)  
Read `dyn.out` and produce a molden input with geometry and velocity of some time step.
- `escalar.f90`  
Get the inner product between two vectors. Used for the phase calculation. (Fortran 90)
- `interpol.f90`  
Interpolate two vectors. Used during the sub-steps of the TDSE integration. (Fortran 90)
- `spectrum.pl`  
Generate absorption spectrum.
- `hist.pl`  
Calculate frequency table for the absorption spectrum histogram.
- `thermostat.f90`  
Program to generate thermal velocities.
- `hybrid_input_fromxyz.pl`  
Tool to create hybrid inputs from XYZ-files (usage of nonstandard-names instead of atomic symbols possible and encouraged). You need for each region in your hybrid setup one XYZ file. Usage is explained in the tutorial.

### 16.7.4 Statistical analysis

- `analysis.pl` (Perl)  
General controller.
- `collect.pl` (Perl)  
Collect results in TRAJi/RESULTS.
- `aninp.f90` (Fortran 90)  
Rewrite input file.
- `rwprop.f90` (Fortran 90)

Rewrite temporary files to final output format.

- `statistics.f90` (Fortran 90)

Compute mean value and standard deviation.

- `smoothangle.pl` (Perl)

Corrects discontinuities in angle-values in prop.3.

- `splithist.pl` (Perl)

Splits prop.protype files in single history-files for each trajectory.

- `collectjumps.pl` (Perl)

Collects information about surface hopping events.

- `diagnostic.pl` (Perl)

Collect information in each trajectory about errors of execution, conservation of energy and conservation of adiabatic population.

- `run-dftb` (Perl)

Run DFTB program and read the results.

## 17 Links to third-party programs

---

***COLUMBUS:***

[www.univie.ac.at/columbus](http://www.univie.ac.at/columbus)

***TURBOMOLE:***

[www.turbomole.com](http://www.turbomole.com)

***DFTB and DFTB+:***

[www.dftb.org](http://www.dftb.org)

***GAUSSIAN:***

[www.gaussian.com](http://www.gaussian.com)

***TINKER:***

[dasher.wustl.edu/tinker](http://dasher.wustl.edu/tinker)

***GAMESS:***

[www.msg.ameslab.gov/games](http://www.msg.ameslab.gov/games)



## 18 References

---

- <sup>1</sup> M. Barbatti, G. Granucci, M. Persico, M. Ruckebauer, M. Vazdar, M. Eckert-Maksić, and H. Lischka, *J. Photochem. Photobiol., A* **190**, 228 (2007).
- <sup>2</sup> M. Barbatti, G. Granucci, M. Ruckebauer, F. Plasser, J. Pittner, M. Persico, and H. Lischka, *NEWTON-X: a package for Newtonian dynamics close to the crossing seam*, [www.newtonx.org](http://www.newtonx.org) (2013).
- <sup>3</sup> H. Lischka, R. Shepard, R. M. Pitzer, I. Shavitt, M. Dallos, T. Müller, P. G. Szalay, M. Seth, G. S. Kedziora, S. Yabushita, and Z. Y. Zhang, *Phys. Chem. Chem. Phys.* **3**, 664 (2001).
- <sup>4</sup> H. Lischka, R. Shepard, I. Shavitt, R. M. Pitzer, M. Dallos, T. Müller, P. G. Szalay, F. B. Brown, R. Ahlrichs, H. J. Boehm, A. Chang, D. C. Comeau, R. Gdanitz, H. Dachsel, C. Ehrhardt, M. Ernzerhof, P. Höchtl, S. Irlle, G. Kedziora, T. Kovar, V. Parasuk, M. J. M. Pepper, P. Scharf, H. Schiffer, M. Schindler, M. Schüler, M. Seth, E. A. Stahlberg, J.-G. Zhao, S. Yabushita, Z. Zhang, M. Barbatti, S. Matsika, M. Schuurmann, D. R. Yarkony, S. R. Brozell, E. V. Beck, J.-P. Blaudeau, M. Ruckebauer, B. Sellner, F. Plasser, and J. J. Szymczak, *COLUMBUS, an ab initio electronic structure program, release 5.9.2*, [www.univie.ac.at/columbus](http://www.univie.ac.at/columbus) (2008).
- <sup>5</sup> R. Ahlrichs, M. Bär, M. Häser, H. Horn, and C. Kölmel, *Chem. Phys. Lett.* **162**, 165 (1989).
- <sup>6</sup> M. Elstner, *Theor. Chem. Acc.* **116**, 316 (2006).
- <sup>7</sup> M. J. Frisch, G. W. Trucks, H. B. Schlegel, G. E. Scuseria, M. A. Robb, J. R. Cheeseman, J. Montgomery, J. A., T. Vreven, K. N. Kudin, J. C. Burant, J. M. Millam, S. S. Iyengar, J. Tomasi, V. Barone, B. Mennucci, M. Cossi, G. Scalmani, N. Rega, G. A. Petersson, H. Nakatsuji, M. Hada, M. Ehara, K. Toyota, R. Fukuda, J. Hasegawa, M. Ishida, T. Nakajima, Y. Honda, O. Kitao, H. Nakai, M. Klene, X. Li, J. E. Knox, H. P. Hratchian, J. B. Cross, V. Bakken, C. Adamo, J. Jaramillo, R. Gomperts, R. E. Stratmann, O. Yazyev, A. J. Austin, R. Cammi, C. Pomelli, J. W. Ochterski, P. Y. Ayala, K. Morokuma, G. A. Voth, P. Salvador, J. J. Dannenberg, V. G. Zakrzewski, S. Dapprich, A. D. Daniels, M. C. Strain, O. Farkas, D. K. Malick, A. D. Rabuck, K. Raghavachari, J. B. Foresman, J. V. Ortiz, Q. Cui, A. G. Baboul, S. Clifford, J. Cioslowski, B. B. Stefanov, G. Liu, A. Liashenko, P. Piskorz, I. Komaromi, R. L. Martin, D. J. Fox, T. Keith, M. A. Al-Laham, C. Y. Peng, A. Nanayakkara, M. Challacombe, P. M. W. Gill, B. Johnson, W. Chen, M. W. Wong, C. Gonzalez, and J. A. Pople, *Gaussian 03, Revision C.02*. Gaussian, Inc., Wallingford CT (2004).
- <sup>8</sup> M. J. Frisch, G. W. Trucks, H. B. Schlegel, G. E. Scuseria, M. A. Robb, J. R. Cheeseman, G. Scalmani, V. Barone, B. Mennucci, G. A. Petersson, H. Nakatsuji, M. Caricato, X. Li, H. P. Hratchian, A. F. Izmaylov, J. Bloino, G. Zheng, J. L. Sonnenberg, M. Hada, M. Ehara, K. Toyota, R. Fukuda, J. Hasegawa, M. Ishida, T. Nakajima, Y. Honda, O. Kitao, H. Nakai, T. Vreven, J. Montgomery, J. A., J. E. Peralta, F. Ogliaro, M. Bearpark, J. J. Heyd, E. Brothers, K. N. Kudin, V. N. Staroverov, R. Kobayashi, J. Normand, K. Raghavachari, A. Rendell, J. C. Burant, S. S. Iyengar, J. Tomasi, M. Cossi, N. Rega, N. J. Millam, M. Klene, J. E. Knox, J. B. Cross, V. Bakken, C. Adamo, J. Jaramillo, R. Gomperts, R. E. Stratmann, O. Yazyev, A. J. Austin, R. Cammi, C. Pomelli, J. W. Ochterski, R. L. Martin, K. Morokuma, V. G. Zakrzewski, G. A. Voth, P. Salvador, J. J. Dannenberg, S. Dapprich, A. D. Daniels, Ö. Farkas, J. B. Foresman, J. V. Ortiz, J. Cioslowski, and D. J. Fox, *Gaussian 09, Revision A.02*. Gaussian, Inc., Wallingford CT (2009).
- <sup>9</sup> M. S. Gordon and M. W. Schmidt, in *Theory and Applications of Computational Chemistry the first forty years*, edited by C. E. Dykstra, G. Frenking, K. S. Kim, and G. E. Scuseria (Elsevier, Amsterdam, 2005), pp. 1167.
- <sup>10</sup> J. W. Ponder and F. M. Richards, *J. Comput. Chem.* **8**, 1016 (1987).
- <sup>11</sup> J. C. Tully, *Faraday Discuss.* **110**, 407 (1998).

- 12 J. C. Tully and R. K. Preston, *J. Chem. Phys.* **55**, 562 (1971).  
13 J. C. Tully, *J. Chem. Phys.* **93**, 1061 (1990).  
14 M. Barbatti, *WIREs: Comp. Mol. Sci.* **1**, 620 (2011).  
15 E. Fabiano, G. Groenhof, and W. Thiel, *Chem. Phys.* **351**, 111 (2008).  
16 C. Lasser and T. Swart, *J. Chem. Phys.* **129**, 034302 (2008).  
17 A. Ferretti, G. Granucci, A. Lami, M. Persico, and G. Villani, *J. Chem. Phys.* **104**, 5517 (1996).  
18 J. Pittner, H. Lischka, and M. Barbatti, *Chem. Phys.* **356**, 147 (2009).  
19 S. Hammes-Schiffer and J. C. Tully, *J. Chem. Phys.* **101**, 4657 (1994).  
20 G. Granucci, M. Persico, and A. Toniolo, *J. Chem. Phys.* **114**, 10608 (2001).  
21 E. Tapavicza, I. Tavernelli, and U. Rothlisberger, *Phys. Rev. Lett.* **98**, 023001 (2007).  
22 U. Werner, R. Mitrić, T. Suzuki, and V. Bonačić-Koutecký, *Chem. Phys.* **349**, 319 (2008).  
23 M. Dallos, H. Lischka, R. Shepard, D. R. Yarkony, and P. G. Szalay, *J. Chem. Phys.* **120**, 7330 (2004).  
24 H. Lischka, M. Dallos, P. G. Szalay, D. R. Yarkony, and R. Shepard, *J. Chem. Phys.* **120**, 7322 (2004).  
25 C. Y. Zhu, S. Nangia, A. W. Jasper, and D. G. Truhlar, *J. Chem. Phys.* **121**, 7658 (2004).  
26 G. Granucci and M. Persico, *J. Chem. Phys.* **126**, 134114 (2007).  
27 U. Muller and G. Stock, *J. Chem. Phys.* **107**, 6230 (1997).  
28 A. W. Jasper, S. N. Stechmann, and D. G. Truhlar, *J. Chem. Phys.* **116**, 5424 (2002).  
29 L. Verlet, *Phys. Rev.* **159**, 98 (1967).  
30 W. C. Swope, H. C. Andersen, P. H. Berens, and K. R. Wilson, *J. Chem. Phys.* **76**, 637 (1982).  
31 T. Atsumi and H. Nakai, *J. Chem. Phys.* **128** (2008).  
32 H. C. Andersen, *J. Chem. Phys.* **72**, 2384 (1980).  
33 V. Lukes, R. Solc, M. Barbatti, M. Elstner, H. Lischka, and H.-F. Kauffmann, *J. Chem. Phys.* **129**, 164905 (2008).  
34 M. Ruckebauer, M. Barbatti, T. Muller, and H. Lischka, *J. Phys. Chem. A* **114**, 6757 (2010).  
35 S. Grimme and M. Waletzke, *J. Chem. Phys.* **111**, 5645 (1999).  
36 B. Sellner, M. Barbatti, and H. Lischka, *J. Chem. Phys.* **131**, 024312 (2009).  
37 J. P. Bergsma, P. H. Berens, K. R. Wilson, D. R. Fredkin, and E. J. Heller, *J. Phys. Chem.* **88**, 612 (1984).  
38 C. Ciminelli, G. Granucci, and M. Persico, *Chem. Eur. J.* **10**, 2327 (2004).  
39 M. Barbatti, A. J. A. Aquino, and H. Lischka, *Phys. Chem. Chem. Phys.* **12**, 4959 (2010).  
40 A. Amadei, A. B. M. Linszen, and H. J. C. Berendsen, *Proteins-Structure Function and Genetics* **17**, 412 (1993).  
41 F. Plasser, M. Barbatti, A. J. A. Aquino, and H. Lischka, *J. Phys. Chem. A* **113**, 8490 (2009).  
42 J. P. Dahl and M. Springborg, *J. Chem. Phys.* **88**, 4535 (1988).  
43 R. Schinke, *Photodissociation Dynamics: Spectroscopy and Fragmentation of Small Polyatomic Molecules*. (Cambridge University Press, Cambridge, 1995).  
44 R. C. Hilborn, *Am. J. Phys.* **50**, 982 (1982).  
45 J. R. Lakowicz, *Principles of fluorescence spectroscopy*, 3rd ed. (Springer, Singapore, 2006).  
46 E. A. Koopman and C. P. Lowe, *J. Chem. Phys.* **124** (2006).  
47 J. Butcher, *J. Assoc. Comp. Mach.* **12**, 124 (1965).  
48 C. F. F. Karney, *J. Mol. Graph. Model.* **25**, 595 (2007).