

A LETTERS JOURNAL EXPLORING THE FRONTIERS OF PHYSICS

OFFPRINT

Three-dimensional nonparaxial accelerating beams from the transverse Whittaker integral

YIQI ZHANG, MILIVOJ R. BELIĆ, JIA SUN, HUAIBIN ZHENG, HAIXIA CHEN, CHANGBIAO LI, ZHIGUO WANG and YANPENG ZHANG

EPL, **107** (2014) 34001

Please visit the website
www.epljournal.org

Note that the author(s) has the following rights:

- immediately after publication, to use all or part of the article without revision or modification, **including the EPLA-formatted version**, for personal compilations and use only;

- no sooner than 12 months from the date of first publication, to include the accepted manuscript (all or part), **but not the EPLA-formatted version**, on institute repositories or third-party websites provided a link to the online EPL abstract or EPL homepage is included.

For complete copyright details see: https://authors.epletters.net/documents/copyright.pdf.

THE FRONTIERS OF PHYSICS AN INVITATION TO SUBMIT YOUR WORK

A LETTERS JOURNAL EXPLORING

www.epljournal.org

The Editorial Board invites you to submit your letters to EPL

EPL is a leading international journal publishing original, innovative Letters in all areas of physics, ranging from condensed matter topics and interdisciplinary research to astrophysics, geophysics, plasma and fusion sciences, including those with application potential.

The high profile of the journal combined with the excellent scientific quality of the articles ensures that EPL is an essential resource for its worldwide audience. EPL offers authors global visibility and a great opportunity to share their work with others across the whole of the physics community.

Run by active scientists, for scientists

EPL is reviewed by scientists for scientists, to serve and support the international scientific community. The Editorial Board is a team of active research scientists with an expert understanding of the needs of both authors and researchers.



www.epljournal.org

A Letters Journal Exploring the Frontiers of Physics







10,755 citations in 2013

"We greatly appreciate the efficient, professional and rapid processing of our paper by your team."

Cong Lin Shanghai University

Six good reasons to publish with EPL

We want to work with you to gain recognition for your research through worldwide visibility and high citations. As an EPL author, you will benefit from:



Quality – The 50+ Co-editors, who are experts in their field, oversee the entire peer-review process, from selection of the referees to making all final acceptance decisions.



Convenience – Easy to access compilations of recent articles in specific narrow fields available on the website.



Speed of processing – We aim to provide you with a quick and efficient service; the median time from submission to online publication is under 100 days.



High visibility – Strong promotion and visibility through material available at over 300 events annually, distributed via e-mail, and targeted mailshot newsletters.



International reach – Over 2600 institutions have access to EPL, enabling your work to be read by your peers in 90 countries.



Open access – Articles are offered open access for a one-off author payment; green open access on all others with a 12-month embargo.

Details on preparing, submitting and tracking the progress of your manuscript from submission to acceptance are available on the EPL submission website **www.epletters.net**.

If you would like further information about our author service or EPL in general, please visit **www.epljournal.org** or e-mail us at **info@epljournal.org**.

EPL is published in partnership with:









IOP Publishing

www.epljournal.org



Three-dimensional nonparaxial accelerating beams from the transverse Whittaker integral

YIQI ZHANG^{1(a)}, MILIVOJ R. BELIĆ^{2(b)}, JIA SUN¹, HUAIBIN ZHENG¹, HAIXIA CHEN¹, CHANGBIAO LI¹, ZHIGUO WANG¹ and YANPENG ZHANG^{1(c)}

 Key Laboratory for Physical Electronics and Devices of the Ministry of Education & Shaanxi Key Lab of Information Photonic Technique, Xi'an Jiaotong University - Xi'an 710049, China
 ² Science Program, Texas A&M University at Qatar - P.O. Box 23874 Doha, Qatar

received 8 May 2014; accepted in final form 8 July 2014 published online 23 July 2014

PACS 42.25.Fx – Diffraction and scattering PACS 42.25.Bs – Wave propagation, transmission and absorption

Abstract – We investigate three-dimensional nonparaxial linear accelerating beams arising from the transverse Whittaker integral. These beams accelerate along a semicircular trajectory, with almost invariant nondiffracting shapes. The transverse patterns of accelerating beams are determined by their angular spectra, which are constructed from the Mathieu functions, Weber functions, and Fresnel integrals. Our results not only enrich the understanding of multidimensional nonparaxial accelerating beams, but also display their real applicative potential —owing to the usefulness of Mathieu and Weber functions, and Fresnel integrals in describing a wealth of wave phenomena in nature.

Copyright \bigodot EPLA, 2014

Introduction. – Accelerating beams have attracted a lot of attention in the photonics research community in the last decade [1–10]. Being solutions of the linear Schrödinger equation —which is equivalent to the paraxial wave equation in linear dielectric materials— Airy wave packets from quantum mechanics have made a debut in optics [11]. Paraxial accelerating beams propagating in free space or in linear dielectric media are mostly related but not limited to the Airy or Bessel functions [12–16]. Very recently, it was demonstrated that Fresnel diffraction patterns can also exhibit accelerating properties [9]. Interest in these beams stems from the fact that they exhibit self-acceleration, self-healing, and nondiffraction over many Rayleigh lengths [1,2,4,17]. To date, accelerating beams have been discovered in Kerr media [8,18,19], Bose-Einstein condensates [20], on a metal surface [21–24], in photonic crystals [25], and elsewhere. However, the transverse acceleration of paraxial accelerating beams is restricted to small angles.

To extend the theory to nonparaxial range, other kinds of linear accelerating beams have been created and attracted a lot of attention recently. These nonparaxial beams, for example Mathieu and Weber beams, are found by solving the Helmholtz wave equation, and can be quite involved —especially in the multidimensional cases [26-31]. Since the solutions of the three-dimensional (3D) Helmholtz equation can be represented by a reduced form of the Whittaker integral [32], the 3D accelerating beams with different angular spectra can be investigated using this representation. If the angular spectrum can be written in terms of the spherical harmonic functions, parabolic, oblate and prolate spheroidal fields, the 3D accelerating spherical fields with a semicircular trajectory can be obtained [33,34]. In light of nonparaxial accelerating beams ability to bend sharply (~ 90°) or even make U-turns, they are likely to find applications in microparticle manipulations. Thus, nonparaxial accelerating beams deserve further scrutiny.

In this letter, we study the 3D nonparaxial accelerating beams by considering different angular spectra introduced into the transverse Whittaker integral. Since Mathieu and Weber functions are the exact solutions of the Helmholtz equation, it is meaningful to construct the corresponding spectra by using the Mathieu and Weber angular functions. In addition, since the Fresnel integrals are used to describe diffracted waves, it makes sense to employ them to produce nondiffracting nonparaxial Fresnel accelerating beams. Thus, in the following we also use the Fresnel

⁽a) E-mail: zhangyiqi@mail.xjtu.edu.cn

⁽b)E-mail: milivoj.belic@qatar.tamu.edu

⁽c)E-mail: ypzhang@mail.xjtu.edu.cn

integrals to compose an appropriate angular spectrum in the Whittaker integral. Even though Mathieu, Weber, and Fresnel waves are not fully rotationally symmetric, they still propagate along part-circular trajectories and represent interesting linear wave systems in physics. Therefore, here we will investigate novel 3D Mathieu, Weber, and Fresnel nonparaxial accelerating beams together, from the same standpoint. We should stress that although we use Mathieu and Weber functions to construct angular spectra, the Mathieu and Weber beams reported here are different from the Mathieu and Weber beams reported elsewhere. This is obvious from the fact that these beams accelerate along elliptic and parabolic trajectories, whereas our beams accelerate along circular trajectories.

Theoretical model. – The transverse Whittaker integral [32,34] can be written as

$$\psi(\mathbf{r}) = \iint \sin\theta d\theta d\phi A(\theta, \phi) \\ \times \exp\left[ik(x\sin\theta\sin\phi + y\cos\theta + z\sin\theta\cos\phi)\right], \quad (1)$$

where $A(\theta, \phi)$ is the angular spectrum function of the wave ψ , with $\theta \in [0, \pi]$ and $\phi \in [-\pi/2, \pi/2]$. In eq. (1), both Cartesian and spherical coordinates are involved; to clarify the notation used, the relation between coordinates is displayed in fig. 1. Whittaker integral originates from diffraction theory of electromagnetic waves and passes under different names in different fields, for example as the Debye-Wolf diffraction formula in the imaging theory [35]. Arbitrary spectra are allowed, but we will consider the spectrum functions that can be separated in the variables, $A(\theta, \phi) = g(\theta) \exp(im\phi)$, where $g(\theta)$ is a complex functional symmetry in the problem. In this case, eq. (1) can be rewritten as

$$\psi(\mathbf{r}) = \iint \sin\theta d\theta d\phi [g(\theta) \exp(im\phi)] \\ \times \exp\left[ik(x\sin\theta\sin\phi + y\cos\theta + z\sin\theta\cos\phi)\right], \quad (2)$$

which is still quite general. If we restrict ourselves to the value of the colatitude angle $\theta = \pi/2$, eq. (2) will reduce to the two-dimensional case [33]. Utilizing eq. (2), we can display the transverse intensity distributions of shape-invariant beams in an arbitrary plane z = const and the accelerating trajectory in an arbitrary plane y = const.

Although at this point $g(\theta)$ is an arbitrary function, it is natural to assume that it is associated with the solutions of the 2D Helmholtz equation. The procedure is that we first specify solutions of the 2D Helmholtz equation in different coordinates, and then construct an appropriate angular spectrum function $g(\theta)$. The corresponding Helmholtz wave equation is of the form

$$\left(\frac{\partial^2}{\partial z^2} + \frac{\partial^2}{\partial x^2} + k^2\right)f(z,x) = 0,$$
(3)

the solutions of which include many functions —the plane wave functions, Bessel functions, Mathieu functions, and



Fig. 1: (Color online) Relation between Cartesian coordinates (x, y, z) and spherical coordinates (r, θ, ϕ) used.

Weber functions that can be expressed in the Cartesian, circular cylindrical, elliptic cylindrical, and parabolic cylindrical coordinates [36–38], respectively. Here k^2 is the eignevalue of the corresponding Laplace operator's boundary value problem (connected with the wavenumber of the corresponding solutions). In the following sections, we will investigate the three cases for which the angular spectra are given in forms of the Mathieu and Weber functions, and the Fresnel integrals, respectively.

Mathieu beams. – In the elliptic cylindrical coordinates $z = h \cosh \xi \cos \eta$ and $x = h \sinh \xi \sin \eta$, with $\xi \in [0, +\infty)$ and $\eta \in [0, 2\pi)$, the solutions of the Helmholtz equation are the Mathieu functions. By utilizing variable separation, that is, by writing the solution of eq. (3) as $f(\xi, \eta) = R(\xi)\Phi(\eta)$, Helmholtz equation separates into two ordinary differential equations [36,38]

$$\frac{\partial^2 R(\xi)}{\partial \xi^2} - (a - 2q \cosh 2\xi) R(\xi) = 0, \qquad (4a)$$

$$\frac{\partial^2 \Phi(\eta)}{\partial \eta^2} + (a - 2q\cos 2\eta)\Phi(\eta) = 0, \qquad (4b)$$

where a is the separation constant, $q = k^2 h^2/4$ is a parameter related to the ellipticity of the coordinate system, h is the interfocal separation, and $k = 2\pi/\lambda$ is the wave number (λ being the wavelength in the medium). The solutions of eqs. (4a) and (4b) are the radial and angular Mathieu functions. A number of such solutions is known [38].

We utilize the angular Mathieu functions to construct the corresponding spectral function g. Generally, there are four categories of the angular Mathieu solutions of eq. (4b), denoted as S_{ee} , S_{eo} , S_{oe} , and S_{oo} for the eveneven, even-odd, odd-even, and odd-odd cases, respectively. They also carry an order m, which is suppressed here. Thus, the term $g(\theta)$ in eq. (2) may be written as

$$g_1(\theta) = S_{ee}(\theta) + iS_{oo}(\theta), \tag{5}$$

$$g_2(\theta) = S_{oe}(\theta) + iS_{eo}(\theta), \tag{6}$$

and in this manner we introduce the novel Mathieu accelerating beams, by utilizing the transverse Whittaker integral. This is different from the Mathieu nonparaxial



Fig. 2: (a) Angular spectrum function $g_1(\theta)$. (b) Angular spectrum function $g_2(\theta)$. The order of the Mathieu functions is 10, and q = 4.

beams introduced elsewhere [29], where combinations of the radial and angular Mathieu functions have been utilized. Before going to details, we show the angular spectrum functions g_1 and g_2 in figs. 2(a) and (b). The order of the involved Mathieu functions is 10, with q = 4.

Based on the spectral functions g_1 and g_2 , we display the transverse $|\psi(\mathbf{r})|$ distributions of the shapeinvariant Mathieu beams at z = 0 in figs. $3(a_1)-(d_1)$ and figs. $3(a_2)-(d_2)$, respectively. It is seen that the amplitude distributions $|\psi(\mathbf{r})|$ of the beams in the z = 0 plane are symmetric about y = 0, and that they become quite complex as the order is increased. Different from the q_1 case, the amplitude $|\psi(\mathbf{r})|$ at the y = 0 plane is always the smallest for the q_2 case. To observe the accelerating properties as the beams propagate, we choose the 10th order Mathieu functions as an example; numerical simulations are shown in figs. 3(e1) and (e2), which correspond to the cases presented in figs. 3(d1) and (d2), respectively. It is seen that the beams are almost invariant along the z-direction, slowly expanding and accelerating along a quarter-circular trajectory. Note that the beams mirror-extend below the z = 0 plane, so that the beams in the y = const plane accelerate along a semicircular trajectory. This comes from the following argument.

As indicated in the section on the model, we set the polar angle ϕ in the interval $[-\pi/2, \pi/2]$; that means the beam propagates in the positive z-direction. If $\phi \in [\pi/2, 3\pi/2]$, the beam will propagate in the negative z-direction and it will also accelerate along a quarter-circular trajectory, which is symmetric about the plane z = 0 to the positive quarter-circular trajectory. When both the positively and negatively propagating beams are considered together, one observes a beam accelerating along a semicircular trajectory. For convenience, we only consider the case of $\phi \in [-\pi/2, \pi/2]$ throughout the letter. From the 3D plot of the acceleration process, we see that the beams exhibit a nonparaxial accelerating property during propagation —as they should.

Weber beams. – The solutions of the Helmholtz equations in the parabolic cylindrical coordinates are the Weber functions. The relation between the Cartesian and parabolic cylindrical coordinates is $z + ix = (\xi + i\eta)^2/2$,



Fig. 3: (Color online) Cross-section in the z = 0 plane of the amplitude $|\psi(\mathbf{r})|$ of the shape-invariant Mathieu beams with q = 4 and m = 50. The 1st, 4th, 7th, and 10th order Mathieu functions are displayed from left to right. (a1)–(d1) The case of g_1 . (a2)–(d2) The case of g_2 . (e1), (e2) 3D plots of the propagating 10th Mathieu beam, corresponding to g_1 and g_2 .

with $\xi \in [0, +\infty)$ and $\eta \in (-\infty, +\infty)$. When eq. (3) is expressed in the parabolic cylindrical coordinates, and the method of separation of variables applied, the Helmholtz equation transfers into a system of equations [29,37,38]

$$\frac{\partial^2 R(\xi)}{\partial \xi^2} + \left(k^2 \xi^2 - 2ka\right) R(\xi) = 0, \qquad (7a)$$

$$\frac{\partial^2 \Phi(\eta)}{\partial \eta^2} + \left(k^2 \eta^2 + 2ka\right) \Phi(\eta) = 0, \qquad (7b)$$

in which 2ka is the separation constant. If one sets $\sqrt{2k\xi} = u$ and $\sqrt{2k\eta} = v$, eqs. (7a) and (7b) transform into the two standard Weber's differential equations

$$\frac{\partial^2 R(u)}{\partial u^2} + \left(\frac{u^2}{4} - a\right) R(u) = 0, \qquad (8a)$$

$$\frac{\partial^2 \Phi(v)}{\partial v^2} + \left(\frac{v^2}{4} + a\right) \Phi(v) = 0.$$
 (8b)

Therefore, the solutions of eqs. (7a) and (7b) are the Weber functions with different eigenvalues. If we denote the even and odd solutions of eq. (7b) as P_e and P_o , we can define the angular spectrum function $g(\theta)$ as

$$g(\theta) = P_e(\theta, a) + iP_o(\theta, a), \tag{9}$$

in which

$$P_{e,o}(\theta, a) = \sum_{n=0}^{\infty} c_n \frac{\theta^n}{n!},$$
(10a)

and the coefficients c_n satisfy the following recurrence relation:

$$c_{n+2} = ac_n - \frac{n(n-1)c_{n-2}}{4}.$$
 (10b)

To obtain P_e or P_o , one sets the first two c_n coefficients to $c_0 = 1, c_1 = 0$ or to $c_0 = 0, c_1 = 1$, respectively [37].



Fig. 4: (Color online) (a)–(f) Wave amplitude $|\psi(\mathbf{r})|$ distribution at z = 0 of the shape-invariant Weber beams, for m = 120and a = 1, 7, 10, 13, 16 and 19, respectively. (g) 3D plot of the Weber beam with a = 1.

This choice for the angular spectrum functions is again different from the choice made in [29], in which the combinations of solutions of both eqs. (7a) and (7b) are used to construct the full angular spectrum of Weber accelerating beams. These Weber beams accelerate along parabolic trajectories and are different from the ones introduced here.

To display our Weber beams, we repeat the same procedure as for the Mathieu beams; the only difference is that now there are not two g functions but one, expressed in terms of the even and odd Weber eigenfunctions. The amplitude $|\psi(\mathbf{r})|$ distribution of Weber beams at z = 0 for m = 120 and different a, according to eqs. (9) and (10), is shown in figs. 4(a)–(f). It is seen that $|\psi(\mathbf{r})|$ changes slightly with a increasing, especially around the point (x = -m, y = 0). Regardless of what the $|\psi(\mathbf{r})|$ distribution at z = 0 plane is, the beam still accelerates along a quarter-circular trajectory. As an example, fig. 4(g) depicts the accelerating process of the beam with a = 1, which corresponds to fig. 4(a). It is clear that the shape of the beam is nearly invariant as it propagates, and the accelerating trajectory is a quarter of the full circle in the plane y = 0 (the bending angle is approximately 90°). So, the acceleration is nonparaxial.

Fresnel integrals. – In optics, Fresnel integrals are commonly used to describe Fresnel diffraction from a straight edge or a rectangular aperture. The Fresnel cosine and sine integrals are defined as [39]

$$\mathcal{C}(t) = \int_0^t \cos\left(\frac{\pi}{2}\tau^2\right) \mathrm{d}\tau, \qquad (11a)$$

$$\mathcal{S}(t) = \int_0^t \sin\left(\frac{\pi}{2}\tau^2\right) \mathrm{d}\tau, \qquad (11b)$$

or in the complex form:

$$\mathcal{F}(t) = \mathcal{C}(t) + i\mathcal{S}(t). \tag{12}$$



Fig. 5: (Color online) (a) Spectrum function $g(\theta)$ of Fresnel integrals. (b) Transverse intensity distribution of the beam at z = 0. (c) Same as (b) but at z = 40. (d) Circular trajectories of the beam at the plane y = 0. (e) Panels (b), (c) and (d) put together.

Thus, we define the new Fresnel nonparaxial accelerating wave by introducing the spectrum function $g(\theta)$ as

$$g(\theta) = \left(\mathcal{C}(\theta) + \frac{1}{2}\right) + i\left(\mathcal{S}(\theta) + \frac{1}{2}\right).$$
(13)

These Fresnel waves are different from the ones introduced in [9], which propagate paraxially along parabolic trajectories. The spectral function $g(\theta)$ is shown in fig. 5(a), which is the Cornu spiral [39]; the two branches of the spiral approach points P_1 and P_2 , with the coordinates $(1/\sqrt{2}, 1/\sqrt{2})$ and (0, 0).

According to eq. (13) and the Whittaker integral, we can now present the Fresnel nonparaxial accelerating beams. We display intensity distributions of the accelerating beam at z = 0 plane, z = 40 plane, and y = 0 plane in figs. 5(b), (c), and (d), respectively. Concerning the intensity at the z = 0 plane, it is similar to the ones shown in figs. 3(a1) and 4(a), even though the angular spectra are quite different. Comparing figs. 5(b) and (c), it is seen that the beam accelerates —all the lobes move rightward along the positive x-direction. In fig. 5(d) we can follow the acceleration of the main lobe and other higher-order lobes in the y = 0 plane. It is evident that the lobes accelerate along a quarter-circular trajectory, with a bending angle $\sim 90^{\circ}$, which demonstrates that the acceleration is nonparaxial. As before, the beams in fig. 5(d) extend to the region $z \leq 0$, so that the acceleration actually proceeds along semicircular trajectories. In order to see the accelerating process more clearly, we put figs. 5(b)-(d) together in fig. 5(e).

Conclusion. – Based on the transverse Whittaker integral, we have investigated the novel 3D nonparaxial accelerating beams, by constructing the angular spectra using Mathieu functions, Weber functions, and Fresnel integrals. In the transverse plane, the intensity distributions of the accelerating beams are related to the angular spectra. In the longitudinal direction, the beams accelerate along a quarter-circular trajectory, with the bending angle approximately 90°. Even though the angular Mathieu functions, Weber functions, and Fresnel integrals used here are not strictly rotationally symmetric, the study of such 3D nonparaxial accelerating beams is still meaningful and deserves appropriate attention. Since Mathieu and Weber functions are the solutions of the Helmholtz equation, and the Fresnel integrals are used to describe optical diffraction, which are all related to common wave phenomena in the real world, we hope to have revealed the importance of such beams for potential applications.

* * *

This work was supported by the 973 Program (2012CB921804), CPSF (2012M521773), and the Qatar National Research Fund NPRP 6-021-1-005 project. The projects NSFC (61308015, 11104214, 61108017, 11104216, 61205112) are also acknowledged.

REFERENCES

- SIVILOGLOU G. A. and CHRISTODOULIDES D. N., Opt. Lett., 32 (2007) 979.
- [2] SIVILOGLOU G. A., BROKY J., DOGARIU A. and CHRISTODOULIDES D. N., Phys. Rev. Lett., 99 (2007) 213901.
- [3] BANDRES M. A. and GUTIÉRREZ-VEGA J. C., Opt. Express, 15 (2007) 16719.
- [4] ELLENBOGEN T., VOLOCH-BLOCH N., GANANY-PADOWICZ A. and ARIE A., *Nat. Photon.*, **3** (2009) 395.
- [5] CHONG A., RENNINGER W. H., CHRISTODOULIDES D. N. and WISE F. W., Nat. Photon., 4 (2010) 103.
- [6] EFREMIDIS N. K. and CHRISTODOULIDES D. N., Opt. Lett., 35 (2010) 4045.
- [7] EICHELKRAUT T. J., SIVILOGLOU G. A., BESIERIS I. M. and CHRISTODOULIDES D. N., *Opt. Lett.*, **35** (2010) 3655.
- [8] ZHANG Y., BELIĆ M., WU Z., ZHENG H., LU K., LI Y. and ZHANG Y., Opt. Lett., 38 (2013) 4585.
- [9] ZHANG Y., BELIĆ M. R., ZHENG H., WU Z., LI Y., LU K. and ZHANG Y., EPL, **104** (2013) 34007.
- [10] ZHANG Y., BELIĆ M. R., ZHENG H., CHEN H., LI C., LI Y. and ZHANG Y., Opt. Express, 22 (2014) 7160.
- [11] BERRY M. V. and BALAZS N. L., Am. J. Phys., 47 (1979) 264.
- [12] DURNIN J., J. Opt. Soc. Am. A, 4 (1987) 651.
- [13] DURNIN J., MICELI J. J. and EBERLY J. H., Phys. Rev. Lett., 58 (1987) 1499.
- [14] GREENFIELD E., SEGEV M., WALASIK W. and RAZ O., *Phys. Rev. Lett.*, **106** (2011) 213902.

- [15] FROEHLY L., COURVOISIER F., MATHIS A., JACQUOT M., FURFARO L., GIUST R., LACOURT P. A. and DUDLEY J. M., Opt. Express, 19 (2011) 16455.
- [16] HU Y., BONGIOVANNI D., CHEN Z. and MORANDOTTI R., *Phys. Rev. A*, 88 (2013) 043809.
- [17] BROKY J., SIVILOGLOU G. A., DOGARIU A. and CHRISTODOULIDES D. N., Opt. Express, 16 (2008) 12880.
- [18] HU Y., HUANG S., ZHANG P., LOU C., XU J. and CHEN Z., Opt. Lett., 35 (2010) 3952.
- [19] KAMINER I., SEGEV M. and CHRISTODOULIDES D. N., *Phys. Rev. Lett.*, **106** (2011) 213903.
- [20] EFREMIDIS N. K., PALTOGLOU V. and VON KLITZING W., *Phys. Rev. A*, 87 (2013) 043637.
- [21] ZHANG P., WANG S., LIU Y., YIN X., LU C., CHEN Z. and ZHANG X., Opt. Lett., 36 (2011) 3191.
- [22] MINOVICH A., KLEIN A. E., JANUNTS N., PERTSCH T., NESHEV D. N. and KIVSHAR Y. S., *Phys. Rev. Lett.*, **107** (2011) 116802.
- [23] LI L., LI T., WANG S. M., ZHANG C. and ZHU S. N., *Phys. Rev. Lett.*, **107** (2011) 126804.
- [24] LI L., LI T., WANG S., ZHU S. and ZHANG X., Nano Lett., 11 (2011) 4357.
- [25] KAMINER I., NEMIROVSKY J., MAKRIS K. G. and SEGEV M., Opt. Express, 21 (2013) 8886.
- [26] KAMINER I., BEKENSTEIN R., NEMIROVSKY J. and SEGEV M., Phys. Rev. Lett., 108 (2012) 163901.
- [27] COURVOISIER F., MATHIS A., FROEHLY L., GIUST R., FURFARO L., LACOURT P. A., JACQUOT M. and DUDLEY J. M., Opt. Lett., 37 (2012) 1736.
- [28] ZHANG P., HU Y., CANNAN D., SALANDRINO A., LI T., MORANDOTTI R., ZHANG X. and CHEN Z., *Opt. Lett.*, **37** (2012) 2820.
- [29] ZHANG P., HU Y., LI T., CANNAN D., YIN X., MORANDOTTI R., CHEN Z. and ZHANG X., *Phys. Rev. Lett.*, **109** (2012) 193901.
- [30] ALEAHMAD P., MIRI M.-A., MILLS M. S., KAMINER I., SEGEV M. and CHRISTODOULIDES D. N., *Phys. Rev. Lett.*, 109 (2012) 203902.
- [31] BANDRES M. A. and RODRÍGUEZ-LARA B. M., New J. Phys., 15 (2013) 013054.
- [32] WHITTAKER E. T. and WATSON G. N., A Course of Modern Analysis, 4th edition (Cambridge University Press, Cambridge) 1996.
- [33] ALONSO M. A. and BANDRES M. A., Opt. Lett., 37 (2012) 5175.
- [34] BANDRES M. A., ALONSO M. A., KAMINER I. and SEGEV M., Opt. Express, 21 (2013) 13917.
- [35] WOLF E., Proc. R. Soc. A, 253 (1959) 349.
- [36] GUTIÉRREZ-VEGA J. C., ITURBE-CASTILLO M. D. and CHÁVEZ-CERDA S., Opt. Lett., 25 (2000) 1493.
- [37] BANDRES M. A., GUTIÉRREZ-VEGA J. C. and CHÁVEZ-CERDA S., Opt. Lett., 29 (2004) 44.
- [38] ABRAMOWITZ M. and STEGUN I. A., Handbook of Mathematical Functions (Dover Publications Inc., New York) 1970.
- [39] BORN M. and WOLF E., Principles of Optics, 7th edition (Cambridge University Press, Cambridge) 1999.